

Quantum Gravity

Shoaib Akhtar

Lec 1: Introduction and overview by Bianca Ditrich on the different Quantum Gravity approaches.

For now, no experimentally accessible situations

$$\frac{F_{\text{grav}}}{F_{\text{electromag}}} (\text{proton-electron}) \sim 10^{-40}$$

$SU(3) \times SU(2) \times U(1)$ Standard Model.

Quantum Mechanics Framework.

General Relativity.

Open Problems:

(i) Dark Matter	} These may be related.
(ii) Unification	
(iii) Quantum Gravity	

Quantum Gravity = Quantum Mechanics + General Relativity.

G.R. $S = \frac{1}{16\pi G} \int \sqrt{g} \cdot R$

Q.M Classical description of Hydrogen atom unstable.

In Quantum description, we have levels of energy.

QM acts like repulsive force to prevent singularity.

Planck Scale

$$L_{\text{Planck}} = \sqrt{\frac{\hbar G}{c^3}} = 1.6 \times 10^{-35} \text{ m}$$

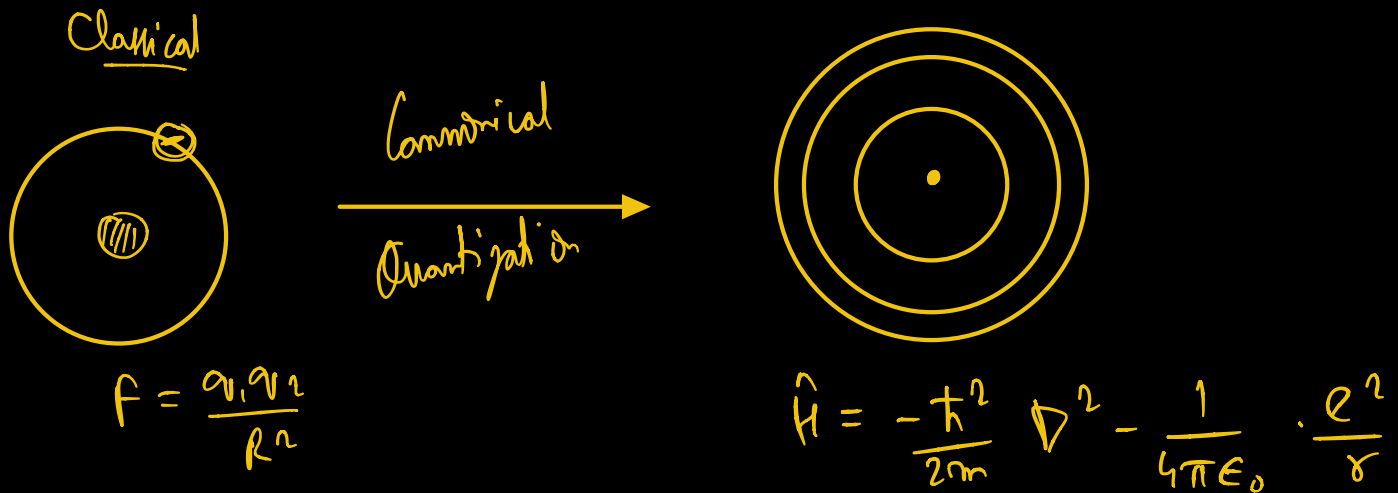
- QM, Compton wavelength: $\lambda_{\text{min}} \sim \frac{\hbar}{mc}$
- GR, Schwarzschild radius: $R \sim \frac{m G}{c^2}$

$$L_{\text{min}} \equiv \lambda_{\text{min}} = R \Rightarrow L_{\text{min}} = L_{\text{Planck}}$$

From Q.M we expect something small.

From G.R " " " dense.

Quantization



Electromagnetic force $\xrightarrow{\text{Feynman}}$ Standard Model.

Gravity $\xrightarrow{?}$ Quantum Gravity.

Perturbative approach : $g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$ with $|h_{\mu\nu}| \ll 1$
does not work!

..... String Theory

- Loop Quantum Gravity. (Dirac's Approach, canonical quantization)

GR (back ground independence, $g_{\mu\nu}$ field)

+
QM (Uncertainty Principle)

"
Theory of Quantum Gravity.

