

# Continuum Mechanics

— Shoab Alkhatir

Lee 1 (8<sup>th</sup> Jan, 2021 ; 2:00pm to 3:30pm)

## Continuum Description

Classical Mechanics

$L(\{q_m, \dot{q}_m\}, t)$

—————> Continuum Mechanics.

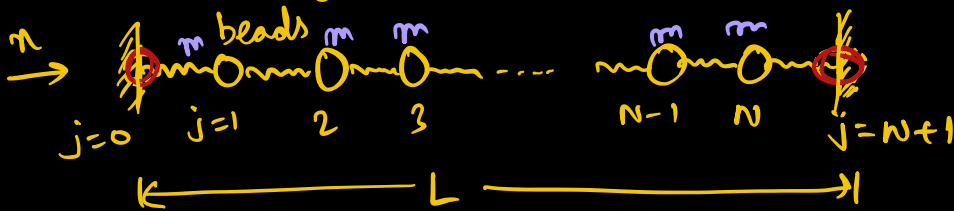
—————>  $\mathcal{L}(q(x), q'(x), \dot{q}(x), t)$

$x$  is continuous parameter.

### Example

Discrete system ———> Continuum limit.

1D string between walls.



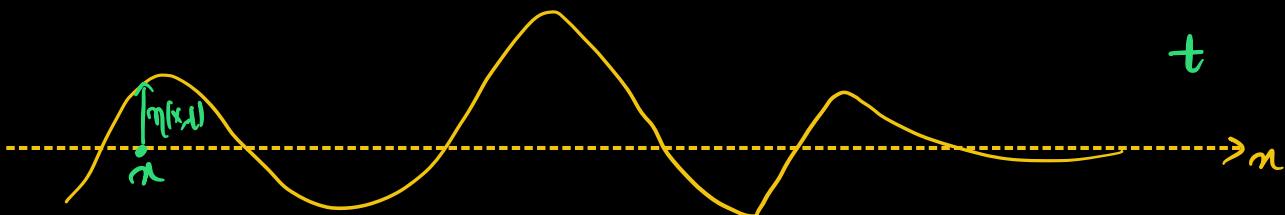
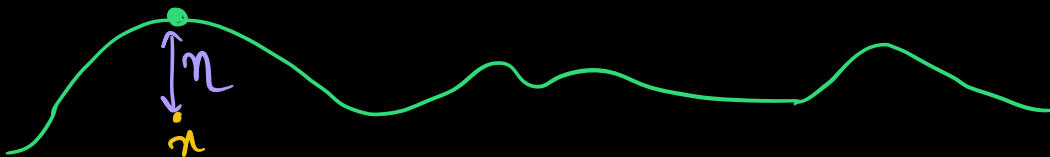
Discrete Counterpart of Wave Equation.

$$\partial_t^2 \eta(x, t) = c^2 \nabla^2 \eta(x, t)$$

Wave Equation

Solutions are well known.

Tells about movement of say string.



$\frac{1}{2}k(\Delta x)^2$  is potential energy of spring in terms of stretch  $(\Delta x)$ .

Assume only longitudinal elongation.

length of whole chain  $= L$ .

$l_0$  = unstretched length of the Hookean spring.

$a$  = spring length at Equilibrium. (Tension)

$x_j^{(eq)} = ja$  : Equilibrium position of beads. ;  $j \in \mathbb{Z}$

we have total  $(N+1)$  springs.

so,  $(N+1)l_0 = L_0 =$  unstretched length.

for simplicity, take  $l_0 \approx a$ . (No tension or compression in string at equilibrium)

Using newton's law:

$$m \ddot{x}_3 = k(x_4 - x_3) - k(x_3 - x_2) \quad x_0 = 0, \quad x_{N+1} = L$$

We will try to do this by lagrangian.

Now we disturb the string;

$$x_j = x_j^{(eq)} + \eta_j(t)$$

$$L = \underbrace{\frac{m}{2} \sum_{j=1}^N \dot{\eta}_j^2}_{\text{Kinetic}} - \underbrace{\frac{k}{2} \sum_{j=0}^N (x_{j+1} - x_j - l_0)^2}_{\text{Potential.}}$$



$$L = \frac{m}{2} \sum_{j=1}^N \dot{\eta}_j^2 - \frac{k}{2} \sum_{j=0}^N (\eta_{j+1} - \eta_j)^2 \quad \left. \vphantom{\sum_{j=0}^N} \right\} \text{for } \frac{a=l_0}{\text{case}}$$

$$\eta_{j+1} - \eta_j - l_0$$

$$= (j+1)a + \eta_{j+1} - ja - \eta_j - l_0$$

$$= (\eta_{j+1} - \eta_j) + (a - l_0)$$

$$\text{so, } L = \frac{m}{2} \sum_{j=1}^N \dot{\eta}_j^2 - \frac{k}{2} \sum_{j=0}^N ((a - l_0) + (\eta_{j+1} - \eta_j))^2$$

$a = \text{eq}^m \text{ length}$  (no motion; but string could be stretched or compressed)

so,

$$L = \frac{m}{2} \sum_{j=1}^N \dot{\eta}_j^2 - \frac{k}{2} \sum_{j=0}^N (\eta_{j+1} - \eta_j)^2 + \text{constant.}$$

The linear term looks like.

$$\frac{k}{2} 2(a - l_0) \sum_j (\eta_{j+1} - \eta_j)$$

$$= k(a - l_0) (\underbrace{\eta_{N+1} - \eta_0}_L) = k(a - l_0) \underbrace{L}_{\text{This is also then constant.}}$$

Now, we apply Lagrange's eq<sup>n</sup> of motion.

$$\frac{d}{dt} \left( \frac{\partial L}{\partial \dot{\eta}_j} \right) = \frac{\partial L}{\partial \eta_j}$$

$\Rightarrow$  *Microscopic eq<sup>n</sup> of motion*  $m \ddot{\eta}_j = k (\eta_{j+1} - 2\eta_j + \eta_{j-1})$  *Discrete Equation of Motion.*

(can also be written by just using Newton's Laws of Motion)

$$\ddot{\eta}_j = \omega^2 (\eta_{j+1} - 2\eta_j + \eta_{j-1})$$

$$\omega = \sqrt{\frac{k}{m}}$$



$$\begin{aligned} m \ddot{\eta}_j &= k (\eta_{j+1} - \eta_j) - k (\eta_j - \eta_{j-1}) \\ &= k (\eta_{j+1} - 2\eta_j + \eta_{j-1}) \end{aligned}$$



Our microscopic model.

$m, k, a$  are microscopic quantities.

Continuum Limit ( $N \rightarrow \infty$ )

Whole string has mass  $M$ ,  
Young's Modulus  $Y$ ,  
Length  $L$

} Macroscopic quantities.  
(finite quantities)

The connection between  $m, k, a$  with  $M, Y, L$  is with  $N$ .

$mN = M$  ,  $aN = L$  (for equilibrium)  
 $L_0 N = L_0$  for unstretched

$$\frac{k}{2} = Y \quad \text{Need to derive.}$$

→ This tells  $x = \frac{1}{N}$   
 ↗ female

50.  $N \rightarrow \infty \Rightarrow k \rightarrow \infty$ .

$$a, m \propto \frac{1}{n}$$

So:  $a_n \rightarrow 0$  for  $n \rightarrow \infty$

but  $k \propto N$  so;  $k \rightarrow \infty$  for  $N \rightarrow \infty$ .

divide  $\Sigma OM$  by  $a$ .

$$\frac{m}{a} \cdot \ddot{x}_j = a \cdot \frac{k}{a^2} \cdot (x_{j+1} - 2x_j + x_{j-1})$$

$$\frac{m}{a} = \mu = \text{mass density} = \frac{M}{L} = \text{finite quantity}$$

$$\mu \cdot \ddot{\eta}_j = a \cdot \frac{k}{a^2} \cdot (\eta_{j+1} - 2\eta_j + \eta_{j-1})$$

In the continuum limit

$$\eta_j(t) \xrightarrow{j \in \mathbb{Z}} \eta(x, t)$$

ie, at every point  
we have particle.

$$\text{ie; } \eta_x(t) \equiv \eta(x, t) \\ x \in \mathbb{R}$$

$$\mu \ddot{\eta}(x, t) = k a \frac{\left[ \frac{(\eta_{j+1} - \eta_j)}{a} - \frac{(\eta_j - \eta_{j-1})}{a} \right]}{a}$$

$$\boxed{\frac{\mu}{a} \ddot{\eta}_j = \frac{k}{a} (\eta_{j+1} - 2\eta_j + \eta_{j-1}))}$$

$$\xrightarrow{a} \frac{\eta'(x+a) - \eta'(x)}{a}$$

$$\frac{\eta_{j+1} - \eta_j}{a} \longrightarrow \text{in continuum limit, it becomes spatial derivative.}$$

$$= \eta'(x)$$

double derivative.

$$\mu \ddot{\eta}(x, t) = Y \eta''(x, t)$$

$$c \equiv \sqrt{Y/\mu} \quad \text{ie; } c^2 = \frac{Y}{\mu}$$

$$\text{Youngs Modulus} = \frac{\text{Force / area}}{\text{strain.}}$$

$$= \frac{(k(\eta_{j+1} - \eta_j) / 1)}{\frac{(\eta_{j+1} - \eta_j)}{a}} \quad \text{strain}$$

nothing to divide



(No area in 1d limit.)

so, area shrinks to point)

$$\text{Youngs Modulus} = k a \Rightarrow \boxed{Y = k a}$$

$$\Rightarrow k = \frac{\gamma}{a} \Rightarrow \boxed{k \propto N}$$

$$N \longrightarrow \infty$$

$$m \longrightarrow 0$$

$$a \longrightarrow 0$$

such that  $Na = L$ ,  $mN = M$  finite.

} Continuum limit.

Forward derivative

$$\eta'_j = \frac{\eta_{j+1} - \eta_j}{a}$$

Backward Derivative

$$\eta'_j = \frac{\eta_j - \eta_{j-1}}{a}$$

Here, we are using this,  
Backward Derivative.  
while defing first  
derivative.

→ While writing double derivative ; we  
used Forward derivative.

$$\frac{d}{dx} \left( \frac{d\eta}{dx} \right) = \frac{d}{dx} \left( \frac{\eta_{j+1} - \eta_j}{a} \right)$$

→ using forward derivative.

$$= \frac{\eta'_{j+1} - \eta'_j}{a} \rightarrow \text{using forward.}$$

$$= \frac{(\eta_{j+2} - \eta_{j+1}) - (\eta_{j+1} - \eta_j)}{a^2}$$

$$= \frac{\eta_{j+2} - 2\eta_{j+1} + \eta_j}{a^2}$$

In continuum limit; we could shift particle by 1 because there are  $\infty$  particles.

$$\text{so, } \frac{\eta_{j+2} - 2\eta_{j+1} + \eta_j}{a^2} \rightarrow \frac{\eta_{j+1} - 2\eta_j + \eta_{j-1}}{a^2}$$

This is more symmetric definition.

So we work with this.

B; Backward derivative, F; Forward derivative.

for double derivative, we could use

BB, FF, BF, FB

$\rightarrow$  all equivalent in continuum limit.

$$L = \frac{m}{2} \sum_{j=1}^N \dot{\eta}_j^2 - \frac{k}{2} \sum_{j=0}^N (\eta_{j+1} - \eta_j)^2$$

$\rightarrow$  should be finite quantity

$$L = \frac{a}{2} \left[ \frac{m}{a} \sum \dot{\eta}_j^2 - ka \sum \left( \frac{\eta_{j+1} - \eta_j}{a} \right)^2 \right]$$

Taking continuum limit.

$$L = \frac{a}{2} \sum \left[ \mu \cdot \dot{\eta}^2(x, t) - \gamma \cdot \left( \frac{d\eta(x, t)}{dx} \right)^2 \right]$$

$a$  is actually  $\Delta x$

so; The summation becomes integral.

$$L = \Delta x \sum \frac{1}{2} \left[ \mu \cdot \dot{\eta}^2(x, t) - \gamma \cdot \left( \frac{d\eta}{dx} \right)^2 \right]$$

$$= \int dx \cdot \left[ \frac{\mu}{2} \dot{\eta}^2(x, t) - \frac{\gamma}{2} \left( \frac{d\eta}{dx} \right)^2 \right]$$

$\mathcal{L}(\eta, \eta', \dot{\eta}, t)$  : Lagrangian density

$$\eta' = \frac{\partial \eta}{\partial x}, \quad \dot{\eta} = \frac{\partial \eta}{\partial t},$$

$$L = \int dx \mathcal{L}(\eta, \eta', \dot{\eta}, t)$$

$$\mathcal{L} = \frac{\mu}{2} \dot{\eta}^2(x, t) - \frac{\gamma}{2} \eta'^2(x, t)$$

$x, t$  are independent variables.

$$\frac{\partial \eta}{\partial x} = \frac{d\eta}{dx}, \quad \frac{\partial \eta}{\partial t} = \frac{d\eta}{dt}$$

because  $x$  &  $t$  are independent variables.

if  $\eta(x(t), t)$

Then

$$\frac{d\eta}{dt} \neq \frac{\partial \eta}{\partial t}$$

but for  $\eta(x, t) \Rightarrow \frac{d\eta}{dt} = \frac{\partial \eta}{\partial t}$

$$\frac{d\eta}{dx} = \frac{\partial \eta}{\partial x}$$

$$L = \int \mathcal{L}(\eta(x, t), \eta', \dot{\eta}, t) dx$$

In Continuum Limit

$$L(\{a_j, \dot{a}_j\}, t) \longrightarrow \mathcal{L}(\eta(x, t), \eta', \dot{\eta}, t)$$

$$j \longrightarrow x$$

$$a_j(t) \longrightarrow \eta(x, t)$$

$$\dot{a}_j(t) \longrightarrow \frac{d\eta}{dx}, \quad \frac{d\eta}{dt}$$

$$\begin{array}{c} L(a, \dot{a}, t) \\ t \end{array}$$

$$\frac{d}{dt} \left( \frac{\partial L}{\partial \dot{a}} \right) = \frac{\partial L}{\partial a}$$

Lagrange's Equation.

$$\begin{array}{c} \mathcal{L}(\eta, \eta', \dot{\eta}, x, t) \\ x, t \end{array}$$

$$\frac{d}{dt} \left( \frac{\partial \mathcal{L}}{\partial \dot{\eta}} \right) + \frac{d}{dx} \left( \frac{\partial \mathcal{L}}{\partial \eta'} \right) = \frac{\partial \mathcal{L}}{\partial \eta}$$

Euler Lagrange Equations.

(Can be derived)

$$\text{we have } \mathcal{L} = \frac{\mu}{2} \dot{\eta}^2 - \frac{\gamma}{2} \eta'^2 \Rightarrow \underline{\mu \ddot{\eta} - \gamma \eta'' = 0} \quad \text{Wave Equation.}$$



Recap:



$$\eta(0, t) = 0 = \eta(L, t)$$

$\leftarrow a \rightarrow$  eq<sup>n</sup> position of spring

$$m \ddot{\eta}_j = k(\eta_{j+1} - 2\eta_j + \eta_{j-1})$$

$\eta_j$  are deviation from equilibrium position.

$$\eta_j = \eta_j^{(eq)} + \eta_j(t) \quad \text{ie; } \eta_j - \eta_j^{(eq)} \equiv \eta_j(t)$$

If we have friction in the medium; then after long time the particle will come to rest at positions  $\eta_j^{(eq)}$

$$\eta_j^{(eq)} = ja$$

Continuum Limit :  $\mu \partial_t^2 \eta(x, t) = \gamma \partial_x^2 \eta(x, t)$

$m \rightarrow 0, N \rightarrow \infty, \text{ such that } mN = M = \mu \cdot L = \text{finite.}$   
 $a \rightarrow 0$

spring constant  $k$  ;  $ka = \gamma = \text{finite}$

$$a \propto \frac{1}{N} \Rightarrow k \propto N$$

$\mathcal{L}(\eta, \dot{\eta}, \eta', x, t)$  : Lagrangian Density.

Here  $\mathcal{L}(\eta, \dot{\eta}, \eta', x, t) = \frac{\mu}{2} (\partial_t \eta)^2 - \frac{\gamma}{2} (\partial_x \eta)^2$  (\*)

(No explicit  $x$  and  $t$  dependence here)

$\eta_x \equiv \partial_x \eta$  ,  $\eta_t \equiv \partial_t \eta$  (Notation)

Analogous of L - eq<sup>n</sup> in discrete particle dynamics is called Euler-Lagrange's eq<sup>n</sup>.

$$\frac{d}{dt} \left( \frac{\partial \mathcal{L}}{\partial \eta_t} \right) + \frac{d}{dx} \left( \frac{\partial \mathcal{L}}{\partial \eta_x} \right) = \frac{\partial \mathcal{L}}{\partial \eta} \quad \text{E-L Equation.}$$

Using the  $\mathcal{L}$  (\*), the E-L. eq<sup>n</sup> gives.


$$\Rightarrow \mu \partial_t^2 \eta = \gamma \partial_x^2 \eta.$$

Can have different boundary conditions:

ex)  $\eta(0, t) = 0 = \eta(L, t)$

Boundary Condition

$x=0$   $x=L$




$\eta(0, t) = 0$

$$\left. \frac{\partial \eta}{\partial x} \right|_{x=L} = 0$$

$$\begin{aligned} m \ddot{\eta}_N &= -k (\eta_N - \eta_{N-1}) \\ &= -ka \frac{(\eta_N - \eta_{N-1})}{a} \end{aligned}$$

More general situation.

$x=0$   $x=L$



$\eta(0, t) = 0$

$$m \ddot{\eta}_N = -ka \cdot \left( \frac{\eta_N - \eta_{N-1}}{a} \right) + F$$

$$m \ddot{\eta}_n = -\gamma \cdot \frac{d\eta}{dx} + F \quad \Bigg| \quad \text{continuum limit}$$

$m \ddot{\eta}_n \rightarrow 0$  in continuum limit, if  $m=0$

so,

$$\boxed{-\gamma \cdot \frac{d\eta}{dx} \Big|_{x=L} + F = 0} \quad \text{Boundary Condition.}$$

$$\boxed{F = \gamma \cdot \frac{d\eta}{dx} \Big|_{x=L}}$$

$\Rightarrow$

$$\boxed{\frac{d\eta}{dx} \Big|_{x=L} = \frac{F}{\gamma}}$$

Individual  $\eta_j$  are microscopic quantity.


so, difference is also microscopic.

$$\left. \begin{array}{l} \frac{\eta_{j+1} - \eta_j}{a} \} \rightarrow \text{microscopic} \\ \} \text{ so can be finite.} \end{array} \right\}$$

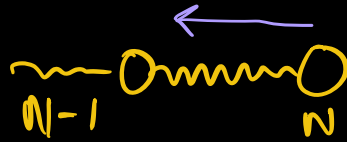
but for  $\ddot{\eta}_n$ , we don't expect this to go infinity.

for this to go to  $\infty$ , we need infinite force.

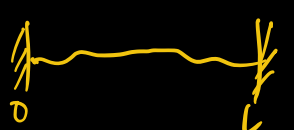
so, we don't expect  $\eta_n$  &  $\ddot{\eta}_n$  to blow up.

$$-k(\eta_j - \eta_{j-1}) \quad \longleftrightarrow \quad +k(\eta_{j+1} - \eta_j)$$


for last particle



No mass at right.



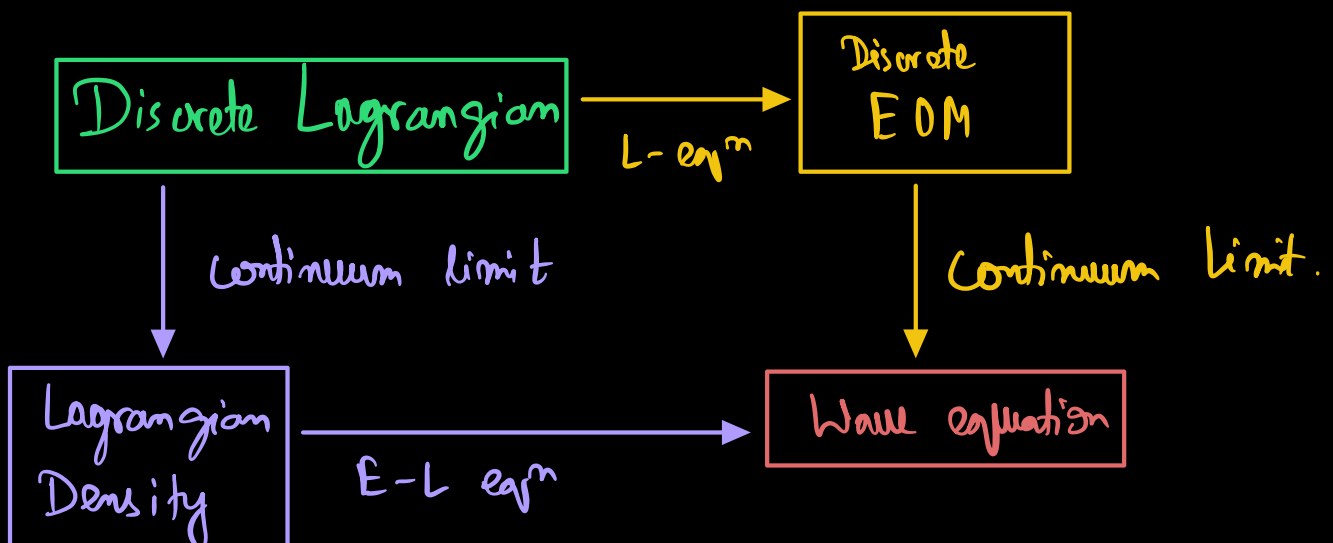
$$\eta(x,t) = \sum A_n \sin k_n x e^{\lambda_n t}$$

$$\mu \ddot{\eta} = Y \eta'' \rightarrow \sin kx \cdot e^{\lambda t} \Rightarrow kL = n\pi \Rightarrow k_n = n\pi/L$$

$$\mu \lambda^2 = -Y k^2 \Rightarrow \lambda_n = i \sqrt{\frac{Y}{\mu}} \cdot k_n$$

$$\text{So; } \eta(x,t) = \sum A_n \sin(k_n x) \cdot e^{i \sqrt{\frac{Y}{\mu}} \cdot k_n \cdot t}$$

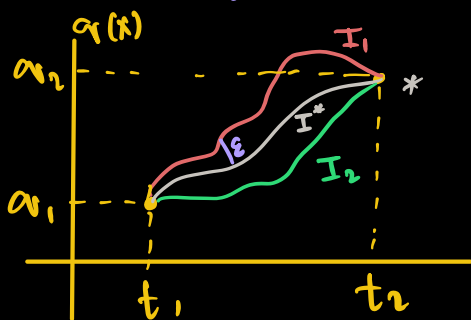
Using boundary conditions, we can find  $A_n$ 's  
and hence  $\eta(x,t)$  is obtained.



Continuum limit

$$\begin{aligned} t &\longrightarrow \eta, t \\ j &\longrightarrow \eta \\ q_j(t) &\longrightarrow \eta(x, t) \\ \dot{q}_j(t) &\longrightarrow \dot{\eta}, \eta' \end{aligned}$$

## Principle of Least Action



$$I = \int_{t_1}^{t_2} L(q, \dot{q}, t) dt \quad \text{Action.}$$

Which path  $q(t)$  will extremize  $I$  ?  
The physical path !

$I(q, \dot{q}, t)$  is a functional, depends on  $q(t)$ .

let  $q(t)$  be the optimum path or the physical path, then we consider the deviation

$$q(t) \longrightarrow q(t) + \delta q(t).$$

$\delta I = I_1 - I_*$  should be zero

$$\delta I = \mathcal{O}(\epsilon^2) + \mathcal{O}(\epsilon)$$

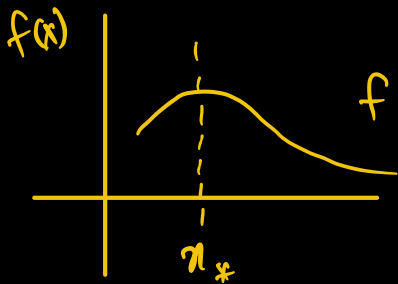
$$\left. \frac{\delta I}{\epsilon} \longrightarrow 0 \text{ as } \epsilon \longrightarrow 0 \right\} \Rightarrow \text{that } * \text{ is the extremum path.}$$

$$\delta I = A\epsilon + B\epsilon^2 + \dots$$

$$\frac{\delta I}{\epsilon} = A + B\epsilon + \dots$$

$$\text{in limit } \epsilon \rightarrow 0 \quad \left. \frac{\delta I}{\epsilon} \right|_{\epsilon \rightarrow 0} = A$$

so; for the path to be extremum, we will need  $A = 0$ .



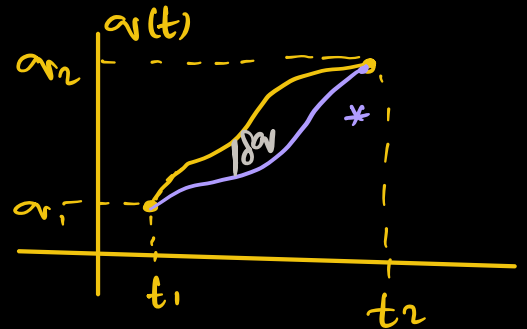
$$f(x_* + \epsilon) = f(x_*) + \epsilon f'(x_*) + \epsilon^2 f''(x_*) + \dots$$

$$\frac{f(x_* + \epsilon) - f(x_*)}{\epsilon} = \underbrace{f'(x_*) + \epsilon f''(x_*) + \dots}_{\substack{\text{This has to be} \\ \text{zero.}}}$$

$$q \rightarrow q + \delta q \Rightarrow L(q) \rightarrow L(q + \delta q)$$

$$\delta I = \int_{t_1}^{t_2} [L(q + \delta q) - L(q)] dt$$

$\underbrace{\hspace{10em}}_{q_1 \text{ path.}} \quad \underbrace{\hspace{10em}}_{q_2 \text{ path.}}$



doing the analysis in path space.

$$\delta q(t_1) = 0 = \delta q(t_2)$$

: End point variations are fixed.

$$\delta \dot{q} = \frac{d}{dt} \delta q$$

$$q_1 = q + \delta q$$

Actually.

$$\delta I = \int_{t_1}^{t_2} [L(q_r + \delta q_r, \dot{q} + \delta \dot{q}, t) - L(q_r, \dot{q}, t)] dt$$

$$= \int_{t_1}^{t_2} \left[ \cancel{L(q_r, \dot{q}, t)} + \frac{\partial L}{\partial q_r} \delta q_r + \frac{\partial L}{\partial \dot{q}} \delta \dot{q} - \cancel{L(q_r, \dot{q}, t)} \right] dt$$

$$= \int_{t_1}^{t_2} \left[ \frac{\partial L}{\partial q_r} \delta q_r + \frac{\partial L}{\partial \dot{q}} \delta \dot{q} \right] dt$$

$$= \int_{t_1}^{t_2} \left( \left[ \delta q_r \cdot \frac{\partial L}{\partial q_r} \right] + \left[ \frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}} \delta q_r \right) - \frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}} \right) \cdot \delta q_r \right] \right) dt$$

$$\text{Note } \int_{t_1}^{t_2} dt \cdot \frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}} \delta q_r \right) = \left. \frac{\partial L}{\partial \dot{q}} \delta q_r \right|_{t_1}^{t_2}$$

$$= \left. \frac{\partial L}{\partial \dot{q}} \right|_{t_2} \cdot \delta q_r(t_2) - \left. \frac{\partial L}{\partial \dot{q}} \right|_{t_1} \cdot \delta q_r(t_1)$$

$$= 0$$

$$\delta I = \int_{t_1}^{t_2} dt \cdot \delta q_r \left[ \frac{\partial L}{\partial q_r} - \frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}} \right) \right]$$

$$\delta I = 0 \implies \frac{\partial L}{\partial q_r} - \frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}} \right) = 0$$

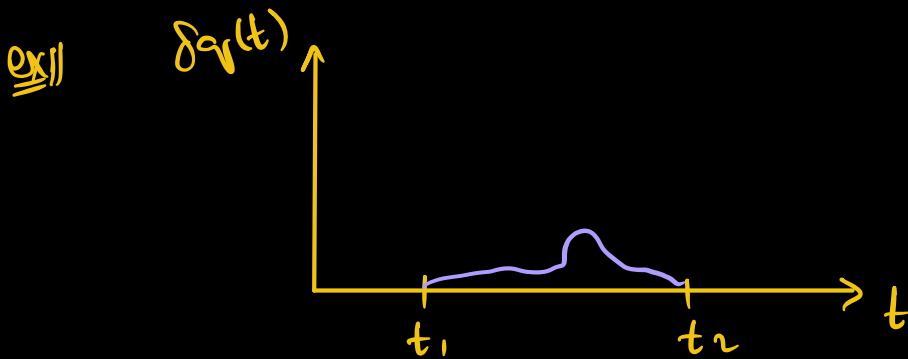
Note  $\int_{x_1}^{x_2} f(x) dx = 0 \not\Rightarrow f(x) = 0$

Here we could conclude because of  $\delta q$ .

and the deviation  $\delta q$  is arbitrary.

so,  $\delta I = 0 \quad \forall \delta q$  arbitrary.

so, we get  $\frac{\partial L}{\partial q} - \frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}} \right) = 0$



## Derivation of Euler-Lagrange's Equation.

given  $\mathcal{L}(q(x,t), q', q, x, t)$  Lagrangian density

$$L = \int \mathcal{L} dx$$

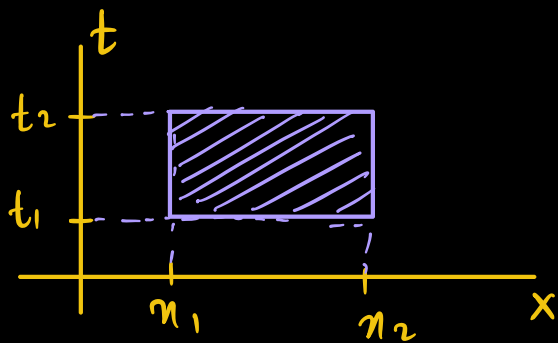
$L \Rightarrow$  Lagrangian.

$\mathcal{L} \Rightarrow$  Lagrangian density.

$$\text{Action} = \int L \cdot dt = \int \mathcal{L} \cdot dx dt = I$$

Need to find  $\delta I$ .





$\delta\eta$

Surface  $\eta(x,t)$  changed to  $\eta(x,t) + \delta\eta(x,t)$ .

$$\eta(x,t) \longrightarrow \eta(x,t) + \delta\eta(x,t)$$

$$\delta I = 0 \implies \eta^*(x,t).$$

$$\left. \delta\eta(x,t) \right|_{\square} = 0 \quad \text{ie; } \delta\eta \text{ is zero on boundary of integration.}$$

$$\text{ie; } I = \int_{\Omega} dx dt \mathcal{L}$$

$$\text{so; } \delta\eta|_{\partial\Omega} = 0.$$

$$\begin{aligned} \delta I &= \int dx dt \left[ \mathcal{L}(\eta + \delta\eta, \eta' + \delta\eta', \dot{\eta} + \delta\dot{\eta}) - \mathcal{L}(\eta, \eta', \dot{\eta}) \right] \\ &= \int dx dt \left[ \mathcal{L}(\eta, \eta', \dot{\eta}) + \frac{\partial \mathcal{L}}{\partial \eta} \delta\eta + \frac{\partial \mathcal{L}}{\partial \eta'} \delta\eta' + \frac{\partial \mathcal{L}}{\partial \dot{\eta}} \delta\dot{\eta} \right. \\ &\quad \left. - \mathcal{L}(\eta, \eta', \dot{\eta}) \right]. \end{aligned}$$

we are doing double Taylor expansion.

Notation:  $\delta\eta' = \frac{d}{dx} \delta\eta$ ,  $\delta\dot{\eta} = \frac{d}{dt} \delta\eta$

$$\delta I = \int dx dt \left[ \frac{\partial \mathcal{L}}{\partial \eta} \delta\eta + \frac{\partial \mathcal{L}}{\partial \eta'} \delta\eta' + \frac{\partial \mathcal{L}}{\partial \dot{\eta}} \delta\dot{\eta} \right]$$

Note we also have terms like

$$\frac{\partial^2 \mathcal{L}}{\partial \eta' \partial \dot{\eta}} \delta\eta' \delta\dot{\eta} + \dots$$

but we don't keep the higher order terms.

so,

$$\delta I = \int dx dt \left[ \frac{\partial \mathcal{L}}{\partial \eta} \delta\eta + \frac{\partial \mathcal{L}}{\partial \eta'} \delta\eta' + \frac{\partial \mathcal{L}}{\partial \dot{\eta}} \delta\dot{\eta} + \mathcal{O}(\delta\eta^2) \right]$$

$$= \int dx dt \left[ \frac{\partial \mathcal{L}}{\partial \eta} \delta\eta + \underbrace{\frac{d}{dx} \left[ \frac{\partial \mathcal{L}}{\partial \eta'} \delta\eta \right]}_{\substack{\text{boundary term} \\ \text{vanishes}}} - \frac{d}{dx} \left[ \frac{\partial \mathcal{L}}{\partial \eta'} \right] \delta\eta + \underbrace{\frac{d}{dt} \left[ \frac{\partial \mathcal{L}}{\partial \dot{\eta}} \delta\eta \right]}_{\substack{\text{boundary term} \\ \text{vanishes}}} - \frac{d}{dt} \left[ \frac{\partial \mathcal{L}}{\partial \dot{\eta}} \right] \delta\eta + \mathcal{O}(\delta\eta^2) \right]$$

doing just dx integral.

$$\int dx \rightarrow \left. \frac{\partial \mathcal{L}}{\partial \eta'} \delta\eta \right|_{x_1}^{x_2} = 0$$

→ similarly this term drops out after dt integral.

$$\delta I = \int dx dt \cdot \delta \eta \left[ \frac{\partial \mathcal{L}}{\partial \eta} - \frac{d}{dx} \left( \frac{\partial \mathcal{L}}{\partial \eta'} \right) - \frac{d}{dt} \left( \frac{\partial \mathcal{L}}{\partial \dot{\eta}} \right) + \mathcal{O}(\delta \eta) \right]$$

$\delta \eta(x, t)$  is completely arbitrary.

$$\text{so, } \delta I = 0 \Rightarrow \frac{\partial \mathcal{L}}{\partial \eta} - \frac{d}{dx} \left( \frac{\partial \mathcal{L}}{\partial \eta'} \right) - \frac{d}{dt} \left( \frac{\partial \mathcal{L}}{\partial \dot{\eta}} \right) = 0$$


---

Question)  $\mathcal{L}(\eta, \eta', \eta'', x, t)$

$$\mathcal{L}(\ddot{\eta}, \ddot{\eta}, \eta', \dots)$$

when we have higher order derivatives.

# Homework. Continuum Mechanics

— Shoab Akhtar

Given  $d+1$  dimensional spacetime

$$\mu \in \{0, 1, 2, \dots, d\}$$

$\downarrow$  time index  
spatial index.

$x^\mu$

Given the Lagrangian Density

$$\mathcal{L}(\eta(x^\sigma), \partial_\mu \eta, \partial_\mu \partial_\nu \eta, \dots, \partial_\mu \partial_{\mu_1} \dots \partial_{\mu_n} \eta, x^\mu)$$

We write the action

$$I[\eta] = \int_{\Omega} \mathcal{L} \cdot d^{d+1}x$$

$\Omega$  is the spacetime region over which we integrate.