- Einstein-Hilbert formulation (1915)

$$S = \frac{1}{16\pi G} \int d^4x \sqrt{g} \cdot R.$$

- Ashtekar Formulation (1986) (gravity looks as Yang-Mill theory)
in this language.

$$S = \frac{1}{8\pi G \gamma} \int d^4x \left(\tilde{E}^a_i \dot{A}^i_a + N \epsilon_{ijk} \tilde{E}^a_i \tilde{E}^b_j F^k_{ab} + \lambda^i (D_a \tilde{E}^a)^i \right)$$

γ : Barbero - Immirzi parameter

$$\{A^i_a(x), \tilde{E}^b_j(y)\} = 8\pi G \gamma \cdot \delta^a_b \cdot \delta^i_j \cdot \delta^3(x-y)$$

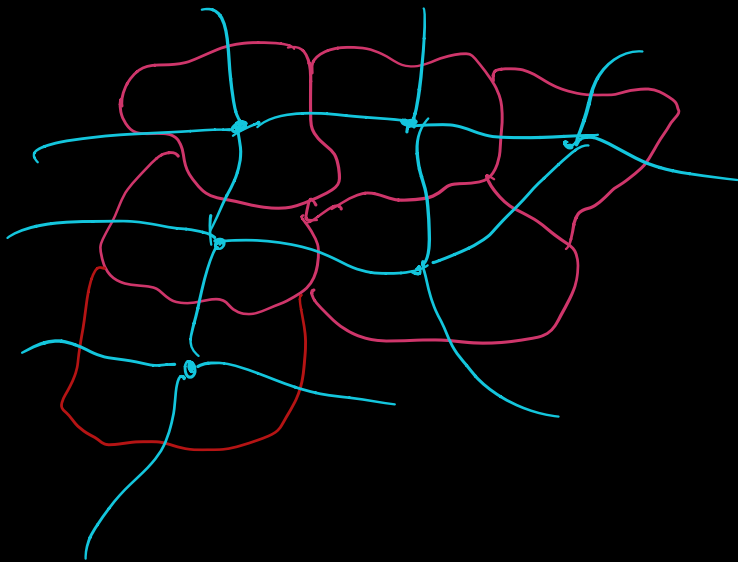
can use loops to measure curvature, so called Loop Quantum Gravity.

★ Spin Networks: basis of the kinematical Hilbert space of Loop Quantum Gravity.

* Carlo Rovelli:

$$A|\Psi\rangle = 8\pi G \cdot L_{\text{plank}}^2 \sum_{p \in \text{sur}} \sqrt{j_p(j_p+1)} |\Psi\rangle$$

$$\text{with } L_{\text{plank}} = \sqrt{\frac{\hbar G}{c^3}} = 1.6 \times 10^{-35} \text{ m.}$$



★ Spin Foam Framework:

Transition amplitude between spin network state of LQG.

General Relativity

Action

Second order
Einstein-Hilbert
Action

1st order
Palatini action

Symmetries
Gauge

Canonical Analysis

Equation of motion
Constraints.

Quantum Mechanics

Quantification Approaches

Dirac

Feynman

perturbative
approach.

Spin
Foams.

Loop Quantum
Gravity

Gravity = Space Time

Quantum Gravity \Rightarrow Quantum Space Time.

Quantum space?
Quantum time?

Back ground
for QFT.

Universe



LHC



Planck Scale

• Top Down Approach



find theories consistent with what we know.

(larger scale to smaller scale)

• Bottom Up Approach



make an intuition about Quantum Gravity; and then show we do get the large scale physics what we know.

ex|| Top Down approach (Asymptotic Safety)

- Takes QFT input

ex|| LQG : standard tool of Quantization.
applied to Space-time Geometry.

- "Quantum Geometry"

★ Bottom Up Approaches

- Causal sets.
- Causal Triangulations.
- Non Commutative Spacetime.
- String Theory.

Quantum Gravity

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Lec 2: Constrained systems and Hamiltonian formalism: Gauge Symmetries.

3D Euclidean gravity with zero Cosmological Constant.

- Canonical formulation of constrained systems.
- Gauge Transformations.

* Why are we interested in "Constrained System" \equiv (Q1)

* Can we relate constraints / gauge symmetries. \equiv (Q2)

(Q1) $g_{\mu\nu} : \quad 4D : 10 \longrightarrow 2$
 $3D : 6 \longrightarrow 0 \quad (\text{Topological Theory})$

$$e^i_\mu$$

$\mu \Rightarrow$ space time indices
 $i \Rightarrow$ internal (Triad) indices

Triad / connection.

Constraint on phase space : $\phi(a_i, p_i) = 0$

* constraints hold at every instant of time in the evolution of the system.

This means; constraint is conserved quantity.

(Q2) Conserved Quantities \longleftrightarrow Symmetries.


Lagrangian formalism :

Symmetry \Rightarrow Conserved quantities (Noether Theorem)

Symmetries $\left\{ \begin{array}{l} \underline{\text{global}} \\ \underline{\text{gauge}} \end{array} \right.$: local and non-trivial transformation of the fields.

local \Rightarrow gauge transformations do not affect any boundary value.

Non-trivial \Rightarrow act non-trivially on extrema of action.



These two evolution are physically equivalent when they are related by gauge transformations.

★ Read on Gauge Transformations

Review of the Hamiltonian Formalism

$$S = S[q_i] = \int L[q_i, \dot{q}_i] dt$$

$$p_i \equiv \frac{\partial L}{\partial \dot{q}_i}, \quad H(q_i, p_i) \equiv p_i \cdot \dot{q}_i - L$$

$$\dot{q} = \{q, H\}.$$

$\{q_i\} \rightarrow$ Configuration space

$\{q_i, p_i\} \rightarrow$ Phase space.

$$S = \int \left[\sum p_i (a_i, \dot{a}_i) - H(p_i, a_i) \right] dt$$

Associate with this symplectic structure in terms of conjugate variables.

$$\{a_i, p_j\} = \delta_{ij}$$

Time Evolution: $\dot{f} = \{f, H\}$: Flow of Hamiltonian.

Any phase space function can generate flow

Hamiltonian formulation for a constrained system.

If is not invertible, then

$$\Rightarrow \Psi_\alpha(a_i, p_i) = 0 \quad \text{Primary Constraints.}$$



Primary Constraint surface.
(all the physical content of theory lie here)

$$H = H_0 + u^\alpha \Psi_\alpha$$

→ Add constraints in terms of Lagrange Multiplier u^α .

$$\{p_i, H\} = \{p_i, H_0\} + \{p_i, u^\alpha \Psi_\alpha\}$$

$$= \{p_i, H_0\} + \{p_i, u^\alpha\} \Psi_\alpha + \{p_i, \Psi_\alpha\} u^\alpha$$

→ zero on constraint surface because of $\Psi_\alpha = 0$.

So: \approx Weak Equality (equal under constraint surface)

$$\boxed{\{p, H\} \approx \{p, H_0\} + \{p, \psi_\alpha\} U^\alpha} \Rightarrow \text{Weak equality:}$$

i.e. it holds true on constraint surface.

U^α : free or fixed?

Example $L = \frac{1}{2}m(\dot{v}_1^2 + \dot{v}_2^2) - \frac{1}{3}a_3(\dot{v}_1^2 + \dot{v}_2^2 - \lambda^2)$

$$\psi_1(a, p) \equiv p_3 = 0$$

$$\dim(\text{phase space}) = 6$$

↓

$$\dim(\text{physical}) = 2$$

$$H = \frac{1}{2m}(\dot{p}_1^2 + \dot{p}_2^2) + \frac{1}{3}a_3(\dot{v}_1^2 + \dot{v}_2^2 - \lambda^2) + U\psi_1$$

$$\{\psi_1, H\} \approx 0 = \psi_2$$

\Rightarrow If not zero; we impose this secondary constraint.

$$\psi_1 \longrightarrow \psi_2$$

$$\psi_2 \longrightarrow \psi_3$$

$$\psi_3 \longrightarrow \psi_4$$

$$\dot{\psi}_4 \longrightarrow \underline{\underline{u=0}}$$

Ex 2.11 $L = \frac{1}{2} m (\dot{v}_1 + \dot{v}_2)^2 - V(a_1 + a_2)$

$\Psi(a, p) \equiv p_1 - p_2 \Rightarrow$ Primary Constraint.

$\dot{\Psi}(a, p) = \{\Psi, H\} = 0$ so done with constraints.

Now to interpret these constraints?

Change of variable: $Q_1 = a_1 + a_2$ $Q_2 = a_1 - a_2$

Lagrangian only acts on Q_1 , not on Q_2 .

Symmetry? $\delta a_1 = u \{a_1, \Psi\} = u$ $u \Rightarrow$ large range multiplier.
 $\delta a_2 = u \{a_2, \Psi\} = -u$

$\delta(a_1 + a_2) = \delta a_1 + \delta a_2 = 0.$

$\delta Q_1 = 0$, but $\delta Q_2 = 2u$

So, Q_1 is physical quantity.

Q_2 is "internal" degree of freedom.

Constraints: primary, secondary,;
 first class, second class.

Distinction between 1st class / 2nd class:

Consider a system with constraints Ψ_α (all constraints, primary, secondary...)