Finding Generalized Euler Lagrange equation.

$$\eta \rightarrow \eta + \delta\eta \quad \text{s.t.} \quad \delta\eta|_{\partial\Omega} = 0$$

under this infinitesimal transformation, let's find the change in action  $\delta I$

$$\delta I = \int_{\Omega} \left[ \frac{\partial \mathcal{L}}{\partial \eta} \delta\eta + \frac{\partial \mathcal{L}}{\partial \eta_{\mu_1}} \delta\eta_{\mu_1} + \frac{\partial \mathcal{L}}{\partial \eta_{\mu_1 \mu_2}} \delta\eta_{\mu_1 \mu_2} + \dots + \frac{\partial \mathcal{L}}{\partial \eta_{\mu_1 \mu_2 \dots \mu_n}} \delta\eta_{\mu_1 \mu_2 \dots \mu_n} \right] d^{d+1}x$$

NOTATION:

$$\eta_{\mu_1 \mu_2 \dots \mu_j} \equiv \partial_{\mu_1} \partial_{\mu_2} \dots \partial_{\mu_j} \eta$$

$$\delta I = \int_{\Omega} \left[ \frac{\partial \mathcal{L}}{\partial \eta} \delta \eta + \sum_{j=1}^N \frac{\partial \mathcal{L}}{\partial \eta_{\mu_1 \dots \mu_j}} \cdot \delta \eta_{\mu_1 \dots \mu_j} \right] d^{d+1}x$$

consider the piece

$$\int_{\Omega} \frac{\partial \mathcal{L}}{\partial \eta_{\mu_1 \dots \mu_j}} \cdot \delta \eta_{\mu_1 \dots \mu_j} \cdot d^{d+1}x$$

$$= \int_{\Omega} \left[ \partial_{\mu_1} \left( \frac{\partial \mathcal{L}}{\partial \eta_{\mu_1 \dots \mu_j}} \cdot \delta \eta_{\mu_2 \dots \mu_j} \right) - \partial_{\mu_1} \left( \frac{\partial \mathcal{L}}{\partial \eta_{\mu_1 \dots \mu_j}} \right) \delta \eta_{\mu_2 \dots \mu_j} \right] d^{d+1}x$$

$$= \underbrace{\int_{\Omega} \partial_{\mu_1} \left( \frac{\partial \mathcal{L}}{\partial \eta_{\mu_1 \dots \mu_j}} \cdot \delta \eta_{\mu_2 \dots \mu_j} \right) d^{d+1}x}_{\Downarrow} - \int_{\Omega} \partial_{\mu_1} \left( \frac{\partial \mathcal{L}}{\partial \eta_{\mu_1 \dots \mu_j}} \right) \delta \eta_{\mu_2 \dots \mu_j} d^{d+1}x$$

This becomes the boundary  
integral over  $\partial \Omega$

$\Rightarrow$  so it vanishes

$$\text{if } \delta \eta_{\mu_2 \dots \mu_j} \Big|_{\partial \Omega} = 0$$

$$= - \int_{\Omega} \partial_{\mu_1} \left( \frac{\partial \mathcal{L}}{\partial \eta_{\mu_1 \dots \mu_j}} \right) \delta \eta_{\mu_2 \dots \mu_j} \eta_{\mu_1} d^{d+1}x$$

We again re-iterate the steps

$$= (-1)^j \int_{\Omega} \partial_{\mu_1} \partial_{\mu_2} \dots \partial_{\mu_j} \left( \frac{\partial \mathcal{L}}{\partial \eta_{\mu_1 \dots \mu_j}} \right) \cdot \delta \eta \cdot d^{d+1}x$$

---

NOTATION :  $\partial_{\mu_1 \dots \mu_j} = \partial_{\mu_1} \partial_{\mu_2} \dots \partial_{\mu_j}$

Using this we get

$$\delta I = \int_{\Omega} \left[ \frac{\partial \mathcal{L}}{\partial \eta} + \sum_{j=1}^N (-1)^j \partial_{\mu_1 \dots \mu_j} \left( \frac{\partial \mathcal{L}}{\partial \eta_{\mu_1 \dots \mu_j}} \right) \right] \delta \eta \cdot d^{d+1}x$$

under the restriction.

$$\delta \eta_{\mu_1 \dots \mu_j} \Big|_{\partial \Omega} = 0 \quad \text{where } j \in \{1, \dots, N-1\}$$

$$\text{and } \delta \eta \Big|_{\partial \Omega} = 0$$

so  $\delta I = 0$  for such arbitrary  $\delta \eta$

gives us the Generalized Euler Lagrange Equation.

$$\boxed{\frac{\partial \mathcal{L}}{\partial \eta} + \sum_{j=1}^N (-1)^j \partial_{\mu_1 \dots \mu_j} \left( \frac{\partial \mathcal{L}}{\partial \eta_{\mu_1 \dots \mu_j}} \right) = 0}$$



lec 3 (15<sup>th</sup> Jan, 2021 ; 2:00 pm to 3:30 pm)

given  $\mathcal{L}(\eta, \eta_x, \eta_t, x, t) \rightarrow \frac{d}{dx} \left( \frac{\partial \mathcal{L}}{\partial \eta_x} \right) + \frac{d}{dt} \left( \frac{\partial \mathcal{L}}{\partial \eta_t} \right) = \frac{\partial \mathcal{L}}{\partial \eta}$

$$\mathcal{L} = \frac{\mu}{2} (\partial_t \eta)^2 - \frac{Y}{2} (\partial_x \eta)^2$$

Independent variable  $x, t$ .

we could have  $x, y, z, t \rightarrow x_\mu : \mu = 0, 1, 2, 3$ .

$$\vec{E}(\vec{x}, t), \quad \vec{B}(\vec{x}, t) \quad (\vec{x}, t) = x^\mu$$

$$\begin{pmatrix} E_1(x^\mu) \\ E_2(x^\mu) \\ E_3(x^\mu) \end{pmatrix}$$

$$x^\mu = x, y, z, t$$

$$\sum_\mu \frac{d}{dx_\mu} \left( \frac{\partial \mathcal{L}}{\partial \eta_\mu} \right) = \frac{\partial \mathcal{L}}{\partial \eta}$$

notation:  $\eta_\mu = \frac{\partial \eta}{\partial x^\mu}$

(For now, don't differentiate between contravariant & covariant)

$$x_\mu = x, t, \quad \eta_\rho$$

$\rho$  stands for the components of the field  $\eta$   
 $\rho = 1, 2, 3$

$$\begin{pmatrix} E_1 \\ E_2 \\ E_3 \end{pmatrix} \equiv [\eta_\rho]$$

ie;  $\eta_1$  is like  $E_1$   
 $\eta_2$  " "  $E_2$   
 $\eta_3$  " "  $E_3$

for each  $\rho$ , we have Euler-Lagrange equations.

for each  $p$

$$\sum_{\mu} \frac{d}{dx_{\mu}} \left( \frac{\partial \mathcal{L}}{\partial \eta_{p,\mu}} \right) = \frac{\partial \mathcal{L}}{\partial \eta_p}$$

$\mu = x, t$

NOTATION:  $\eta_{p,\mu} \equiv \frac{\partial \eta_p}{\partial x_{\mu}}$

$p$  stands for components of the vector.

when  $\mathcal{L}(\eta_p, \eta_{p,\mu}, x_{\mu})$

$$\mathcal{L}(\eta, \eta_x, \eta_{xx}, \dots)$$

$-\frac{d}{dx^2} \left( \frac{\partial \mathcal{L}}{\partial \eta_{x,x}} \right)$  term comes on LHS

$$\mathcal{L}(\dots, \eta_{xx} + \delta \eta_{xx}, \dots)$$

expanding gives

$$\frac{\partial \mathcal{L}}{\partial \eta_{xx}} \delta \eta_{xx} = \frac{d}{dx} \left( \frac{\partial \mathcal{L}}{\partial \eta_{xx}} \delta \eta_x \right) - \frac{d}{dx} \left( \frac{\partial \mathcal{L}}{\partial \eta_{xx}} \right) \delta \eta_x$$

This becomes boundary term after integral.

because it gives  $\frac{\partial \mathcal{L}}{\partial \eta_{xx}} \cdot \delta \eta_x \Big|_{\text{boundary}}$

$\delta \eta_x = 0$  on boundary.

$$-\frac{d}{dx} \left( \frac{\partial \mathcal{L}}{\partial \eta_{xx}} \right) \delta \eta_x = \frac{d}{dx} \left( \frac{d}{dx} \left( \frac{\partial \mathcal{L}}{\partial \eta_{xx}} \right) \delta \eta \right) + \frac{d^2}{dx^2} \left( \frac{\partial \mathcal{L}}{\partial \eta_{xx}} \right) \cdot \delta \eta$$

becomes boundary term after integration & vanishes.  $\delta \eta = 0$  on boundary.



so; The extra term we get is

$$\int dx dt \left[ \dots + \frac{d^2}{dx^2} \left( \frac{\partial \mathcal{L}}{\partial \eta_{xx}} \right) + \dots \right] \delta \eta$$

$$\text{i.e.} \int dx dt \left[ \frac{\partial \mathcal{L}}{\partial \eta} - \frac{d}{dx} \left( \frac{\partial \mathcal{L}}{\partial \eta_x} \right) + \frac{d^2}{dx^2} \left( \frac{\partial \mathcal{L}}{\partial \eta_{xx}} \right) \right] \delta \eta = 0$$

→ This is zero.

$$\frac{\partial \mathcal{L}}{\partial \eta} = \frac{d}{dx} \left( \frac{\partial \mathcal{L}}{\partial \eta_x} \right) - \frac{d^2}{dx^2} \left( \frac{\partial \mathcal{L}}{\partial \eta_{xx}} \right)$$

$$\eta \rightarrow \eta + \delta \eta(x)$$

hold term

$$\eta + \alpha \cdot f(x)$$

$$\delta \eta(x) = \alpha \cdot f(x)$$

$\alpha$  small number.

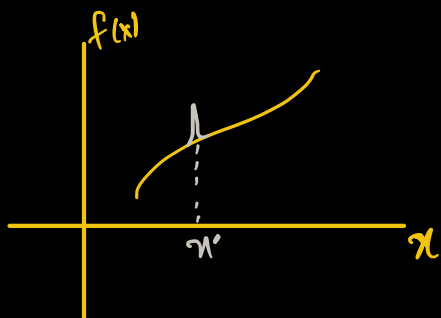
$$I = I[f]$$

Functional.

ex||  $I[f] = \int_0^1 f^2(x) dx$

we want to find  $\frac{\delta I}{\delta f(x)} = g(x)$

→ This will be some function of  $x$ .



$$\mathcal{L}(\eta, \eta_x, \dots)$$

$$\eta \rightarrow \eta + \delta \eta(x)$$

$$\frac{\delta I[f]}{\epsilon} = \lim_{\epsilon \rightarrow 0} \frac{1}{\epsilon} \int \left\{ \left[ f(x) + \underbrace{\epsilon \cdot \delta(x-x')}_{\substack{\text{changing by small} \\ \text{amount at a point } x'}} \right]^2 - [f^2(x)] \right\} dx$$

definition of functional derivative

$$\frac{\delta I[f]}{\delta f(x')}$$

$$= \lim_{\epsilon \rightarrow 0} \frac{1}{\epsilon} \int dx \left[ f^2 + 2\epsilon f(x) \cdot \delta(x-x') + \cancel{\epsilon^2 \cdot \delta^2(x-x')} - f^2 \right] \xrightarrow{\text{neglect.}}$$

$$\int \delta(x-x') dx = 1$$

$$\begin{aligned} \int (\delta(x-x'))^2 dx &= \int \delta(x-x') \delta(x-x') dx \\ &= \delta(0) \end{aligned}$$

$$\text{so, } \int \epsilon^2 \cdot \delta^2(x-x') dx = \epsilon^2 \delta(0)$$

$$\frac{1}{\epsilon} \int \epsilon^2 \cdot \delta^2(x-x') dx = \epsilon \cdot \delta(0)$$

$$\lim_{\epsilon \rightarrow 0} \epsilon \cdot \delta(0) = 0$$

$\epsilon$  is stronger 0

as compared to  $\delta(0)$  being infinity.

choice of  $\epsilon$  is arbitrary

so make it arbitrarily small

$$\text{so that } \epsilon \delta(0) = 0$$

$$\frac{\delta I[f]}{\epsilon} = 2 \cdot f(x') \quad \text{ie} \quad \frac{\delta I}{\delta f(x')} = 2 \cdot f(x')$$

$$\frac{\delta I}{\delta f(x')} = \lim_{\epsilon \rightarrow 0} \frac{I[f(x) + \epsilon \delta(x-x')] - I[f(x)]}{\epsilon} \quad \left. \vphantom{\frac{\delta I}{\delta f(x')}} \right\} \text{Functional differentiation.}$$

Example 1]  $I = \int f^2 dx \rightarrow \frac{\delta I}{\delta f(x')} = 2f(x')$

Example 2]  $I = \int \left(\frac{df}{dx}\right)^2 dx \rightarrow \frac{\delta I}{\delta f(x')} = -2 \cdot \frac{d^2}{dx'^2} f(x')$

$$I[f + \epsilon \delta(x-x')] = \int \left(\frac{d}{dx}(f(x) + \epsilon \delta(x-x'))\right)^2 dx$$

$$\delta I = \int \left[ \left(\frac{d}{dx}(f(x) + \epsilon \delta(x-x'))\right)^2 - \left(\frac{d}{dx}f(x)\right)^2 \right] dx$$

$$= \int \left[ (f'(x) + \epsilon \delta'(x-x'))^2 - (f'(x))^2 \right] dx$$

$$= \int \left[ f'(x)^2 + \epsilon^2 (\delta'(x-x'))^2 + 2f'(x) \epsilon \delta'(x-x') - (f'(x))^2 \right] dx$$

$$= \int \left[ \epsilon^2 \cdot (\delta'(x-x'))^2 + 2\epsilon f'(x) \delta'(x-x') \right] dx$$

$$\Rightarrow \frac{\delta I}{\delta f(x')} = \frac{1}{\epsilon} \delta I = \int 2f'(x) \delta'(x-x') dx$$

$$= 2 \left[ f'(x) \delta(x-x') \Big|_{-\infty}^{+\infty} - \int_{-\infty}^{+\infty} f''(x) \cdot \delta(x-x') dx \right]$$

$$= 2 \cdot \left[ 0 - \frac{d^2 f}{dx'^2} \Big|_{x'} \right]$$

$$= -2 \frac{d^2 f(x')}{dx'^2}$$

$$\int g(x) \delta'(x-x') dx = -g'(x')$$

$$\mathcal{L}(\eta, \eta_x, x, t)$$

$$\delta I = 0$$

$$\eta \rightarrow \eta + \delta\eta$$

$$\eta \rightarrow \eta + \epsilon \cdot \delta(x-x')$$

$$\delta I = \int \left( \frac{\partial \mathcal{L}}{\partial \eta} \delta\eta + \frac{\partial \mathcal{L}}{\partial \eta_x} \delta\eta_x \right) dx dt$$

$$\delta\eta = \epsilon \cdot \delta(x-x')$$

$$\delta I = \int \epsilon \left[ \frac{\partial \mathcal{L}}{\partial \eta} \delta(x-x') + \frac{\partial \mathcal{L}}{\partial \eta_x} \cdot \delta'(x-x') \right] dx dt$$

$$= \epsilon \cdot \left[ \frac{\partial \mathcal{L}}{\partial \eta(x')} - \frac{d}{dx'} \left( \frac{\partial \mathcal{L}}{\partial \eta_{x'}} \right) \right]$$

$$\frac{\delta I}{\epsilon} = \frac{\partial \mathcal{L}}{\partial \eta(x')} - \frac{d}{dx'} \left( \frac{\partial \mathcal{L}}{\partial \eta_{x'}} \right)$$

Example 1)  $\left[ \frac{\hbar^2}{2m} \cdot \nabla^2 + V \right] \Psi = i\hbar \cdot \frac{\partial \Psi}{\partial t}$  Time dependent Schrodinger equation.

This equation follows from the Lagrangian density.

$$\mathcal{L} = \frac{\hbar^2}{2m} \bar{\nabla} \Psi \cdot \bar{\nabla} \Psi^* + V \Psi^* \Psi$$

$$+ \frac{\hbar}{2i} (\Psi^* \dot{\Psi} - \Psi \dot{\Psi}^*)$$

→ current term.

Note, here we have  $\Psi$  &  $\Psi^*$

There are actually two degrees of freedom

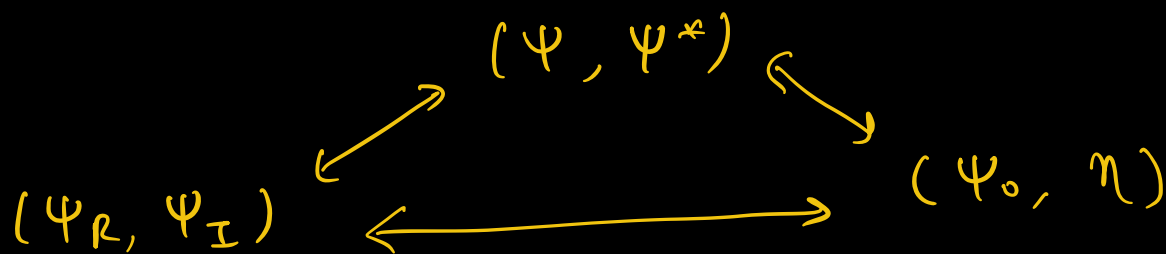
$$\Psi = \Psi_R + i\Psi_I$$

$\Psi_R, \Psi_I$  independent

or  $\psi(x) = \underbrace{\psi_0(x)}_{\text{There are two fields.}} e^{i\eta(x)}$

or we can also think of  $\psi$  &  $\psi^*$  as independent degree of freedom.

It has 2 degrees of freedom.



They convey same information.

$$\psi^* : \frac{d}{dx} \left( \frac{\partial \mathcal{L}}{\partial \psi_n^*} \right) + \frac{d}{dt} \left( \frac{\partial \mathcal{L}}{\partial \psi_t^*} \right) = \frac{\partial \mathcal{L}}{\partial \psi^*} \quad \left. \vphantom{\frac{d}{dx}} \right\} \begin{array}{l} \text{This gives} \\ \text{Time dependent} \\ \text{Schrodinger} \\ \text{equation.} \end{array}$$

$$\psi : \frac{d}{dx} \left( \frac{\partial \mathcal{L}}{\partial \psi_n} \right) + \frac{d}{dt} \left( \frac{\partial \mathcal{L}}{\partial \psi_t} \right) = \frac{\partial \mathcal{L}}{\partial \psi} \quad \left. \vphantom{\frac{d}{dx}} \right\} \begin{array}{l} \text{This gives} \\ \text{complex} \\ \text{conjugate of Time} \\ \text{dependent Schrodinger equation.} \end{array}$$

Example 2) Electromagnetic field Lagrangian  
 $\vec{E}, \vec{B}$

$$\vec{B} = \vec{\nabla} \times \vec{A} \quad , \quad \vec{E} = -\vec{\nabla} \phi - \frac{1}{c} \frac{\partial \vec{A}}{\partial t}$$

We construct a Tensor quantity  $F_{\mu\nu}$

$$F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu \quad (\text{Antisymmetric})$$

$$= \frac{\partial}{\partial x^\mu} A_\nu - \frac{\partial}{\partial x^\nu} A_\mu$$

$$\vec{A} = (\vec{A}, i\phi)$$

↑ four vector

$$\mathcal{L} = -\frac{1}{4c} F_{\mu\nu} F_{\mu\nu}$$

$$[F_{\mu\nu}]_{4 \times 4} = \begin{bmatrix} 0 & B_3 & -B_2 & -iE_1 \\ -B_3 & 0 & 0 & 0 \\ +B_2 & 0 & 0 & 0 \\ iE_1 & 0 & 0 & 0 \end{bmatrix}$$

Equation for  $\mu$  component of A four vector

$$\frac{d}{dx_\lambda} \left( \frac{\partial \mathcal{L}}{\partial \left( \frac{\partial A_\mu}{\partial x_\lambda} \right)} \right) = \frac{\partial \mathcal{L}}{\partial A_\mu}$$

→ This is zero here

$$\text{So, } \frac{d}{dx_\lambda} \left( \frac{\partial \mathcal{L}}{\partial \left( \frac{\partial A_\mu}{\partial x_\lambda} \right)} \right) = 0$$

⇒ gives Maxwell equation for zero current & zero charge case.

$$\text{we get } \partial_\mu F^{\mu\nu} = 0 \quad \checkmark$$



lec 4 (19<sup>th</sup> Jan, 2021 ; 3:30 pm to 5:00 pm)

## Symmetries of $\mathcal{L}$

symmetries of  $L(q, \dot{q}, t) \longrightarrow L + \frac{dF}{dt}(q, t) = L'$

This keeps  $L$ -eq<sup>n</sup> unchanged.

i.e.  $\frac{d}{dt} \left( \frac{\partial L'}{\partial \dot{q}} \right) = \frac{\partial L'}{\partial q}$  is satisfied.

$\frac{\partial L'}{\partial \dot{q}} \Rightarrow$  has an additional term

$$\begin{aligned} \frac{\partial}{\partial \dot{q}} \left( \frac{dF}{dt} \right) &= \frac{\partial}{\partial \dot{q}} \left( \frac{\partial F}{\partial t} + \frac{\partial F}{\partial q} \cdot \dot{q} \right) && \text{F not dependent explicitly on } \dot{q}. \\ &= \frac{\partial}{\partial \dot{q}} \left( \frac{\partial F}{\partial t} \right) + \frac{\partial}{\partial \dot{q}} \left( \frac{\partial F}{\partial q} \dot{q} \right) \\ &\quad \downarrow \qquad \qquad \downarrow \\ &\quad \text{This is independent of } q_i \end{aligned}$$

$$= 0 + \cancel{\frac{\partial}{\partial \dot{q}} \left( \frac{\partial F}{\partial q} \right)} \cdot \dot{q} + \frac{\partial F}{\partial q} \frac{\partial \dot{q}}{\partial \dot{q}}$$

$$= \frac{\partial F}{\partial q}$$

$\frac{d}{dt} \left( \frac{\partial F}{\partial q} \right) \Rightarrow$  extra term from LHS

we can show, that the RHS  $\frac{\partial L'}{\partial q}$  will also generate same term.

$$\frac{\partial L'}{\partial q} \Rightarrow \frac{\partial}{\partial q} \left( \frac{dF}{dt} \right)$$

It can be shown

$$\frac{d}{dt} \left( \frac{\partial F}{\partial \dot{q}} \right) = \frac{\partial}{\partial \dot{q}} \left( \frac{dF}{dt} \right)$$

when  $F$  is not dependent explicitly on  $\dot{q}$ .

Variational approach

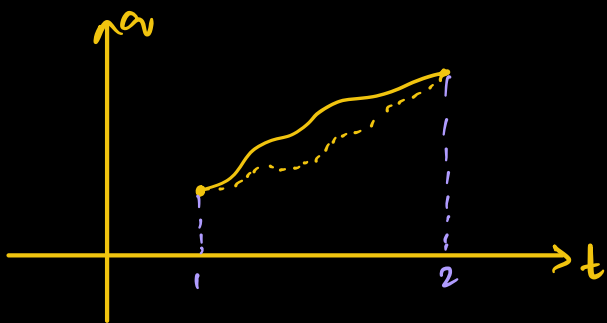
$$\begin{aligned} I' &= \int_{(1)}^{(2)} L' \cdot dt = \int_{(1)}^{(2)} L \cdot dt + \int_{(1)}^{(2)} \frac{dF}{dt} \cdot dt \\ &= \left[ \int_{(1)}^{(2)} L \cdot dt \right] + (F(2) - F(1)) \end{aligned}$$

$$\Rightarrow I' = I + (F(2) - F(1))$$

$\underbrace{\hspace{10em}}_{\text{This is extra term.}}$

when we do variation of  $I$ , we keep the end points fixed.

$$\text{so, } \delta(F(2) - F(1)) = 0$$



$\frac{dF}{dt}$  term is boundary term as far as the action is considered.

Energy Conservation

$$L = L(q, \dot{q})$$

not explicitly dependent on time "t".

Then something is conserved (usually called Energy)



$$\frac{dL}{dt} = \frac{\partial L}{\partial r} \dot{r} + \frac{\partial L}{\partial \dot{r}} \ddot{r} + \frac{\partial L}{\partial t}$$

$$= \frac{d}{dt} \left( \frac{\partial L}{\partial \dot{r}} \right) \dot{r} + \frac{\partial L}{\partial \dot{r}} \ddot{r} + \frac{\partial L}{\partial t}$$

→ used Lagrange's equation

$$\Rightarrow \frac{dL}{dt} = \frac{d}{dt} \left( \frac{\partial L}{\partial \dot{r}} \dot{r} \right) + \frac{\partial L}{\partial t}$$

$$\Rightarrow \frac{d}{dt} \left( \frac{\partial L}{\partial \dot{r}} \dot{r} - L \right) = - \frac{\partial L}{\partial t}$$

when  $L = L(r, \dot{r})$

so  $\frac{\partial L}{\partial t} = 0$

$$\Rightarrow \frac{d}{dt} \left( \frac{\partial L}{\partial \dot{r}} \dot{r} - L \right) = 0$$

→ This is called Hamiltonian.

$$H = \frac{\partial L}{\partial \dot{r}} \dot{r} - L$$

Analogy: for fields  $\eta_\rho(x_\mu)$   $\xrightarrow{x, y, z, t}$   
↖ vector component.

$$\sum_\nu \frac{d}{dx_\nu} \left[ \underbrace{\sum_\rho \left( \frac{\partial \mathcal{L}}{\partial \eta_{\rho, \nu}} \cdot \frac{\partial \eta_\rho}{\partial x_\mu} \right) - \mathcal{L}}_{T_{\mu\nu}} \cdot \delta_{\mu\nu} \right] = 0 \quad \text{if} \quad \frac{\partial \mathcal{L}}{\partial x_\mu} = 0$$

so;  $\frac{d}{dX_\nu} T_{\mu\nu} = 0$

ie; divergence of  $T_{\mu\nu}$  is zero.

$T_{\mu\nu} \Rightarrow$  Energy - Momentum Tensor.

ex|| from ---  $\eta(x)$   $p=1$  (only one component)

$x, t$  independent variables.

$$\sum_\nu \frac{d}{dX_\nu} T_{\mu\nu} = 0 \Rightarrow \frac{d}{dt} T_{\mu 0} + \frac{d}{dx} T_{\mu 1} = 0$$

$$\mu, \nu \in \{0, 1\}$$

$\uparrow$   $x$  - component.

"  $t$  ": Time component

Analogue of " $L \rightarrow L' = L + \frac{dF}{dt}$ "

$$\mathcal{L} \rightarrow \mathcal{L}' = \mathcal{L} + \frac{d}{dX_\nu} F_\nu (\{\eta_p\}, \{x_\mu\})$$

$F_\nu$  should not be function of derivatives of fields

ie;  $\frac{\partial F_\nu}{\partial \left( \frac{\partial \eta_p}{\partial X_\mu} \right)} = 0$

This is divergence term.

$F_\nu$  is a vector of same dimension as the space-time.

# Tensors

\* Rotations are Linear Transformations.

$$\vec{A}' = R \vec{A}$$

initial vector

rotated vector

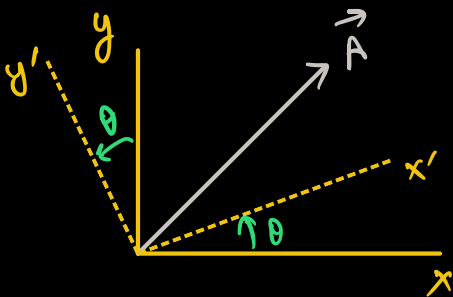
Rotation.

$$R_z(\theta) = \begin{pmatrix} \cos\theta & \sin\theta & 0 \\ -\sin\theta & \cos\theta & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Passive Rotation (most used one)

Active Rotation.

## Passive Rotation



vector does not move,  
only coordinate frame moves.

Object is not physically changed, only the coordinates are transformed.

$$\vec{A} = A_x \hat{i} + A_y \hat{j}$$

In new coordinate system

$$\vec{A} = A'_x \hat{i}' + A'_y \hat{j}'$$

$$\vec{A} = A_x \hat{i} + A_y \hat{j} = A'_x \hat{i}' + A'_y \hat{j}'$$

$$\hat{x} = \cos \theta \hat{x}' - \sin \theta \hat{y}'$$

$$\hat{y} = \sin \theta \hat{x}' + \cos \theta \hat{y}'$$

so;  $A_x \hat{x} + A_y \hat{y} = A_x (\cos \theta \hat{x}' - \sin \theta \hat{y}') + A_y (\sin \theta \hat{x}' + \cos \theta \hat{y}')$

$$= (A_x \cos \theta + A_y \sin \theta) \hat{x}' + (-A_x \sin \theta + A_y \cos \theta) \hat{y}'$$

so;  $A_{x'} = A_x \cos \theta + A_y \sin \theta$

$$A_{y'} = -A_x \sin \theta + A_y \cos \theta$$

$\Rightarrow \boxed{\begin{pmatrix} A_{x'} \\ A_{y'} \end{pmatrix} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} A_x \\ A_y \end{pmatrix}}$  This is how vector transforms

Consider the moment of Inertia matrix  $I_{ij}$ ,  
how it transforms?

$$R : I_{ij} \longmapsto I'_{ij} = R I_{ij} R^{-1}$$

where  $R$  is Rotation matrix

i.e;  $\boxed{I' = R I R^{-1}}$

This is how a matrix transform.

$$R^{-1} = R^T$$

for rotation matrices

# General Transformation rules for Tensors

$$\bar{A}^{ijk} = \sum_{m,n,p} \frac{\partial \bar{x}^i}{\partial x^m} \cdot \frac{\partial \bar{x}^j}{\partial x^n} \cdot \frac{\partial \bar{x}^k}{\partial x^p} \cdot A^{mnp}$$

if  $m, n, p \in \{1, 2\}$ , Then this has  $2^3 = 8$  terms.

$x^i \longrightarrow \bar{x}^i$  (coordinate frame transformation)

$$\bar{x}^i = \bar{x}^i(x^m)$$

Rank = 0 are scalars  $\bar{\phi}(\bar{x}) = \phi(x)$

Rank = 1 are vector  $\bar{A}^i = \frac{\partial \bar{x}^i}{\partial x^m} \cdot A^m$

we know that, under rotation vectors transform as  $\bar{A} = R A$

$$\text{ie; } \bar{A}^i = \sum_m R^i_m A^m$$

$$R = \left[ \frac{\partial \bar{x}^i}{\partial x^m} \right]$$

$$\bar{I}^{ij} = \frac{\partial \bar{x}^i}{\partial x^m} \cdot \frac{\partial \bar{x}^j}{\partial x^n} I^{mn}$$

For Rotation:  $\frac{\partial \bar{x}^i}{\partial x^j} = \frac{\partial x^j}{\partial \bar{x}^i}$

$\downarrow$   
R matrix

$\curvearrowright$

$R^{-1}$  (Inverse matrix)

$R_{ij} = \frac{\partial \bar{x}^i}{\partial x^j}$  takes unprimed to prime

$R^{-1}_{ji} = \frac{\partial x^j}{\partial \bar{x}^i}$  takes primed to unprimed.

$$(R_{ij})^T = R_{ji} = (R^{-1})_{ji}$$

so,  $\bar{I} = R I R^{-1}$  ✓

$$\begin{aligned}\bar{I}_{ij} &= R_{im} I_{mn} (R^{-1})_{nj} \\ &= \frac{\partial \bar{x}^i}{\partial x^m} I_{mn} \cdot \frac{\partial x^n}{\partial \bar{x}^j}\end{aligned}$$

but  $\frac{\partial x^n}{\partial \bar{x}^j} = \frac{\partial \bar{x}^j}{\partial x^n}$

so,  $\bar{I}_{ij} = \frac{\partial \bar{x}^i}{\partial x^m} \cdot \frac{\partial \bar{x}^j}{\partial x^n} I_{mn}$  ✓ Rotational transformation of moment of inertia is compatible to rank-2 tensor transformation property.

Contravariant

$$\bar{A}^{ijk\dots} = \frac{\partial \bar{x}^i}{\partial x^m} \cdot \frac{\partial \bar{x}^j}{\partial x^n} \cdot \frac{\partial \bar{x}^k}{\partial x^p} \dots A^{mnp\dots}$$

Covariant

$$\bar{A}_{ijk\dots} = \frac{\partial x^m}{\partial \bar{x}^i} \cdot \frac{\partial x^n}{\partial \bar{x}^j} \cdot \frac{\partial x^p}{\partial \bar{x}^k} \dots A_{mnp\dots}$$

Mixed Tensor:  $\bar{A}_i{}^{jk} = \frac{\partial x^m}{\partial \bar{x}^i} \cdot \frac{\partial \bar{x}^j}{\partial x^n} \cdot \frac{\partial \bar{x}^k}{\partial x^p} A_m{}^{np}$

## Example of Covariant vector

$$B_i = \frac{\partial \phi(x)}{\partial x^i} \quad \text{This transforms in covariant way.}$$

$$\bar{B}_i = \frac{\partial \bar{\phi}(\bar{x})}{\partial \bar{x}^i} \quad \begin{array}{l} \text{since } \phi(x) \text{ is scalar} \\ \text{so, } \phi(x) = \bar{\phi}(\bar{x}) \end{array}$$

$$\Rightarrow B_i = \frac{\partial \phi(x)}{\partial \bar{x}^i}$$

$$\begin{aligned} \text{if a vector } \bar{B}_i &= \frac{\partial \bar{\phi}(\bar{x})}{\partial \bar{x}^i} \cdot \frac{\partial x^m}{\partial \bar{x}^i} \\ &= \frac{\partial \phi(x)}{\partial x^m} \cdot \frac{\partial x^m}{\partial \bar{x}^i} \end{aligned}$$

$$\Rightarrow \underbrace{\bar{B}_i}_{\text{new } B} = \frac{\partial x^m}{\partial \bar{x}^i} \cdot \underbrace{B_m}_{\text{old } B} \quad \left. \vphantom{\frac{\partial x^m}{\partial \bar{x}^i} \cdot B_m} \right\} \begin{array}{l} \text{satisfies} \\ \text{covariant} \\ \text{transformation} \\ \text{rule.} \end{array}$$

upper index  $\Rightarrow$  Contravariant  
lower index  $\Rightarrow$  Covariant } Convention.

$B_{ij} = \frac{\partial A_i}{\partial x^j}$  does not follow transformation rule of rank 2 tensors.  
 $\hookrightarrow$  not tensor in general.



## Rotation Matrices

$$① R_{ij} = \frac{\partial \bar{x}^i}{\partial x^j} \equiv a_{ij}$$

$$R^{-1} = R^T \Rightarrow \frac{\partial \bar{x}^i}{\partial x^j} = \frac{\partial x^j}{\partial \bar{x}^i}$$

② Isotropic :  $A_{ij} = \lambda \cdot \delta_{ij}$  (does not change under rotation)  
under rotation

$$A \rightarrow A' = R A R^T$$

if  $A$  is isotropic

These types of matrices are called Isotropic.

$$\text{Then } A' = R (\lambda \cdot I) R^T = \lambda \cdot R R^T = \lambda \cdot I$$

$$\Rightarrow A' = \lambda I = A \Rightarrow A' = A$$

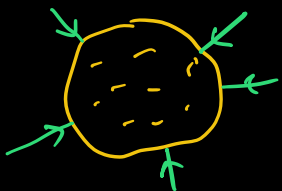
Stress :  $\sigma_{ij}$

$$\sigma_{ij} = -p \cdot \delta_{ij}$$

} Isotropic Stress

$p \Rightarrow$  pressure

pressure is isotropic. (normal to surface)



we know, that if  $A_{ijk} \dots$  is tensor

then

$$A'_{ijk} = \frac{\partial \bar{x}^i}{\partial x^m} \cdot \frac{\partial \bar{x}^j}{\partial x^n} \cdot \frac{\partial \bar{x}^k}{\partial x^p} \dots A_{mnp} \dots$$

③  $\epsilon_{ijk} \dots$

What is transformation property of

Levi-Civita?

It is actually pseudo-tensor



$\epsilon_{ijk} \dots$  does not transform exactly as tensor.  
 $\hookrightarrow$  fully antisymmetric w.r.t. any exchange of indices.

$$\epsilon_{ijk} = -\epsilon_{ikj} = -\epsilon_{kij}$$

Any pair we change,  
get us negative sign.

$$\epsilon_{ijk} \rightarrow ijk \rightarrow iljk \rightarrow lijk$$



so,  $\epsilon_{ijk} = -\epsilon_{lik}$

(a) cyclic change may give sign change if even # of indices

(b)  $\epsilon_{ijk} \dots$  all indices have to be different, otherwise it will be zero due to fully antisymmetric nature.

(c) Third Rank Levi-Civita  $\epsilon_{ijk}$

$\epsilon_{ijk}$  3! # of elements = 6 non-zero elements.

$\hookrightarrow$  27 elements. (out of which  $i, j, k \in \{1, 2, 3\}$  are zero)

declare  $\epsilon_{123} = 1$   $27 - 6 = 21$  are zero

Then, the other components are fixed.

i.e.  $\epsilon_{132} = -\epsilon_{123} = -1$ , etc.

$\epsilon_{112} = 0$ , etc.

• 3 of them is +1

• 3 of them is -1

$$(4) (A \times B)_i = \epsilon_{ijk} A_j B_k$$

$$(5) \text{ Determinant: } A = \begin{pmatrix} a_{11} & a_{12} & a_{13} \\ a_{21} & a_{22} & a_{23} \\ a_{31} & a_{32} & a_{33} \end{pmatrix}$$

$$|A| = a_{11}(a_{22}a_{33} - a_{23}a_{32}) - a_{12}(a_{23}a_{33} - a_{22}a_{31}) + a_{13}(a_{23}a_{32} - a_{22}a_{31})$$

can be written compactly in terms of Levi-Civita.

$$|A| = \epsilon_{ijk} a_{1i} a_{2j} a_{3k}$$

$$|A| = \epsilon_{123} a_{11} a_{22} a_{33} + \dots$$

for  $N \times N$  matrix  $A$

we will have  $N!$  terms in  $|A|$

$$|A| \equiv \det(A) = \epsilon_{i_1 i_2 \dots i_N} a_{1i_1} a_{2i_2} \dots a_{Ni_N}$$

$$(6) \epsilon_{ijk} \epsilon_{imn} \equiv A_{jkmn} = \delta_{jm} \delta_{kn} - \delta_{jn} \delta_{km}$$

(sum over  $i$ )

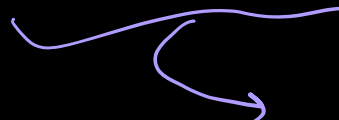
Einstein Notation: repeated indices are summed over.

Aside: Contraction reduces rank of a tensor (but it remains a tensor of lower rank)

ex  
Consider,  $A_{ij} B^j$  show that  $A'^i_j B'^j = \frac{\partial x^m}{\partial x'^i} A_{mn} B^n$

and hence we establish that  $A_{ij} B^j$  is rank 1 object

$$A'_{ij} B'^j = \frac{\partial x^m}{\partial \bar{x}^i} \cdot \frac{\partial x^m}{\partial \bar{x}^j} A_{mn} \cdot \frac{\partial \bar{x}^j}{\partial x^p} B^p$$
$$= \frac{\partial x^m}{\partial \bar{x}^i} \cdot \left( \frac{\partial x^m}{\partial \bar{x}^j} \cdot \frac{\partial \bar{x}^j}{\partial x^p} \right) \cdot A_{mn} \cdot B^p$$



$$\frac{\partial x^m}{\partial x^p} \quad \text{by chain rule}$$

$$= \frac{\partial x^m}{\partial \bar{x}^i} \cdot \frac{\partial x^m}{\partial x^p} \cdot A_{mn} \cdot B^p$$

$$= \frac{\partial x^m}{\partial \bar{x}^i} \cdot \delta_{mp} A_{mn} B^p$$

$$= \frac{\partial x^m}{\partial \bar{x}^i} \cdot (A_{mp} B^p)$$

Ex||

$$(A \times B) \times C = (A \cdot C) B - (B \cdot C) A$$

prove using Levi-Civita

$$[(A \times B) \times C]_i = \epsilon_{ijk} (A \times B)_j C_k$$

$$= \epsilon_{ijk} (\epsilon_{jmn} A_m B_n) C_k$$

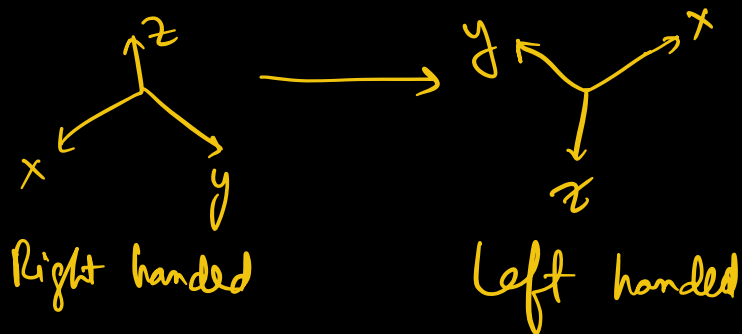
$$= \epsilon_{ijk} \epsilon_{jmn} A_m B_n C_k$$

pseudo Tensor

Recall pseudo-vector, example  $A = B \times C$   $B, C$  vectors

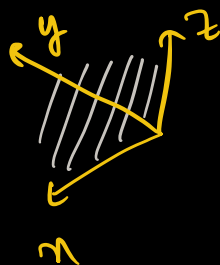
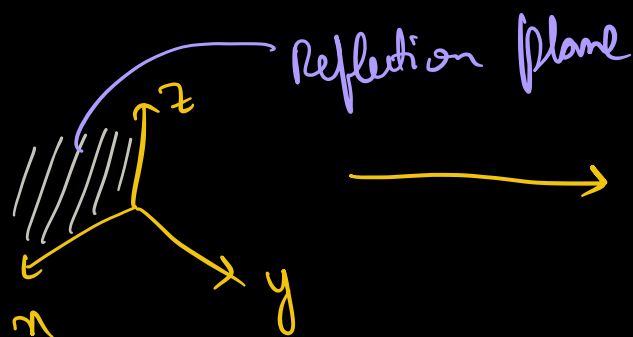
then  $A$  is pseudo vector or axial vector.