* First class: If there is a subset of constraints φ_α whose bracket with all constraints weakly vanishes,
i.e: $\{\varphi_\alpha, \psi_\alpha\} \approx 0$

Then the φ_α are called first class constraints.

\Rightarrow No restriction on associated Lagrange multiplier.

\Rightarrow Gauge symmetries are generated by first class constraints.

* Second class: Remaining constraints denoted by χ_m ,
then the sub-matrix defined as

$\Delta_{mn} = \{\chi_m, \chi_n\}$ is invertible.

χ_m are called 2nd class constraints.

\Rightarrow Corresponding Lagrange multipliers are (weakly) fixed.

let \mathcal{C} be one of the constraint.

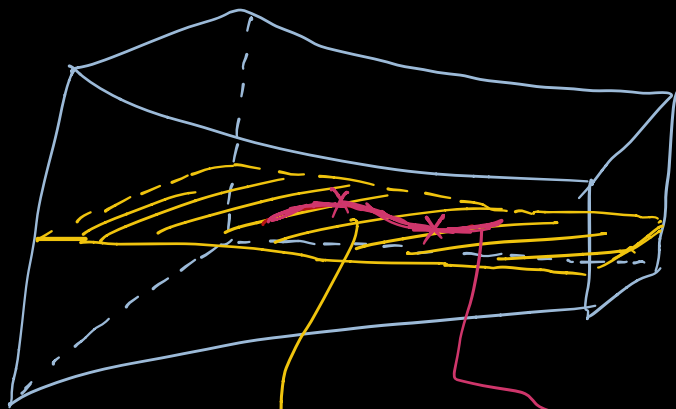
so $\{\mathcal{C}, H\} \approx 0$: The flow generated by H of \mathcal{C}
is zero (at least weakly)

or read as $\{H, \mathcal{C}\} \approx 0$

The flow of H generated by \mathcal{C} is zero.

..... and this is what we call Symmetry.

Constraints & Physical degrees of Freedom.



Kinematical Phase space.



Physical or Reduced Phase space

Constraint surface.

Gauge orbits.

$$\dim(\text{Kinematical phase space}) - \#(\text{1st class}) \times 2 = \dim(\text{physical})$$

because of dividing
by Gauge orbits.

Quantum Gravity

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Lec 3: Examples of parametrized particle: Canonical analysis, Physical phase space & Dirac observables. The Dirac program: case of the parametrized particle.

A totally constrained system - example of the parametrized particle.

Free non-relativistic particle in 1d.

$$S[q] = \int_{t_1}^{t_2} dt \cdot \frac{1}{2} m \dot{q}^2 \quad ; \quad \dot{q} = \frac{dq}{dt} \rightarrow q(t) = q(\tau) + \frac{p(\tau)}{m} (t - \tau)$$

t independent variable

$$S_0[q(s), t(s)] = \int_{s_1}^{s_2} \left(\frac{1}{2} m \cdot \frac{q'^2}{t'} \right) \cdot ds$$

$$q' = \frac{dq}{ds} \quad ' \text{ denotes derivative w.r.t. } s.$$

S_0 is invariant under $s \rightarrow \tilde{s} = f(s)$.

Canonical Analysis

$$p_t = \frac{\delta S}{\delta t'} = -m \cdot \frac{q'^2}{2t'^2}, \quad p_q = \frac{\delta S}{\delta q'} = m \cdot \frac{q'}{t'}$$

We have a primary constraint C

$$C = p_t + \frac{p_q^2}{2m} = 0$$

Hamiltonian of the system: $H = p_t \cdot t' + p_q \cdot q' - L$

$$\Rightarrow \underline{H = \dot{x}' \cdot C}$$

Totally constrained system;
Hamiltonian is proportional to constraints.

$$\underline{H = N(s) \cdot C}$$

flow generated by H ?

$$\alpha' = \{\alpha, H\} = N \cdot \frac{p_\alpha}{m}$$

$$p'_\alpha = 0$$

$$t' = \{t, H\} = N$$

$$p'_t = 0$$

Physical Phase Space: Dirac Observables

Gauge orbits of phase space variables generated by C .

$$\frac{d\alpha}{ds} = \{\alpha, C\} = \frac{p_\alpha}{m} \Rightarrow \alpha(s) = \alpha + \frac{p_\alpha}{m} \cdot s$$

$$\frac{dt}{ds} = \{t, C\} = 1 \rightarrow t(s) = s + t$$

Treating α & t on same footing.

$$\frac{dp_\alpha}{ds} = \{p_\alpha, C\} = 0 \Rightarrow p_\alpha = p_{\alpha_0}$$

$$\frac{dp_t}{ds} = \{p_t, C\} = 0 \Rightarrow p_t = -\frac{p_\alpha^2}{2m}$$

2 independent dirac observables:

$$\hookrightarrow \{F, C\} \approx 0$$

F is dirac observable

To construct a physical observables:

- As many gauge fixing as constraints.

$$t(s) = \tau$$

* Good gauge fixing. } Read.
 \hookrightarrow J. Tambornino.