# Inversion / reflections.



$$\begin{aligned} x &\rightarrow -x \\ y &\rightarrow -y \\ z &\rightarrow -z \end{aligned}$$

Inversion  
or  
Parity  
Transformation



Reflection

Inversion :  $R = \begin{bmatrix} -1 & 0 & 0 \\ 0 & -1 & 0 \\ 0 & 0 & -1 \end{bmatrix}$

Reflection :  $R = \begin{bmatrix} 1 & 0 & 0 \\ 0 & -1 & 0 \\ 0 & 0 & 1 \end{bmatrix}$

Satisfies orthogonality condition.

$$R^T = R^{-1}$$

Improper Rotations.

$$R R^{-1} = I$$

$$R^{-1} = R^T \Rightarrow R R^T = I$$

$$\Rightarrow |R R^T| = |I| = 1$$

$$\Rightarrow |R| |R^T| = 1$$

$$\Rightarrow |R| |R| = 1 \Rightarrow |R|^2 = 1 \Rightarrow |R| = \pm 1$$

$$|R| = 1 \quad (\text{proper Rotation matrix})$$

$$|R| = -1 \quad (\text{improper Rotation matrix})$$

Given a vector  $A = B \times C$  , (B & C are vectors)  
under Improper Rotation , i.e;  $(x, y, z) \rightarrow (-x, -y, -z)$  ,

$$A = (-B) \times (-C) = B \times C$$

A remains same ,

Then A is called Axial Vector or Pseudo Vectors,

B & C are usual vectors,  
called polar vectors.

Change sign under inversion.

Now, we discuss more general element,  
Pseudo Tensor.

Transformation Rule.

let  $E_{ijk}$  be pseudo tensor.

It transforms as

$$\bar{E}_{ijk} = |a| \cdot \frac{\partial \bar{x}^i}{\partial x^m} \cdot \frac{\partial \bar{x}^j}{\partial x^n} \cdot \frac{\partial \bar{x}^k}{\partial x^p} \cdot E_{mnp}$$

→ because of this factor, it is not truly a tensor.

$|a| = \det \left( \frac{\partial \bar{x}^i}{\partial x^m} \right)$  = determinant or Jacobian  
of the transformation

$\bar{\epsilon}_{ijk} = \epsilon_{ijk}$  under above transformation rule.

(if we want to keep definition of Levi-Civita same in any frame, then we have to follow above transformation rule)


$$\epsilon_{123} = \epsilon_{231} = \epsilon_{321} = 1$$

$$\epsilon_{132} = \epsilon_{213} = \epsilon_{312} = -1$$

$$\bar{\epsilon}_{ijk} = |a| a_{im} a_{jn} a_{kp} \cdot \epsilon_{mnp}$$

$$\begin{aligned}\bar{\epsilon}_{123} &= |a| \cdot a_{1m} a_{2n} a_{3p} \epsilon_{mnp} \\ &= |a| |a| = |a|^2 = 1 = \epsilon_{123}\end{aligned}$$

$$\bar{\epsilon}_{213} = |a| \cdot a_{2m} a_{1n} a_{3p} \epsilon_{mnp}$$

  $m$  &  $n$  are dummy variable.

Just interchange them

$$= |a| a_{2m} a_{1n} a_{3p} \epsilon_{mnp}$$

$$= |a| a_{1m} a_{2n} a_{3p} \epsilon_{nmp}$$

$$= |a| a_{1m} a_{2n} a_{3p} (-\epsilon_{mnp})$$

$$= (-1) |a| |a| = (-1) |a|^2 = (-1) \cdot (1) = -1$$

$$= \epsilon_{213}$$

## pseudo vector

$A = B \times C$ , B & C polar, Then A is axial

B & C change sign under inversion

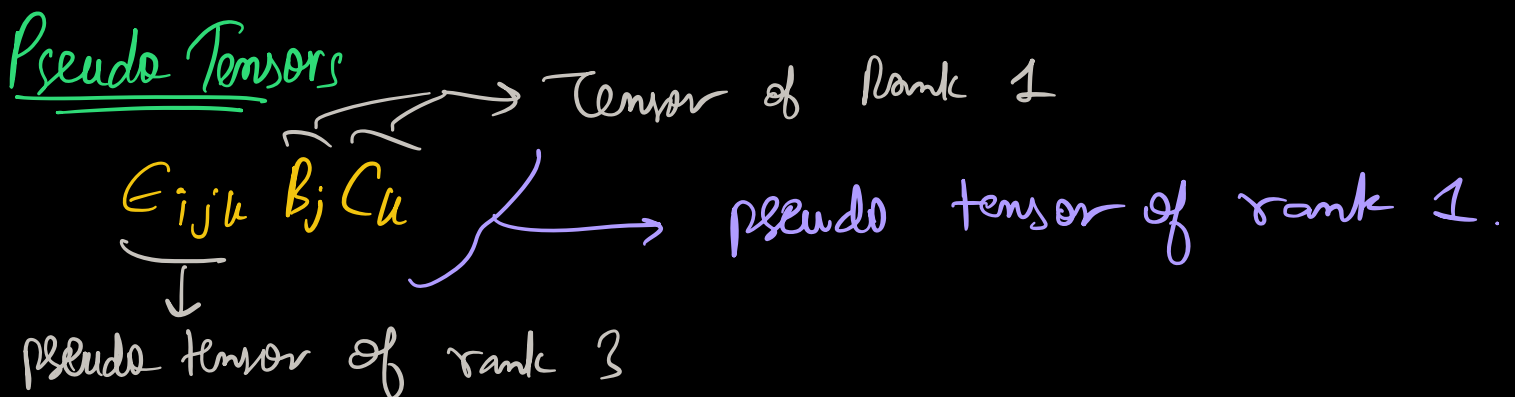
A don't change sign under inversion.

axial = polar  $\times$  polar

polar = polar  $\times$  axial

axial = axial  $\times$  axial

## Pseudo Tensors



Product of odd no. of pseudo tensor is pseudo tensor.  
" " even " " " " tensor.

- If odd # of components are pseudo, then product is pseudo.

pseudo thing is defined w.r.t. inversion.

Under proper rotation, pseudo tensor transforms as usual tensor.

Symmetric / Anti-symmetric Tensor defined w.r.t to pair of indices.  
 $A_{ijkl}$

consider a general matrix  $B_{ij}$

$B_{ij}$  can be written as

$$B_{ij} = A_{ij} + S_{ij}$$

where  $A_{ij}$  is antisymmetric  
 $S_{ij}$  " symmetric.

$$A_{ij} = \frac{B_{ij} - B_{ji}}{2}, \quad S_{ij} = \frac{B_{ij} + B_{ji}}{2}$$

$$\Rightarrow B_{ij} = \left( \frac{B_{ij} - B_{ji}}{2} \right) + \left( \frac{B_{ij} + B_{ji}}{2} \right)$$

note  $A_{ij} = -A_{ji}, \quad S_{ij} = S_{ji}$

show that under rotation, anti-symmetric part transforms to itself; same for symmetric.

$$\begin{array}{l} A_{ij} \rightarrow A'_{ij} \\ S_{ij} \rightarrow S'_{ij} \end{array} \quad \text{in general} \quad \begin{array}{l} A_{ij} \neq A'_{ij} \\ S_{ij} \neq S'_{ij} \end{array}$$

$$\begin{array}{l} \text{but } S'_{ij} = S'_{ji} \\ A'_{ij} = -A'_{ji} \end{array}$$



Example)

$$A_{ij} = \frac{\partial v_i}{\partial x_j} + \frac{\partial v_j}{\partial x_i} - \frac{2}{3} \cdot \delta_{ik} \cdot \left( \frac{\partial v_l}{\partial x_l} \right)$$

↪ sum over  $l$ ,  
so it is just  
 $\bar{D} \cdot \bar{V}$

(a)  $A_{ij}$  is symmetric

(b)  $\text{tr}(A) = 0$

$$\text{tr} A = A_{ii} = \frac{\partial v_i}{\partial x_i} + \frac{\partial v_i}{\partial x_i} - \frac{2}{3} \delta_{ii} (D \cdot V)$$

$$= D \cdot V + D \cdot V - \frac{2}{3} \cdot 3 (D \cdot V)$$

$$= 2 D \cdot V - 2 D \cdot V = 0$$



Lec 6 (29<sup>th</sup> Jan, 2021 ; 3:30 pm to 5:00 pm)

$$\mathcal{L} = -\frac{F_{\mu\nu}F^{\mu\nu}}{16\pi} + \frac{j_\alpha A_\alpha}{c} \quad \text{where} \quad F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$$

$$\text{and} \quad A_\mu = \begin{pmatrix} \vec{A} \\ i\phi \end{pmatrix}, \quad x_\mu = \begin{pmatrix} \vec{x} \\ ict \end{pmatrix}$$

$c \Rightarrow$  speed of light.

$$j_\mu = \begin{pmatrix} \vec{J} \\ ic\rho \end{pmatrix} \quad \vec{B} = \vec{\nabla} \times \vec{A}, \quad \vec{E} = -\vec{\nabla} \phi - \frac{1}{c} \cdot \frac{\partial \vec{A}}{\partial t}$$

for the limit  $j_\alpha = 0$ , i.e;  $\rho = 0$ ,  $\vec{J} = 0$

Maxwell equation becomes  $\vec{\nabla} \times \vec{B} = \frac{1}{c} \cdot \frac{\partial \vec{E}}{\partial t}$

$$\vec{\nabla} \times \vec{E} = -\frac{1}{c} \cdot \frac{\partial \vec{B}}{\partial t}$$

$$\vec{\nabla} \cdot \vec{B} = 0$$

$$\vec{\nabla} \cdot \vec{E} = 0$$

$$F_{\mu\nu} = \begin{bmatrix} 0 & B_3 & -B_2 & -iE_1 \\ -B_3 & 0 & B_1 & -iE_2 \\ B_2 & -B_1 & 0 & -iE_3 \\ iE_1 & iE_2 & iE_3 & 0 \end{bmatrix}$$

$$F_{13} = \partial_1 A_3 - \partial_3 A_1 = -(\vec{\nabla} \times \vec{A})_2 = -B_2$$

$$F_{34} = \frac{\partial A_4}{\partial x_3} - \frac{\partial A_3}{\partial x_4} = i \frac{\partial \phi}{\partial x_3} - \frac{\partial A_3}{ic \partial t} = i \left( \frac{\partial \phi}{\partial x_3} + \frac{\partial A_3}{c \partial t} \right)$$

$$= -iE_3$$

Euler Lagrange equation

$$\frac{d}{dx_\lambda} \left( \frac{\partial \mathcal{L}}{\partial (\partial_\lambda A_\alpha)} \right) = \frac{\partial \mathcal{L}}{\partial A_\alpha}$$

$$\mathcal{L} = \frac{-1}{16\pi} F_{\mu\nu} F_{\mu\nu}$$

$$\frac{d}{dx_\lambda} \left[ \frac{-1}{16\pi} \times 2 \times F_{\mu\nu} \cdot \frac{\partial F_{\mu\nu}}{\partial (\partial_\lambda A_\alpha)} \right] = 0$$

$$F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$$

$$\begin{aligned} \Rightarrow \frac{\partial F_{\mu\nu}}{\partial (\partial_\lambda A_\alpha)} &= \frac{\partial (\partial_\mu A_\nu)}{\partial (\partial_\lambda A_\alpha)} - \frac{\partial (\partial_\nu A_\mu)}{\partial (\partial_\lambda A_\alpha)} \\ &= \delta_{\mu\lambda} \delta_{\nu\alpha} - \delta_{\nu\lambda} \delta_{\mu\alpha} \end{aligned}$$

$$\Rightarrow \frac{d}{dx_\lambda} \left[ F_{\mu\nu} \cdot (\delta_{\mu\lambda} \delta_{\nu\alpha} - \delta_{\nu\lambda} \delta_{\mu\alpha}) \right] = 0$$

$$\Rightarrow \frac{d}{dx_\lambda} [F_{\lambda\alpha} - F_{\alpha\lambda}] = 0 \Rightarrow \frac{d}{dx_\lambda} [F_{\lambda\alpha} + F_{\lambda\alpha}] = 0$$

$$\Rightarrow 2 \cdot \frac{d}{dx_\lambda} F_{\lambda\alpha} = 0$$

$$\Rightarrow \boxed{\frac{d}{dx_\lambda} F_{\lambda\alpha} = 0}$$

$$\boxed{\frac{d}{dx_\lambda} (F_{\lambda\alpha}) = 0}$$

Euler Lagrange Equation for E.M. field.

$$\text{for } \alpha = 3, \quad \frac{dF_{13}}{dx_1} + \frac{dF_{23}}{dx_2} + \frac{dF_{33}}{dx_3} + \frac{dF_{43}}{dx_4} = 0$$

$$\Rightarrow \frac{d}{dx} F_{13} + \frac{dF_{23}}{dy} + \frac{dF_{43}}{ic dt} = 0$$

Third column elements of  $F$

$$\begin{matrix} -B_2 \\ B_1 \\ 0 \\ iE_3 \end{matrix}$$

$$\Rightarrow -\frac{d}{dx_1} B_2 + \frac{d}{dx_2} B_1 + \frac{i}{ic} \frac{dE_3}{dt} = 0$$

$$\Rightarrow -\left[ \frac{dB_2}{dx_1} - \frac{dB_1}{dx_2} \right] + \frac{1}{c} \frac{dE_3}{dt} = 0$$

$$\Rightarrow -(\nabla \times B)_3 + \frac{1}{c} \frac{\partial E_3}{\partial t} = 0$$

for  $\alpha=4$ ,  $\bar{\nabla} \cdot \bar{E} = 0$ , for  $\alpha=1,2,3$

$$\Rightarrow \bar{\nabla} \times \bar{B} = \frac{1}{c} \frac{\partial \bar{E}}{\partial t}$$

$\bar{\nabla} \cdot \bar{B} = 0$

$$\bar{\nabla} \times \bar{E} + \frac{1}{c} \frac{\partial \bar{B}}{\partial t} = 0 \quad \left\{ \begin{array}{l} \text{follows from the identity} \\ \frac{dF_{\mu\nu}}{dx_\alpha} + \frac{dF_{\alpha\mu}}{dx_\nu} + \frac{dF_{\nu\alpha}}{dx_\mu} = 0 \end{array} \right.$$

$F_{\mu\nu}$  is invariant under

$$A_\mu \longrightarrow A_\mu + \partial_\mu \Lambda(\vec{x}, t)$$

Gauge Symmetry. Gauge invariance of Maxwell's Equations.

i.e.,

$$\begin{aligned} \vec{A} &\longrightarrow \vec{A} + \vec{\nabla} \Lambda \\ i\phi &\longrightarrow i\phi + \frac{1}{ic} \frac{\partial \Lambda}{\partial t} \end{aligned}$$

Maxwell equations don't change under these transformations

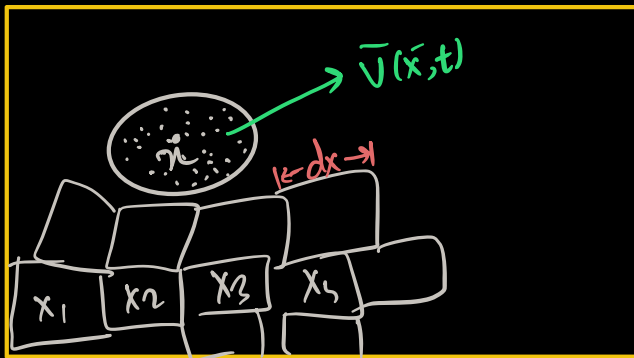
$$F_{\mu\nu} \rightarrow F'_{\mu\nu} = (\partial_\mu A_\nu' - \partial_\nu A_\mu')$$

$$A_\mu' = A_\mu + \partial_\mu \Lambda$$

$$\begin{aligned} \Rightarrow F'_{\mu\nu} &= F_{\mu\nu} + \partial_\mu (\partial_\nu \Lambda) - \partial_\nu (\partial_\mu \Lambda) \\ &= F_{\mu\nu} + [\partial_\mu, \partial_\nu] \Lambda = F_{\mu\nu} \end{aligned}$$

$$\Rightarrow \boxed{F'_{\mu\nu} = F_{\mu\nu}}$$

## Fluid Mechanics



we study a small blob of fluid.  
(still have macroscopic no. of particles inside the blob)

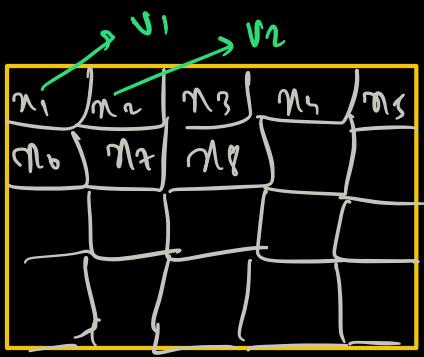
$\bar{v}$  = average velocity of all the particles in the small volume element (the blob)

The whole blob  has only one point,  $x$ .

The blobs (or the boxes) have the size  $dx$ .

These blobs are mesoscopic object. (not microscopic nor macroscopic)

$\bar{v} \Rightarrow$  mesoscopic variable (coarse grained over space)



$$\bar{v}(\bar{x}, t)$$

want to write equation which describes this mesoscopic variable, velocity.

Navier-Stokes equation

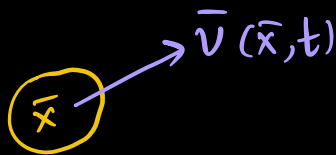
$$\rho \left[ \frac{\partial \bar{v}(\bar{x}, t)}{\partial t} + (\bar{v} \cdot \bar{\nabla}) \bar{v} \right] = -\bar{\nabla} p + \bar{f}_{\text{ext}} + \eta \nabla^2 \bar{v}(\bar{x}, t)$$

$\bar{f}_{\text{ext}} \Rightarrow$  External force density  
= force on unit volume

$\eta \Rightarrow$  viscosity.

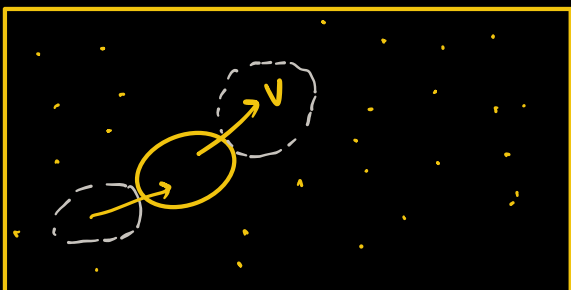
viscous dissipation term

$$\rho \cdot \frac{D}{Dt} \bar{v} = \text{all forces}$$



$\frac{D \bar{v}}{Dt} =$  acceleration of the blob.

$\frac{\partial \bar{v}(\bar{x}, t)}{\partial t} \}$  at fixed  $\bar{x}$



$$\frac{\partial \bar{v}(\bar{x}, t)}{\partial t} = \lim_{\Delta t \rightarrow 0} \frac{v(x, t + \Delta t) - v(x, t)}{\Delta t}$$

When the sensor for measuring velocity is located at some fixed point  $\bar{x}$

at time  $t$  &  $t + \Delta t$ , it measures velocity of different blobs.

but when we follow the blob we get the term  $(\bar{v} \cdot \bar{\nabla}) \bar{v}$



$$\frac{D\bar{v}}{Dt}$$

this is w.r.t. lab frame.

Lagrangian Velocity

Euler Velocity

let us look at any quantity  $Q$  (could be scalar or vector)

$$\begin{aligned} \frac{DQ(\bar{x}, t)}{Dt} &= \lim_{\Delta t \rightarrow 0} \frac{Q(\bar{x} + \bar{v}\Delta t, t + \Delta t) - Q(\bar{x}, t)}{\Delta t} \\ &= \frac{Q(\bar{x}, t) + \frac{\partial Q}{\partial \bar{x}} \cdot \bar{v} \Delta t + \frac{\partial Q}{\partial t} \Delta t - Q(\bar{x}, t)}{\Delta t} \end{aligned}$$

$$\frac{\partial Q}{\partial \bar{x}} = \nabla Q$$

$$\Rightarrow \frac{DQ}{Dt} = (\bar{\nabla} Q) \cdot \bar{v} + \frac{\partial Q}{\partial t}$$

$$\Rightarrow \boxed{\frac{DQ}{Dt} = \frac{\partial Q}{\partial t} + (\bar{v} \cdot \bar{\nabla}) Q}$$

Material Derivative

Eulerian Derivative.

if  $\bar{Q}$  is vector

$$\frac{D\bar{Q}}{Dt} = \frac{\partial \bar{Q}}{\partial t} + (\bar{v} \cdot \bar{\nabla}) \bar{Q}$$

sitting at fixed point  
& measuring  $\bar{Q}$

following the fluid  
& measuring  $\bar{Q}$ .

practically, we measure  $\frac{\partial Q}{\partial t}$  (just the Eulerian derivative w.r.t. time)

$$\text{If } \bar{Q} = \bar{v}(\bar{x}, t)$$

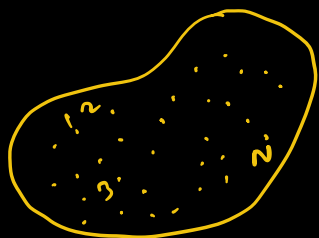
$$\text{Then } \frac{D\bar{v}(\bar{x}, t)}{Dt} = \frac{\partial \bar{v}(\bar{x}, t)}{\partial t} + (\bar{v}(\bar{x}, t) \cdot \bar{\nabla}) \bar{v}(\bar{x}, t)$$

$\rho \cdot \frac{D\bar{v}}{Dt}$  is rate of change of momentum. (also for compressible fluid)

(don't think that  $\rho$  should go inside derivative,  $\frac{D}{Dt}(\rho \bar{v})$  is not needed)



when the blob has particle constituent of same mass.



Blob with  $N$   
particles of  
mass  $m$

$$m \frac{dv_1}{dt} = \dots$$

$$m \cdot \frac{dv_2}{dt} = \dots$$

$\vdots$

$$m \cdot \frac{dv_N}{dt} = \dots$$

$$\bar{v} = \frac{1}{N} \sum_{n=1}^N \bar{v}_n$$

---


$$m \sum_j \frac{d\bar{v}_j}{dt} = \text{Total force.}$$

$$\Rightarrow N m \cdot \frac{d}{dt} \left( \frac{1}{N} \sum_j \bar{v}_j \right) = \text{Total force.}$$

$$\Rightarrow (Nm) \cdot \frac{d}{dt} \bar{v} = \text{Total force}$$

$Nm \Rightarrow$  total mass of blob.

divide by volume  $V$

$$\frac{(Nm)}{V} \cdot \frac{d}{dt} \bar{v} = \frac{\text{force}}{V} = \text{force density}$$

Note  $\rho = \frac{Nm}{V}$

$$\Rightarrow \rho \cdot \frac{d}{dt} \bar{v} = \text{force density.}$$

so; although  $f$  is function of  $\bar{x}$  &  $t$ ,  $f = f(\bar{x}, t)$ ,  
it does not go inside  $\frac{D}{Dt}$ .

Coarse grained : observing with lower resolution.



$$\rho \cdot \left[ \frac{\partial \bar{v}}{\partial t} + (\bar{v} \cdot \bar{\nabla}) \bar{v} \right] = -\bar{\nabla} p + \eta \nabla^2 \bar{v} + \bar{f}_{\text{ext}}(\bar{x})$$

$\frac{D \bar{v}}{D t}$  : material derivative or Time derivative in Lagrangian frame.

$\frac{\partial \bar{v}}{\partial t}$  : Eulerian derivative

$\frac{D}{D t}$  : hard to calculate; need to follow the blob.

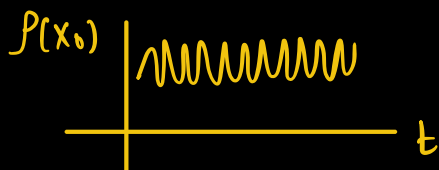
①  $\rho = \rho_0$  : constant : incompressible

ex) Water is practically incompressible

$\rho(\bar{x})$  density oscillates at sound speed.

$$f \cdot \lambda = c$$

$c$  = sound speed.

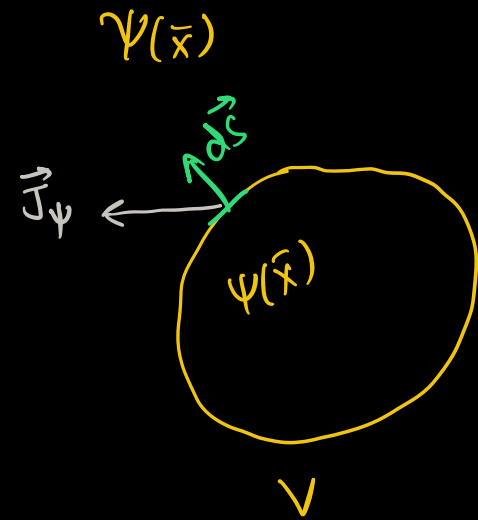


$t_v \Rightarrow$  velocity time scale

if  $f \gg (t_v)^{-1} \Rightarrow \rho$  is practically constant.

Mach number  $= \frac{\bar{v}}{c} \ll 1 \Rightarrow$  density fluctuation is fast as compared to velocity time scale.  
 $\Rightarrow$  and hence practically constant.  
 $c \Rightarrow$  property of material.

② Continuity Equation. for property  $\Psi$ .  $\Psi$  could be mass, energy, entropy, etc.



Content of the property inside volume  $V$  is  $\int \Psi(x) dV$

$$\frac{\partial}{\partial t} \int \Psi(x) dV$$

$\vec{J}_\Psi \Rightarrow$  flux of  $\Psi$

$$= - \oint \vec{J}_\Psi \cdot d\vec{S} + \int g_\Psi dV$$

$\downarrow$   
 rate of generation or destruction of  $\Psi$  inside volume

$$\frac{\partial}{\partial t} \int \Psi(x) dV = - \int \vec{\nabla} \cdot \vec{J}_\Psi dV + \int g_\Psi dV$$

$$\Rightarrow \int_V dV \left[ \frac{\partial \Psi}{\partial t} + \vec{\nabla} \cdot \vec{J}_\Psi - g_\Psi \right] = 0 \quad \text{for any arbitrary volume.}$$

for  $V \rightarrow 0$  (volume)  $\Rightarrow$   $\boxed{\frac{\partial \psi}{\partial t} + \vec{\nabla} \cdot \vec{J}_\psi - g_\psi = 0}$  Continuity Equation

for  $g_\psi = 0 \Rightarrow \boxed{\frac{\partial \psi}{\partial t} = -\vec{\nabla} \cdot \vec{J}_\psi}$

$\boxed{\vec{J}_\psi = \psi \vec{v}}$

amount  
of  $\psi$  leaving

velocity with which it  
is leaving

Continuity Equation : a conservation law if  $g_\psi = 0$

$$\frac{\partial \psi}{\partial t} = -\vec{\nabla} \cdot (\psi \vec{v})$$

If  $\psi = \rho(x, t)$  density

Then  $\boxed{\frac{\partial \rho}{\partial t} = -\vec{\nabla} \cdot (\rho \vec{v})}$

If density  $\rho = \rho_0 \Rightarrow 0 = -\vec{\nabla} \cdot (\rho \vec{v})$   
 $= -\rho \vec{\nabla} \cdot \vec{v} - \vec{v} \cdot (\vec{\nabla} \rho)$

$\Rightarrow \boxed{\vec{\nabla} \cdot \vec{v} = 0}$

$$\text{Incompressibility} \Rightarrow \vec{\nabla} \cdot \vec{v} = 0$$


---

$$\text{Recall } \rho \left[ \frac{\partial \vec{v}}{\partial t} + (\vec{v} \cdot \vec{\nabla}) \vec{v} \right] = -\vec{\nabla} p + \eta \nabla^2 \vec{v} + \vec{f}_{\text{ext}}$$

non linear term

$$\text{let } \vec{f}_{\text{ext}} = 0$$

If  $\rho$  is incompressible, then  $p$  determined by  $\vec{v}$ , but not converse.

( $\Rightarrow$  these types of quantities are called Passive Quantities;

i.e., it does not decide  $\vec{v}$ , but get decided by  $\vec{v}$ )

$$\text{Reynolds Number} = \frac{\rho (\vec{v} \cdot \vec{\nabla}) \vec{v}}{\eta \nabla^2 \vec{v}}$$

$$\approx \frac{\left( \frac{\rho \cdot v_0^2}{L} \right)}{\left( \frac{\eta v_0}{L^2} \right)} = \frac{\rho v_0 L}{\eta}$$

$L$  is the scale of the flow

$$\rho (\vec{v} \cdot \vec{\nabla}) \vec{v} \approx \rho \cdot v_0 \cdot \frac{1}{L} \cdot v_0$$

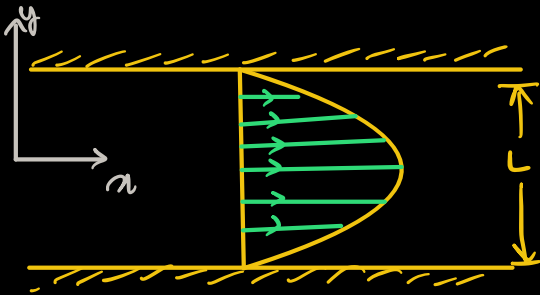
$$\eta \cdot \nabla^2 \vec{v} \approx \eta \cdot \frac{1}{L} \cdot \frac{1}{L} \cdot v_0$$

$\eta \Rightarrow$  dynamic viscosity,  $\nu \Rightarrow$  kinematic viscosity

$$Re = \left(\frac{\rho}{\eta}\right) \cdot v_0 L = \frac{v_0 L}{\nu}$$

Reynolds Number

$$\nu = \frac{\eta}{\rho} \quad \text{Kinematic viscosity}$$



Parabolic flow profile

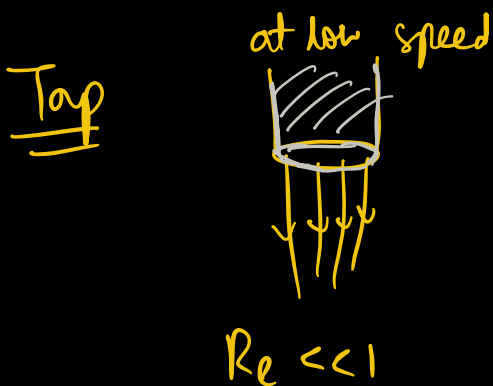
$$v_0 \sim \frac{1}{L} \cdot \int_0^L v_x \cdot dy$$

L is the width over which v is varying.

If  $Re \ll 1 \Rightarrow$  Laminar flow. (smooth),  
ignore the non-linear term.

ex : Chaotic or non-laminar flow

If  $Re \gg 1 \Rightarrow$  Turbulant flow (chaotic)



$$\rho \cdot \frac{\partial v}{\partial t} + \dots = \dots \quad \text{so:} \quad \left[ \rho \cdot \frac{\partial v}{\partial t} \right] = [\eta \nabla^2 v] \quad \text{dimensional analysis}$$

$$\Rightarrow [\eta] = \frac{[\rho]}{[t]} [L]^2 = \frac{M}{L^3} \cdot \frac{L^2}{T} = \frac{M}{LT}$$

$$\Rightarrow [\eta] = M L^{-1} T^{-1}$$

and  $[\nu] = L^2 T^{-1}$  Kinematic Viscosity.

$$\eta_{\text{water}} = 10^{-3} \frac{\text{kg}}{\text{m} \cdot \text{s}} = 10^{-3} \text{ Pa} \cdot \text{s} = 0.01 \frac{\text{gm}}{\text{cm} \cdot \text{sec}}$$

$$\eta_{\text{honey}} = 850 \cdot \eta_{\text{water}}$$

$$\eta_{\text{mercury}} = 1.5 \cdot \eta_{\text{water}}$$

$$\text{Pascal} = \frac{\text{N}}{\text{m}^2}$$

N  $\Rightarrow$  Newton

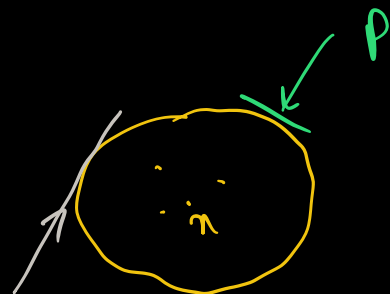
Astrophysical  $\Rightarrow v_0$  high  $\Rightarrow Re$  high.

Ocean  $\Rightarrow L$  high  $\Rightarrow Re$  high.

Blood flow is also Turbulent.

Terms on the RHS of N.S. equation.

$\eta \nabla^2 \bar{v}$  (a kind of force) : Viscous term.



pressure acts perpendicular to surface.

forces on the blob are due to neighbours.

There are other forces which work tangentially to blob (viscous force/shear force)  $\eta \nabla^2 v$

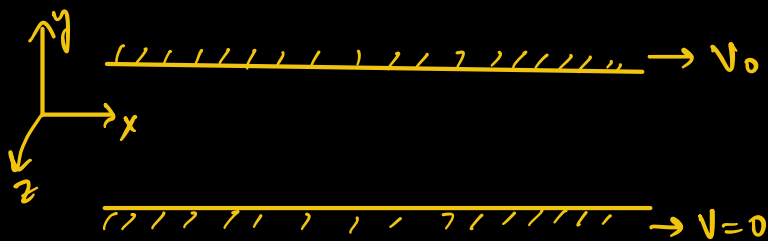


When fluid is compressible, there are other forces than pressure which act perpendicular to surface.

for incompressible fluid, pressure is the only force  $\perp$  to surface

$\Rightarrow$  & viscous term is tangential to surface.

### Shear flow



consider flow between two plates.

- upper plate moving.
- lower plate at rest.

find profile of the flow.

Can think of  $\infty$   $x$ - $z$  plane. For simplicity being suppress  $z$  direction.

### Assumptions


- 1)  $\infty$  plate in  $x$  direction  $\Rightarrow v$  is not function of  $x$
- 2) steady flow (pattern of flow does not change with time)

$$\frac{\partial v}{\partial t} = 0 \quad \text{at a given } \bar{x}, \text{ velocity is not changing.}$$

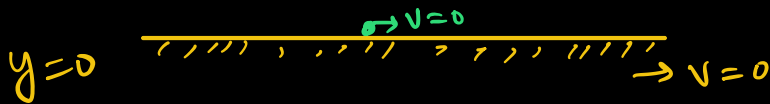
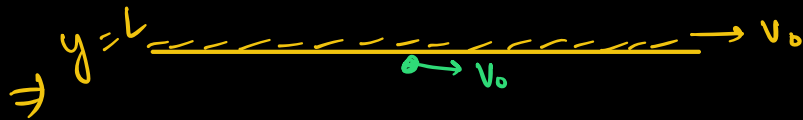
so,  $\vec{v}(x, y)$  Ignore  $z$ , can be uniform along that also.

$$\vec{v}(x, y) = [v_x(y), v_y(y)] \quad \text{because of (1), i.e., Translational invariance.}$$

### 3) No slip Boundary Condition.

 fluid which is in contact with wall, is at rest w.r.t. wall.

fluid element / blob which is in contact with the wall is always at rest w.r.t wall.



### 4) Incompressible fluid

$$\Rightarrow \bar{\nabla} \cdot \bar{v} = 0 \Rightarrow \frac{\partial v_x}{\partial x} + \frac{\partial v_y}{\partial y} = 0$$

$$v_x = v_x(y) \Rightarrow \frac{\partial v_x}{\partial x} = 0 \Rightarrow \boxed{\frac{\partial v_y}{\partial y} = 0}$$

$$\Rightarrow \boxed{v_y = \text{constant}}$$

$$\Rightarrow v_y = v_y(y=0) = v_y(y=L) = 0$$

$$\Rightarrow \boxed{v_y = 0} \quad v_y \text{ is zero at the two walls.}$$

$$\Rightarrow \boxed{\bar{v} = v_x(y) \hat{x}}$$

s) Non turbulent / Laminar / slow flow

$$\Rightarrow (\bar{v} \cdot \bar{\nabla}) \cdot \bar{v} \approx 0$$

but here, we can see  $(\bar{v} \cdot \bar{\nabla}) \cdot \bar{v} = 0$  exactly zero

$$\left( v_x \frac{\partial}{\partial x} + v_y \frac{\partial}{\partial y} \right) v_x = v_x \cancel{\frac{\partial v_x}{\partial x}} + v_y \cancel{\frac{\partial v_x}{\partial y}} = 0$$

$$\text{since } v_y = 0 \Rightarrow \left( v_x \frac{\partial}{\partial x} + v_y \frac{\partial}{\partial y} \right) v_y = 0$$

Here, the non-linearly term is anyway exactly zero.

So, we can ignore this non-linear term from beginning.