• $f(s) \longrightarrow F(\tau)$
 phase space function Associated physical observable.

ex) $\alpha_V(s) \longrightarrow \left\{ \begin{aligned} F_{\alpha}^{\tau} &= \alpha_{V0} + \frac{p_0}{m} (\tau - t_0) \\ F_{p_V} &= p_{V0} \end{aligned} \right\}$

* At Kinematical level constraints are not imposed.

\longrightarrow These form complete set of Dirac observables.

So, we form Poisson Bracket: $\{F_{\alpha}^{\tau}, F_{p}^{\tau}\} = 1$

Dirac Program

a) Find a representation of the phase space variable as operators acting in the kinematical Hilbert space \mathcal{H}_{kin} satisfying the standard commutation relation:

$$\{ \cdot, \cdot \} \longrightarrow -i/\hbar [\cdot, \cdot]$$

b) Promote the constraints to (self-adjoint) operators in \mathcal{H}_{kin}

c) Characterize the phase space of solutions of the constraints
 \longrightarrow Inner product $\mathcal{H}_{\text{phys}}$.

d) Find a complete set of gauge invariant observables.

Case of Parametrized Particle

(a) $\mathcal{H}_{\text{kin}} = \mathcal{L}^2[\mathbb{R}^2]$ functions depend on q & t .

$$\Psi(q, t), \quad \langle \Psi, \Psi \rangle = \int dq \cdot dt \cdot \bar{\Psi}(q, t) \Psi(q, t)$$

Kinematical Inner product.

$$q \longrightarrow \hat{q} \quad \hat{q} \Psi(q, t) = q \cdot \Psi(q, t)$$

$$t \longrightarrow \hat{t} \quad \hat{t} \Psi(q, t) = t \cdot \Psi(q, t)$$

$$p_q \longrightarrow \hat{p}_q \quad \hat{p}_q \Psi(q, t) = -i\hbar \cdot \frac{\partial}{\partial q} \Psi(q, t)$$

$$p_t \longrightarrow \hat{p}_t \quad \hat{p}_t \Psi(q, t) = -i\hbar \frac{\partial}{\partial t} \Psi(q, t)$$

$$(b) \quad C = p_t + \frac{p_q^2}{2m} \longrightarrow \hat{C} = -i\hbar \cdot \frac{\partial}{\partial t} - \frac{\hbar^2}{2m} \cdot \frac{\partial^2}{\partial q^2}$$

So, the equation becomes $\hat{C}|\Psi\rangle = 0$.

This is just Schrodinger's equation. ↪

$$(c) \quad \Psi_{\text{phys}}(q, t) = \exp\left(-\frac{i}{\hbar} \hat{h} t\right) \Psi(q)$$

$\Psi(q)$ initial wave function $\Psi(q) = \Psi(q, t=t_0)$

$$\hat{h} = -\frac{\hbar^2}{2m} \cdot \frac{\partial^2}{\partial q^2}$$

Inner Product

$$\langle \Psi_{\text{phys}}^{(1)} | \Psi_{\text{phys}}^{(2)} \rangle_{\text{kin}} = \int \left(\int \bar{\Psi}^{(1)}(q) \Psi^{(2)}(q) dq \right) dt$$

$$\mathcal{H}_{\text{phys}} = L^2(\mathbb{R})$$

Physical Inner product.

$$\langle \Psi_{\text{phys}}^{(1)} | \Psi_{\text{phys}}^{(2)} \rangle = \int \bar{\Psi}^{(1)}(t_{\text{fixed}}, q) \Psi^{(2)}(t_{\text{fixed}}, q) dq$$

Here we don't integrate over t variable.

(d) Two independent observables:

$$\hat{O}_1 = \hat{q} - \frac{\hat{p}_q}{m} (\hat{t} - t_0)$$

$$\hat{O}_2 = \hat{p}_q$$

we can show $[\hat{O}_1, C] = 0$, $[\hat{O}_2, C] = 0$

Quantum Gravity

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Lec 4: Actions of gravity; The Einstein-Hilbert action - 3D gravity - first order formalism of 3D gravity (triads & connection)

Actions for Gravity.