Now we solve the N-V equation.

$$\left[ \frac{\partial \bar{v}}{\partial t} + (\bar{v} \cdot \bar{\nabla}) \bar{v} \right] = -\frac{\bar{\nabla} p}{\rho_0} + \frac{\eta}{\rho} \bar{\nabla}^2 \bar{v} + \cancel{f_{ext}}$$

$\eta$  equation:

$$\frac{\partial \bar{v}}{\partial t} = 0 \text{ because of steady flow.}$$

$$(\bar{v} \cdot \bar{\nabla}) \bar{v} = 0 \text{ anyway}$$

$$\Rightarrow \boxed{\frac{\partial p}{\partial x} = \eta \cdot \frac{\partial^2 v_x(y)}{\partial y^2}}$$

$$\begin{aligned} \nabla^2 v_x &= \cancel{\frac{\partial^2 v_x}{\partial x^2}} + \frac{\partial^2 v_x}{\partial y^2} \\ &= \frac{\partial^2 v_x}{\partial y^2} \end{aligned}$$

$p$  can only be function of  $y$

so;  $\frac{\partial p}{\partial x} = 0$

$$\Rightarrow \boxed{\frac{\partial^2 v_x(y)}{\partial y^2} = 0}$$

$$\Rightarrow v_x = Ay + B$$

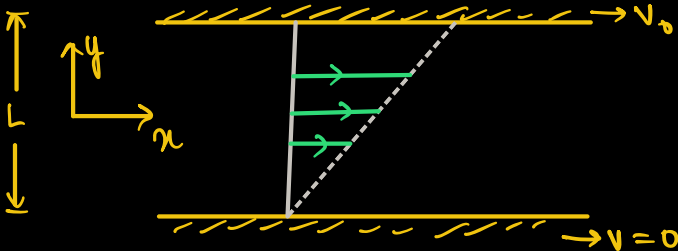
$\xrightarrow{v_0}$   
 $\xrightarrow{0}$

B.C. :  $v_x(0) = 0$  ,  $v_x(L) = v_0$

y-equation

$$\frac{\partial p}{\partial y} = \cancel{\eta} \left( \cancel{\gamma} \right) \cancel{v_y} \rightarrow 0$$

$$\Rightarrow \frac{\partial p}{\partial y} = 0 \Rightarrow \boxed{p = \text{constant}}$$



$$\Rightarrow \boxed{v_x(y) = \frac{v_0}{L} y}$$

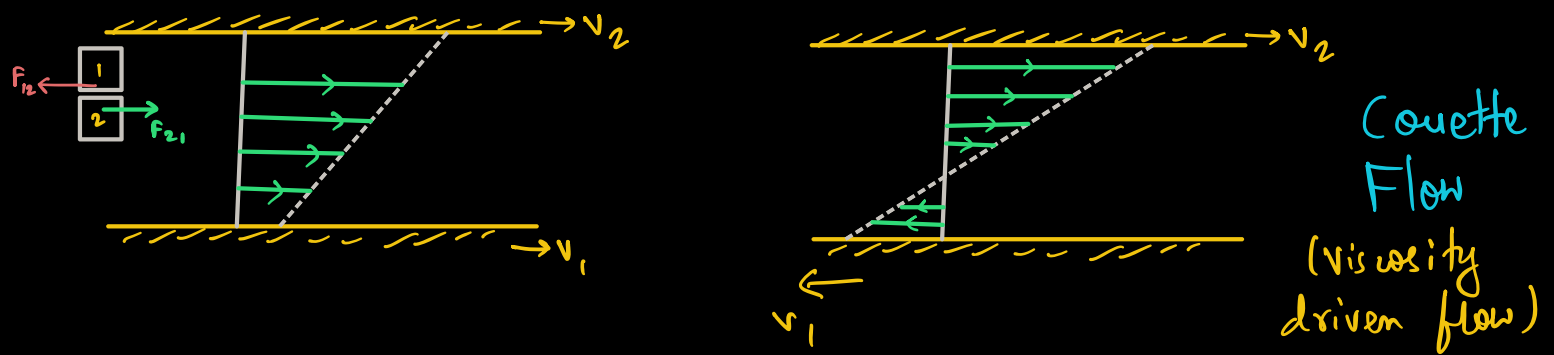
linear profile.

$p = \text{constant}$ .

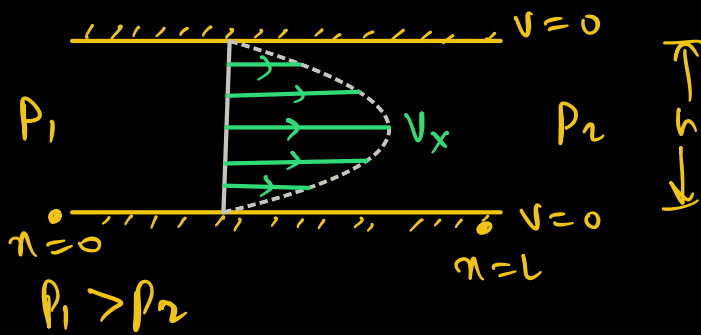
So what driving the flow? Tangential force.  
so called Shear flow.



Lee 9 (12<sup>th</sup> Feb, 2021 ; 3:30 pm to 5:00 pm)



## Poiseuille Flow (pressure driven flow)



### Assumption

- ①  $\frac{\partial \mathbf{v}}{\partial t} = 0$ ,  $(\nabla \cdot \bar{\mathbf{p}})\bar{\mathbf{v}} = 0$ ,  $\rho = \rho_0$  (slow flow)
- ② No slip boundary condition at the walls.

pressure difference breaks the invariance in  $x$  direction.

$P(x)$

$V_x(y)$

$V_y(y)$

For now,  
just ask  
can  $x$  independent  
solution exist?

Since it is incompressible fluid; so even the invariance in  $x$ -direction is broken, we get velocities to be independent of  $x$ .

for  $L \gg h$

, what happens at boundary  $y=0$  or  $y=h$   
will not affect much in the middle.

at  $x=0$ ; pressure is same

$$\text{so; } v_y(x=0, y) = 0$$

so; pressure can't be function of  $y$

$$\frac{\partial v_x}{\partial x} + \frac{\partial v_y}{\partial y} = 0 \quad \text{if } v_x = v_x(y)$$
$$\Rightarrow \frac{\partial v_y}{\partial y} = 0$$

$$\text{so; } v_y = \text{constant}$$

$$\text{at; } h=0 \text{ \& } h=L, \quad v_y = 0$$

$$\Rightarrow \boxed{v_y = 0}$$

So, we look steady state solution  $v_x = v_x(y)$

N-V. equation → steady state ↗  $Re \ll 1$ , slow flow

$$\frac{\partial \bar{v}}{\partial t} + (\bar{v} \cdot \nabla) \cdot \bar{v} = -\bar{\nabla} p + \eta \nabla^2 \bar{v}$$

$$(v_x \cdot \partial_x v_x) = 0 \quad \text{exactly here.} \quad (\text{no need to truly worry about low or high } Re)$$

$$\bar{\nabla} p = \eta \nabla^2 \bar{v}$$

$$x: \quad \frac{\partial p}{\partial x} = \eta \cdot \frac{\partial^2 V_x}{\partial y^2}$$

$$y: \quad \frac{\partial p}{\partial y} = 0 \Rightarrow p = p(x) \quad p \text{ is function of } x \text{ only}$$

$$\underbrace{\frac{\partial p(x)}{\partial x}}_{\text{function of } x} = \underbrace{\eta \cdot \frac{\partial^2 V_x(y)}{\partial y^2}}_{\text{function of } y}$$

LHS : function of  $x$  }  $\Rightarrow$  They have to be constant.  
 RHS : " "  $y$

$$\Rightarrow \frac{\partial p}{\partial x} = \text{constant}$$

$$\text{so; } \boxed{p = p_2 + \frac{x}{L} \cdot (p_2 - p_1)}$$

$$\Rightarrow \frac{\partial p}{\partial x} = \frac{p_2 - p_1}{L}$$

$$\text{so; } \eta \cdot \frac{\partial^2 V_x(y)}{\partial y^2} = \frac{p_2 - p_1}{L}$$

$$\Rightarrow \frac{\partial^2 V_x(y)}{\partial y^2} = \frac{1}{\eta} \cdot \left( \frac{p_2 - p_1}{L} \right)$$

$$\Rightarrow \boxed{V_x(y) = \frac{1}{2\eta} \cdot \left( \frac{P_2 - P_1}{L} \right) \cdot y^2 + Ay + B}$$

$$V_x(y=0) = 0 \Rightarrow B = 0$$

$$V_x(y=h) = 0 \Rightarrow \text{gives us } A$$

$$\frac{1}{2\eta} \cdot \frac{(P_2 - P_1)}{L} \cdot h^2 + Ah = 0$$

$$\Rightarrow A = -\frac{1}{2\eta} \cdot \frac{h}{L} \cdot (P_2 - P_1)$$

$$\boxed{V_x(y) = \frac{1}{2\eta} \cdot \left( \frac{P_2 - P_1}{L} \right) \cdot (y^2 - hy)}$$

$$P_1 - P_2 = \Delta P > 0$$

$$\boxed{V_x(y) = -\frac{\Delta P}{2\eta L} \cdot [y^2 - hy]}$$

This flow is driven by pressure difference.

$$V_x = \frac{-1}{2\eta} \cdot \left( \frac{\Delta P}{L} \right) \cdot (y^2 - hy)$$

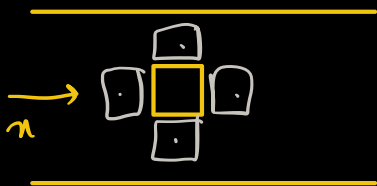
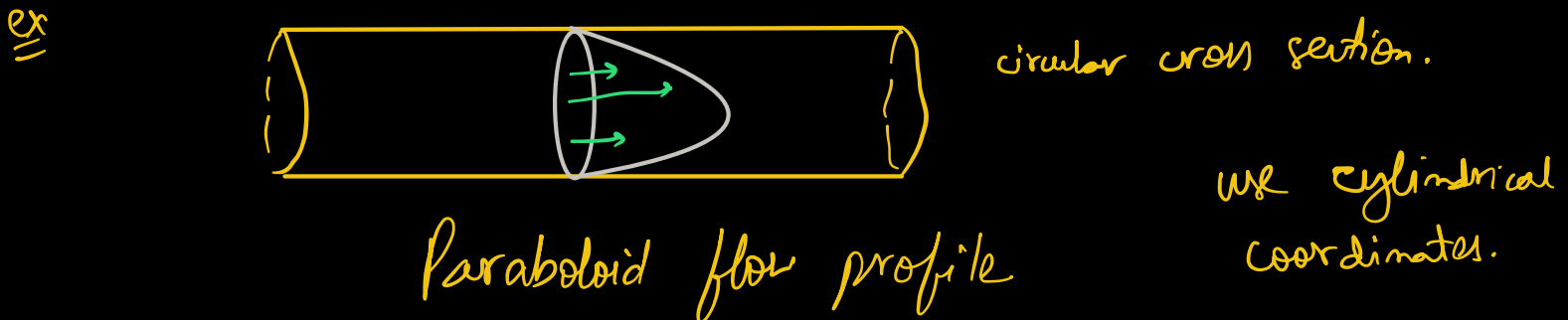
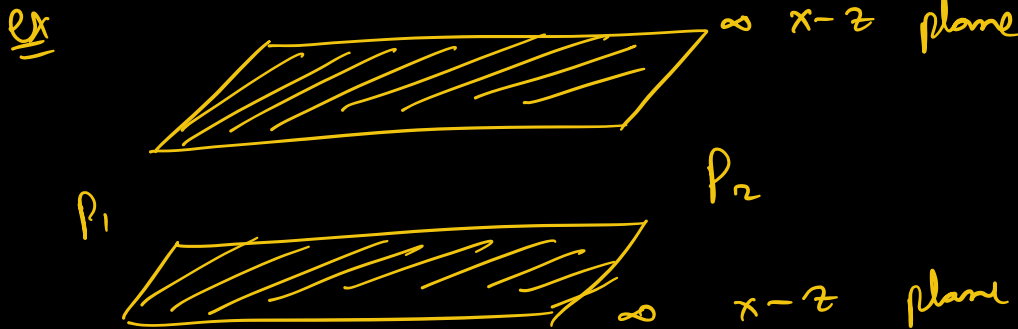
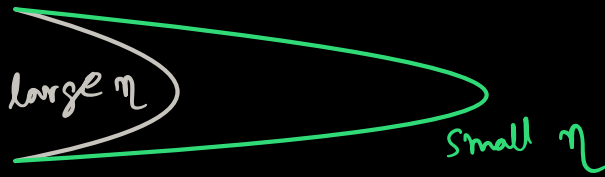
$$\frac{dP}{dx} = -\frac{\Delta P}{L}$$



So,

$$V_x = -\frac{dp}{dx} \cdot \frac{y \cdot (y-h)}{2\eta}$$

flow under constant pressure gradient.



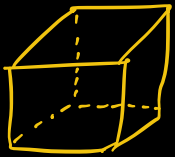
we got steady non accelerating solution along  $x$ .

$$V_x = V_x(y)$$

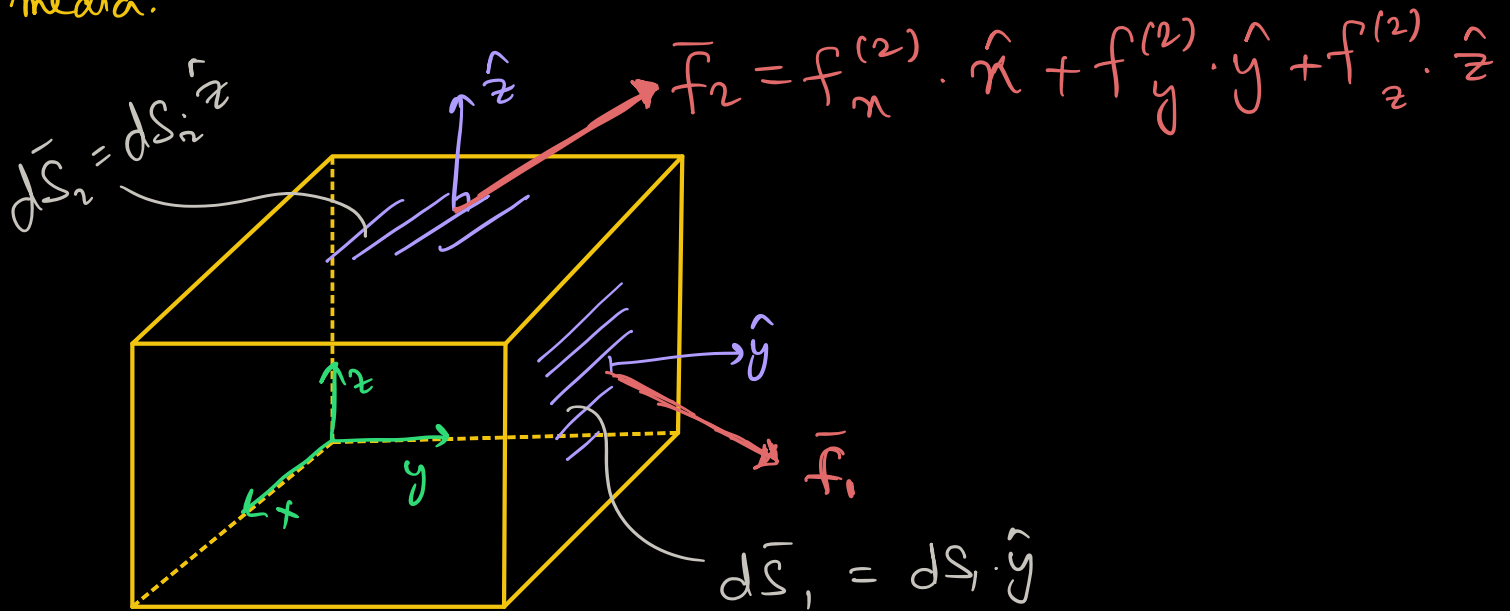
So; net force on fluid element is zero actually.

find force due to neighbours... Drag.

# Stress Tensor $\tilde{\sigma}_{ij}$



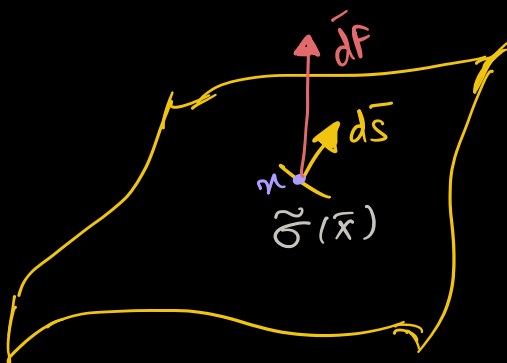
consider any volume element in continuum media.



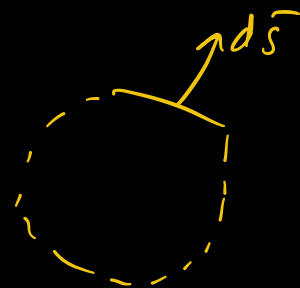
$$\bar{f} = \tilde{\sigma} \cdot d\bar{s}$$

$\sigma$  is a matrix

$$= \begin{bmatrix} \vdots & \vdots & \vdots \end{bmatrix} \begin{bmatrix} \vdots \\ \vdots \\ \vdots \end{bmatrix}$$



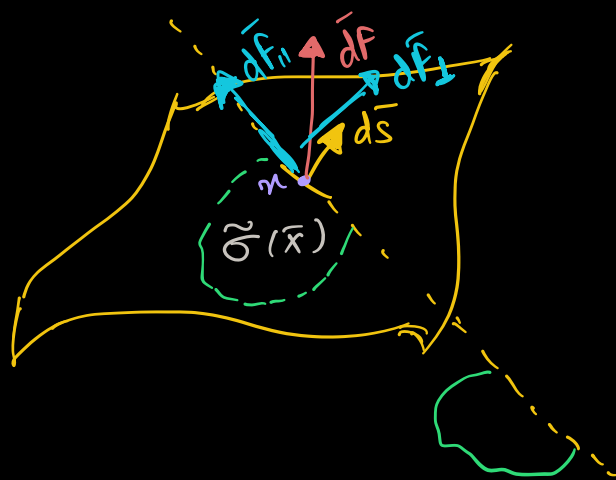
$$d\bar{F} = \tilde{\sigma} \cdot d\bar{s}$$



This surface belongs to some volume.

$d\vec{F} \cdot d\vec{s} > 0 \Rightarrow$  neighbours are pulling the volume element  
along the same direction as  $d\vec{s}$ .  
(ie; outward normal)

$d\vec{F} \cdot d\vec{s} < 0 \Rightarrow$  neighbours are pressing or compressing the  
volume element.