Einstein-Hilbert action.

$$S = \frac{1}{k} \int R \cdot \sqrt{g} \cdot d^3x$$

$$\Rightarrow G_{\mu\nu} = R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} = 0$$

$$\Leftrightarrow R_{\mu\nu} = 0 \quad (\text{for } \dim(\text{spacetime}) \neq 2)$$

Proof of equivalence.

$$R^{\mu}_{\mu} - \frac{1}{2} R \cdot g^{\mu}_{\mu} = 0$$

$$\text{let } D = \dim(\text{spacetime})$$

$$R - \frac{1}{2} \cdot R \cdot D = 0 \Rightarrow \left(1 - \frac{D}{2}\right) R = 0$$

$$\text{for } D \neq 2 \Rightarrow R = 0$$

$$\Rightarrow G_{\mu\nu} = 0 \Rightarrow R_{\mu\nu} = 0 \quad \text{for } D \neq 2.$$

$$R_{\mu\nu} = 0 \Rightarrow R_{\mu\nu} \wedge \delta = 0 \quad \text{in 3d spacetime}$$

$R_{\mu\nu} \wedge \delta$ is antisymmetric in 1st two indices.

$$\Rightarrow \frac{n(n-1)}{2} \Big|_{n=3} = \underline{3 \text{ pairs.}}$$

• antisymmetric in last two indices

$$\Rightarrow \frac{n(n-1)}{2} \Big|_{n=3} = \underline{3 \text{ pairs.}}$$

flat solution.

because
in 3d;
 $R_{\mu\nu}$ &
 $R_{\mu\nu} \wedge \delta$ have
same no. of
degrees of
freedom.

• Symmetric under 1st pair & 2nd pair.

so; $\frac{n(n+1)}{2} \Big|_{n=3} = 6$ independent components in 3D.

$R_{\mu\nu}$: 6 independent components in 3D.

3D gravity = Topological Theory
(no local degrees of freedom)

* We don't expect graviton field in 3D.

because we don't expect local degrees of freedom.

* Topology of the manifold is characterized by the Fundamental group.

In 3D, $\mathcal{M} = \mathbb{R} \times \Sigma$

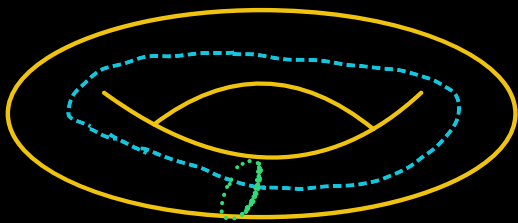
Σ is 2d surface.

The topology of 2d surface is very well classified.

Orientable surfaces are homeomorphic to sphere S^2 or to connected sum of g tori.

g = genus of the surface.

$$g(S^2) = 0 \quad ; \quad \pi_1(S^2) = 0.$$



First order formalism of 3D gravity.

$$\underline{g_{\mu\nu}(x) = e^i_{\mu}(x) \cdot e^j_{\nu}(x) \cdot \delta_{ij}}$$

$\delta_{ij} \Rightarrow$ Minkowski Metric $i, j \in \{1, 2, \dots, D\}$

At each point we raise orthonormal frame

In 3 dimension $e^i_{\mu}(x)$ are called Tetrad.

In 4 " $e^i_{\mu}(x)$ " " Quatriad.

In 3D $g_{\mu\nu} = e^i_{\mu} \cdot e^j_{\nu} \cdot \delta_{ij}$

$g_{\mu\nu}$ has 6 degrees of freedom.

for (e^i_{μ}) we get 9 degrees of freedom.

So we see that from going from metric to triad, we had introduced some irrelevant informations.

So we Additional Gauge Symmetry.

$$e^i_\mu(x) = R^i_j(x) \cdot e^j_\mu(x) \quad ; \quad R \in SO(3)$$

so; we get addition $SO(3)$ gauge symmetry

This transformation keeps the metric unchanged.

Instead of the metric, we can work with triad.

Spin Connection // transport objects with internal indices and also to ensure the proper transformation behavior under internal rotation for the covariant derivative of objects with internal indices.

↳ Spin Connection.

$$\omega_\mu{}^i{}_k \quad D_\mu$$

i, k internal indices.
 μ spacetime index.

ϕ^i is vector for internal indices, and scalar for spacetime index

$$D_\mu \phi^i = \partial_\mu \phi^i + \omega_\mu{}^i{}_k \phi^k$$

$$D_\mu \phi_j = \partial_\mu \phi_j - \omega_\mu{}^k{}_j \phi_k$$

$$D_\mu v^i{}_\nu = \partial_\mu v^i{}_\nu - \Gamma^\rho{}_{\mu\nu} v^i{}_\rho + \omega_\mu{}^j{}_k \cdot v^k{}_\nu$$

Check the derivative of a scalar; $D_\mu \phi = \partial_\mu \phi$

It comes out to be ordinary partial derivative

Covariant
Derivative, ∇_μ ; notation used ∇_μ
(Levi Civita
connection)

Compatibility Condition : $\nabla_\mu e^i{}_\nu = 0$

- ↔ This condition makes sure that
- Derivative commutes with the contraction.
 - $\omega^i{}_{\mu k} = -e^y{}_k \nabla_\mu e^i{}_y$ (*)
 $= e^y{}_k \Gamma^\rho{}_{\mu y} e^i{}_p - e^y{}_k \partial_\mu e^i{}_y$

Spin connection compatibility determined by Γ, e .

proof

Conversely, we can also define a connection using $(*)$ because its invertible. $\tilde{\Gamma}$

$$\tilde{\Gamma}^{\rho}_{\mu\nu} = e^{\rho}_j \omega^j_{\mu k} e^k_{\nu} + e^{\rho}_j \partial_{\mu} e^j_{\nu}$$

Condition on ω / $\tilde{\Gamma} = \Gamma$?

We impose metric compatibility & torsion free on $\tilde{\Gamma}$.

$$\omega_{\mu j k} := \omega_{\mu}^{\ell}{}_{k} \cdot \delta_{\ell j}$$

Metric Compatible : $\tilde{\nabla}_{\mu} g_{\rho\nu} = 0$.

we know $\tilde{\nabla}_{\mu} (\delta_{ij}) = 0$

$$\text{so; } \tilde{\nabla}_{\mu} (e^{\nu}_i e^{\rho}_j g_{\nu\rho}) = 0$$

Now we use metric compatibility.

$$\tilde{\nabla}_{\mu} (e^{\nu}_i e^{\rho}_j g_{\nu\rho}) \underset{\substack{\uparrow \\ \text{metric compatible}}}{=} -\omega_{\mu j i} - \omega_{\mu i j}$$

$$\Rightarrow \omega_{\mu j i} = -\omega_{\mu i j} \quad \text{Antisymmetric}$$

So, we show that $D_{\gamma}(\delta_{jk}) = 0$

\Rightarrow Also metric compatible.

Torsion FREE Condition of $\tilde{\Gamma}$



ω is also Torsion free.

ie; $T^j_{\mu\nu} = \partial_\mu e^j_\nu - \partial_\nu e^j_\mu + \omega^j_{\mu k} e^k_\nu - \omega^j_{\nu k} e^k_\mu$

Remark|| In 1st order formulation, we will use

$\omega_{\nu jk}$ (antisymmetric in j and k)

but where Torsion does not vanish a priori; but
vanishes as Equation of Motion.

Quantum Gravity

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Lec 5: First order formalism of 3D gravity: action & symmetries, Canonical Analysis.

Assume ω is antisymmetric.

We want to write Riemann Tensor as a function of anti-symmetric spin connection.

$$R_{\mu\nu\rho}{}^{\sigma} v_{\sigma} = (\nabla_{\mu} \nabla_{\nu} - \nabla_{\nu} \nabla_{\mu}) v_{\rho} \\ = e^j_{\rho} (D_{\mu} D_{\nu} - D_{\nu} D_{\mu}) v_j$$

$$v_{\rho} = e^j_{\rho} v_j$$

$$\text{so; } R_{\mu\nu\rho}{}^{\sigma} v_{\sigma} = e^j_{\rho} \cdot F_{\mu\nu jk} \cdot e^{\sigma k} \cdot v_{\sigma}$$

Curvature 2-form of the connection ω .

$$F_{\mu\nu jk} = \partial_{\mu} \omega_{\nu jk} - \partial_{\nu} \omega_{\mu jk} + \omega_{\mu jl} \omega_{\nu}{}^l{}_k - \omega_{\nu jl} \omega_{\mu}{}^l{}_k$$

In 3D

enforce the antisymmetry of $\omega_{\nu jk}$, and of $F_{\mu\nu jk}$ by defining.

$$\omega_{\mu jk} = \epsilon_{jlk} \omega_{\mu}{}^l$$

$$F_{\mu\nu jk} = \epsilon_{jlk} F_{\mu\nu}{}^l$$

ϵ_{ijk} totally antisymmetric ; $\epsilon_{123} = 1$

Σ_{ijk} are also the structure constant for $su(2)$, $so(3)$ Lie Algebra.

$$D_\mu \phi_j = \partial_\mu \phi_j + \Sigma_{jkl} \cdot \omega_\mu^k \phi^l$$

$$F_{\mu\nu}^l = \partial_\mu \omega_\nu^l - \partial_\nu \omega_\mu^l + \Sigma^{lijk} \omega_{\mu j} \omega_{\nu k} \quad \text{Curvature form}$$

$$T_{\mu\nu}^l = \partial_\mu e_\nu^l - \partial_\nu e_\mu^l + \Sigma^{lijk} \omega_{\mu j} e_{\nu k} - \Sigma^{lijk} \omega_{\nu j} e_{\mu k}$$

Torsion form

Einstein-Hilbert Action: in terms of (e, ω)

Start with $S = \frac{1}{\kappa} \int \sqrt{g} \cdot R \cdot d^3x$

$$R = R_{\mu\nu} g^{\mu\nu} = R_{\mu\sigma\nu}{}^\sigma \cdot g^{\mu\nu}$$

$$\sqrt{\det g} = |\det(e_i)| = |\det(e^i)|^{-1}$$

→ det of cotriad.

$\det(e^i)$ is determinant of triad.

$$S = - \int e_{\sigma\lambda} F_{\mu\nu}^l(\omega) \cdot \tilde{\Sigma}^{\sigma\mu\nu} \cdot d^3x$$

- Invariant under
- Rotation in internal indices.
- Translation
- diffeomorphism.

In this formulation we can couple fermions with gravity.

→ tensor density with weight 1.

→ "BF" action. (Topological Theory in any dimensions)

Equation of Motion: Vanishing of curvature $\underline{F_{\mu\nu}^I = 0}$

Vanishing of Torsion $\underline{T_{\mu\nu}^I = 0}$

Diffeo can be written as translation & rotation in 3D gravity.

Symmetries:

- Rotation in the internal index.
- Diffeomorphism.
- Translation

Translation characterised by scalar field with one internal index N^i

$$\delta_N^\tau e_\mu^j = D_\mu N^j$$

$$\delta_N^\tau \omega_\mu^j = 0$$

We can check that the Action is indeed invariant.

Canonical Analysis

Hamiltonian analysis.

* chose a slicing of the manifold.

$\mathcal{M} = \mathbb{R} \times \Sigma$ Σ is 2d surface.

$\mu \rightarrow (0, a)$ ————— (spatial coordinate)
 \hookrightarrow (time component)

Σ is equal time surface.

$$S = \frac{1}{2} \int e_{\sigma\lambda} F^{\lambda}_{\mu\nu} \tilde{\Sigma}^{\sigma\mu\nu} d^3x$$

Notation Use A instead of ω to denote Spin Connection.

$$S = \int e^j_b \partial_a A_{0j} \tilde{\Sigma}^{ab} + \underline{e_{0j}} \frac{1}{2} F^j_{ab} \tilde{\Sigma}^{ab} + \underline{A_0^j} (\partial_a e_{bj} + \Sigma_{jlm} A_a^l e_b^m) \tilde{\Sigma}^{ab} d^2x dt$$

Canonical pair of variable

$$\{A_a^j(x), E_k^b(y)\} = \delta_k^j \delta_a^b \cdot \delta(x, y)$$

Lagrange multiplier

(because don't have time derivative term)

$$E_a^j = \frac{\delta S}{\delta(\partial_0 A_a^j)} = \tilde{\Sigma}^{ab} e^j_b = \text{Densitized Triad.}$$

So the constraints are;

$$\mathcal{F}^j = \frac{1}{2} F^j_{ab} \tilde{\Sigma}^{ab}$$

$$\mathcal{G}^j = \frac{1}{2} T_{ab}^j \tilde{\Sigma}^{ab}$$

$$\mathcal{F}^j = \tilde{\Sigma}_{ab} \left(\partial_0 A_a^j + \frac{1}{2} \Sigma^j_{kl} A_a^k A_b^l \right) \left. \vphantom{\frac{1}{2}} \right\} \begin{array}{l} \text{Flatness constraint} \\ \text{generate translational} \\ \text{symmetry} \end{array}$$

$$\mathcal{G}^j = \partial_a E_a^j + \Sigma_{jlm} A_a^l E^{am} \left. \vphantom{\frac{1}{2}} \right\} \text{Gauss Constraint.}$$

Hamiltonian :

$$H = - \int d^2x (N^i \mathcal{F}_i + N^i \mathcal{G}_i)$$

Totally constrained system.

Reference : ROMANOS article (canonical analysis)