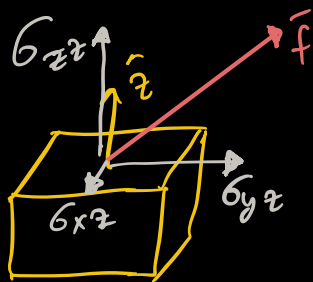
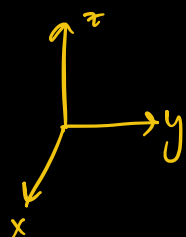


$$d\vec{F} = dF_{||} \cdot \hat{t} + dF_{\perp} \cdot \hat{n}$$

shear force

extensile (when  $> 0$ )  
or  
compressive (when  $< 0$ )

ex Unit surface

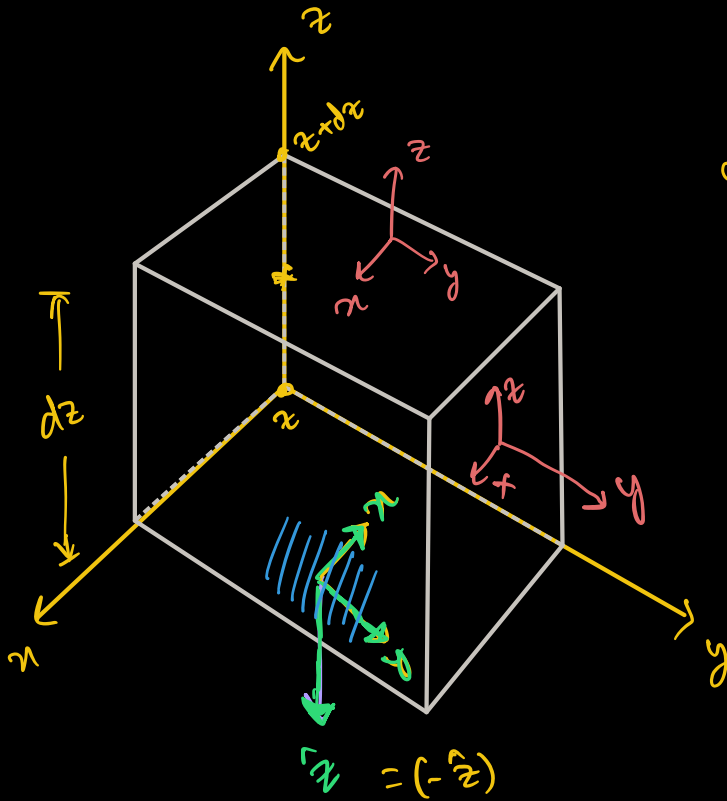
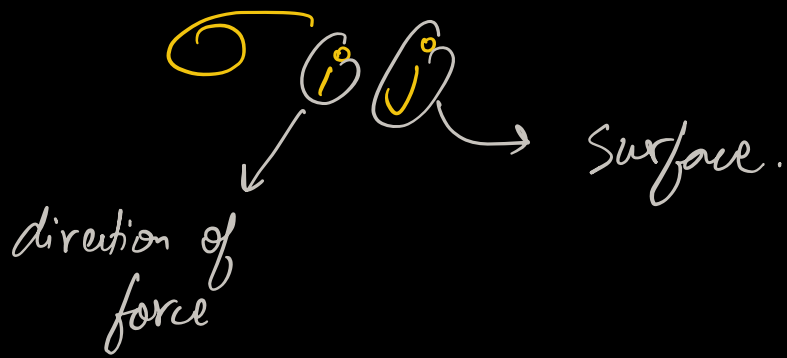


$$\vec{F} = \sigma \cdot \hat{z}$$

$$= \begin{bmatrix} \sigma_{xx} & \sigma_{xy} & \sigma_{xz} \\ \sigma_{yx} & \sigma_{yy} & \sigma_{yz} \\ \sigma_{zx} & \sigma_{zy} & \sigma_{zz} \end{bmatrix} \begin{bmatrix} 0 \\ 0 \\ 1 \end{bmatrix} = \begin{bmatrix} \sigma_{xz} \\ \sigma_{yz} \\ \sigma_{zz} \end{bmatrix}$$

we chose unit surface,  
and now directly interpreting  $\sigma_{ij}$  as forces.

$\sigma_{ij}$  = force on unit area along  $\hat{i}$  direction on the  $j^{\text{th}}$  surface



can take origin at midpoint.

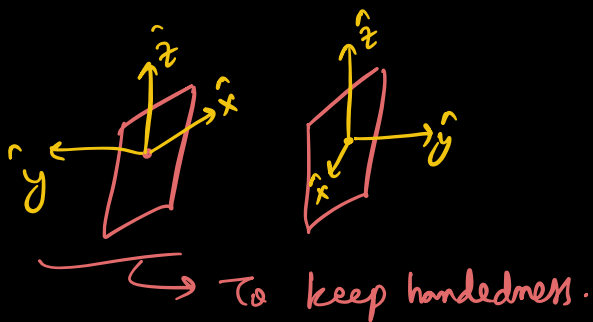
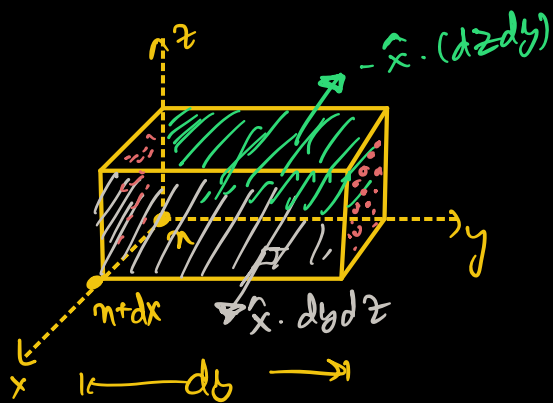
if all elements of  $\vec{\tau}$  are positive  $\Rightarrow$  system is under stretch from all side.

$$(-\sigma_{ij}) \cdot \hat{z} = -ve$$

Here we have 6 faces

Then,

total force along  $x$ -direction.



$$F_x = (\sigma_{xx}(x+dx) - \sigma_{xx}(x))dydz + (\sigma_{xy}(y+dy) - \sigma_{xy}(y))dxdz + (\sigma_{xz}(z+dz) - \sigma_{xz}(z))dxdy$$

$$\frac{F_x}{dxdydz} = \text{force per unit volume}$$

$$\frac{F_x}{dxdydz} = \frac{\sigma_{xx}(x+dx) - \sigma_{xx}(x)}{dx} + \frac{\sigma_{xy}(y+dy) - \sigma_{xy}(y)}{dy} + \frac{\sigma_{xz}(z+dz) - \sigma_{xz}(z)}{dz}$$

$$= \frac{d\sigma_{xx}}{dx} + \frac{d\sigma_{xy}}{dy} + \frac{d\sigma_{xz}}{dz}$$

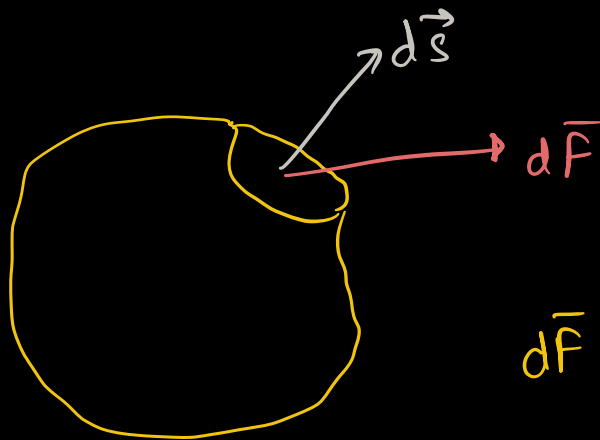
$$= \frac{\partial}{\partial x_i} \sigma_{xi} = f_x \rightarrow \text{force density.}$$

$$f_y = \frac{\partial}{\partial x_i} \tilde{\sigma}_{yi}$$

$$f_z = \frac{\partial}{\partial x_i} \tilde{\sigma}_{zi}$$

$$f_i = \frac{\partial}{\partial x_j} \tilde{\sigma}_{ij}$$

$$\bar{f} = \bar{\nabla} \cdot \tilde{\sigma}$$



$$d\bar{F} = \tilde{\sigma} \cdot d\bar{s}$$

Total force on the volume element  $V$

$$\bar{F} = \oint_{\partial V} \tilde{\sigma} \cdot d\bar{s} \stackrel[\text{Theorem}]{\text{Div}} = \int_V \bar{\nabla} \cdot \tilde{\sigma} dV$$

$$\bar{F} = \int_V \bar{f} dV = \int_V \bar{\nabla} \cdot \tilde{\sigma} dV$$

net force

force density

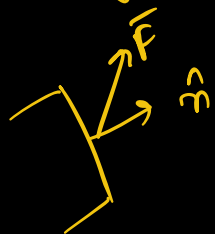
$$\bar{f} = \bar{\nabla} \cdot \tilde{\sigma}$$

$$\bar{F} = \int_V \bar{f} dV$$

defining equation  
for force density  
 $\bar{f}$ .



Identify the surface by identifying normal  $\hat{n}$



$$d\vec{S} = dS \hat{n}$$

$$\vec{F} = \vec{\sigma} \cdot d\vec{S} \quad \text{Force away from the volume element}$$

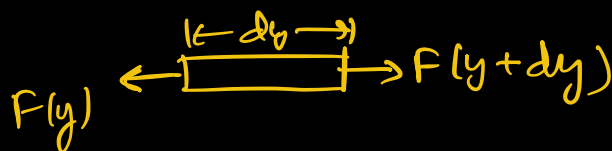
$$= (:) \cdot (:) \quad \text{volume element}$$

$$\begin{pmatrix} \sigma_{11}(x) & \sigma_{12}(x) & \sigma_{13} \\ \sigma_{21}(x) & \sigma_{22}(x) & \sigma_{23} \\ \sigma_{31} & \sigma_{32} & \sigma_{33} \end{pmatrix} \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix} = \begin{pmatrix} \sigma_{11} \\ \sigma_{21} \\ \sigma_{31} \end{pmatrix}$$

$$\begin{pmatrix} \sigma_{11} & \sigma_{12} & \sigma_{13} \\ \sigma_{21} & \sigma_{22} & \sigma_{23} \\ \sigma_{31} & \sigma_{32} & \sigma_{33} \end{pmatrix} \begin{pmatrix} -1 \\ 0 \\ 0 \end{pmatrix} = - \begin{pmatrix} \sigma_{11} \\ \sigma_{21} \\ \sigma_{31} \end{pmatrix}$$

$$\begin{pmatrix} \sigma_{11} & \sigma_{12} & \sigma_{13} \\ \sigma_{21} & \sigma_{22} & \sigma_{23} \\ \sigma_{31} & \sigma_{32} & \sigma_{33} \end{pmatrix} \begin{pmatrix} 0 \\ 1 \\ 0 \end{pmatrix} = \begin{pmatrix} \sigma_{12} \\ \sigma_{22} \\ \sigma_{32} \end{pmatrix}$$

Sx11  
1-D



In 1-D, only longitudinal force.  
Tangential force at not here

If  $F$  is positive definite  $\Rightarrow$  Stretched

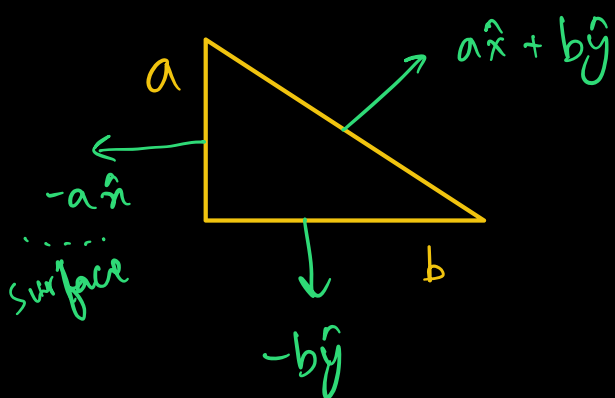
$$\xrightarrow{F(y) < 0} \boxed{\phantom{000}} \rightarrow F(y+dy)$$

Compressed if  $F(y) > F(y+dy)$

Stretched if  $F(y) < F(y+dy)$

Ex 11 2-D

let stress matrix be positive



$$\sigma = \begin{pmatrix} 2 & 1 \\ 1 & 2 \end{pmatrix}$$

constant, positive definite.

Net force on triangle

$$\vec{\sigma}(-a\hat{x}) + \vec{\sigma}(-b\hat{y}) + \vec{\sigma}(a\hat{x} + b\hat{y})$$

$$= \vec{\sigma}(-a\hat{x} - b\hat{y} + a\hat{x} + b\hat{y})$$

$$= \vec{\sigma}(\vec{0}) = \vec{0}$$

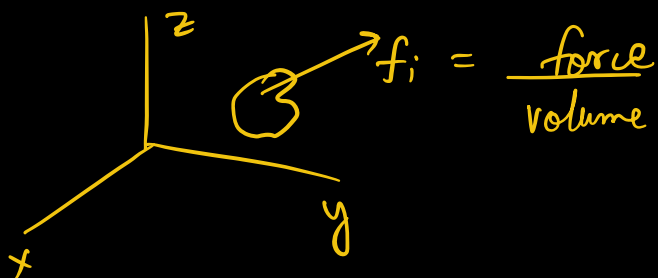
(because  $\vec{\sigma}$  is constant vector)

So, net force on the triangle is zero.

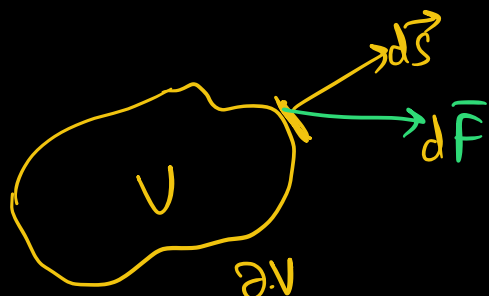
$\rightarrow$  This could not be done if  $\sigma$  varied in space.



We had  $f_i = \frac{\partial \sigma_{ij}}{\partial x_j}$



We are only thinking of contact forces for now; while thinking stress tensor.



$$\bar{F} = \oint_{\partial V} d\bar{F} = \oint_{\partial V} \tilde{\sigma} \cdot d\vec{S} \stackrel{\text{Divergence Theorem}}{=} \int_V (\bar{\nabla} \cdot \tilde{\sigma}) dV$$

In the absence of external force.

$\bar{F} = \int_V \bar{f} \cdot dV$  is the defining eq<sup>n</sup> for force density  $\bar{f}$

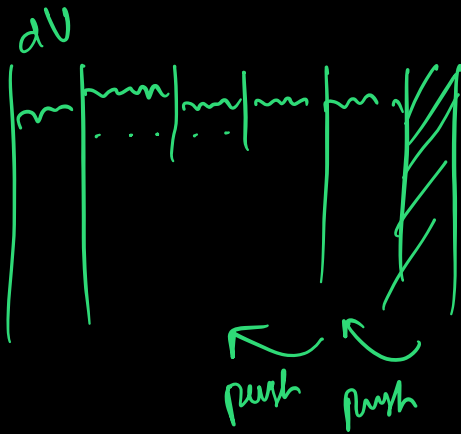
$$\Rightarrow \int_V \bar{f} dV = \int_V (\bar{\nabla} \cdot \tilde{\sigma}) dV$$

$$\Rightarrow \boxed{\bar{f} = (\bar{\nabla} \cdot \tilde{\sigma})}$$

when we have only surface forces.  
(no forces like electrostatics, gravity, ... etc)

$$\int_V dV \cdot [\bar{f} - \bar{\nabla} \cdot \tilde{\sigma}] = 0 \text{ for arbitrary } V$$

$$\Rightarrow \boxed{\bar{f} = \bar{\nabla} \cdot \tilde{\sigma}}$$



So  $\bar{f}$  is a local expression, because forces are short distance forces.

Now does  $\tilde{\sigma}$  look like?

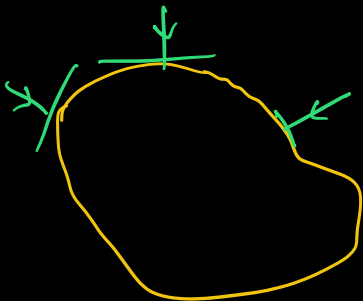
For a fluid,  $\sigma_{ij}(\bar{x}) = -P(\bar{x})\delta_{ij} + \sigma'_{ij}(\bar{x})$

$P \equiv$  conventional pressure

This term will always be there

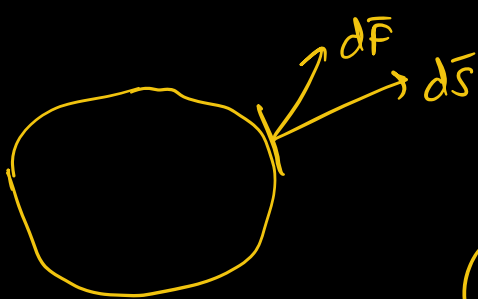
$\sigma'_{ij}$  stress due to non ideal fluid.  
(ex, fluid with friction)  
(friction between volume element of fluids)

Consider a volume element



$-P\delta_{ij}$  ensures that pressure always act normal to surface.

Our notion of pressure (always  $\perp$  to surface, & compresses the volume element)



$$d\bar{F} = \tilde{\sigma}_p \cdot d\bar{s}$$

$$(\tilde{\sigma}_p)_{ij} = -p \cdot \delta_{ij}$$

$$dF_i = -p \cdot \delta_{ij} \cdot dS_j = -p \cdot dS_i$$

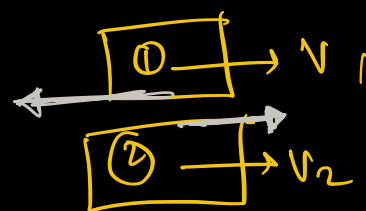
$$\Rightarrow \boxed{d\bar{F} \propto d\bar{s}}$$

i.e;

$$\boxed{d\bar{F} = -p d\bar{s}}$$

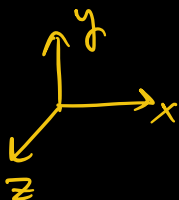
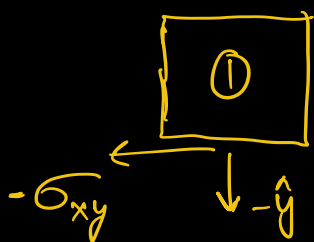
Now  $\sigma'_{ij}$

$$\textcircled{1} \sigma'_{ij}(x) \propto \frac{\partial v_i}{\partial x_j}$$



let  $v_1 > v_2$

so there will be friction

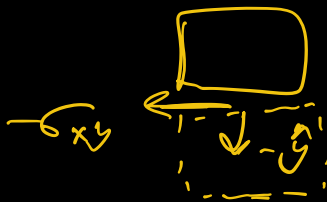


Example 1 To derive  $\sigma'_{ij} \propto \frac{\partial v_i}{\partial x_j}$  or  $\sigma'_{ij} \propto -\frac{\partial v_i}{\partial x_j}$

$$\text{let } \frac{\partial v_x}{\partial y} > 0$$



$$\sigma_{xy} = \frac{\partial v_x}{\partial y}$$



so,  $\sigma'_{ij} \propto \frac{\partial v_i}{\partial x_j}$   
 ✓ correct.

We could choose higher order term

$$\sigma_{ij} \propto \sum_k \frac{\partial v_i}{\partial x_k} \frac{\partial v_j}{\partial x_k}$$

We can't have  $\sigma_{ij} \propto v_i v_j$

because for uniform  $v$  also we will get friction.

So  $v_i v_j$  is not possible as friction term.

$\sum_k \alpha_k v_i v_j v_k$  is possible.

What about  $\sigma'_{ij} \propto v_i \cdot \frac{\partial v_j}{\partial x_k} \cdot v_k$

This is possible, but it is higher order in  $v$ .

$\sigma'_{ij}$  should depend on gradients & has to be 2<sup>nd</sup> order.

②  $\sigma'_{ij} \propto$  Terms symmetric in  $i$  &  $j$

like  $\left( \frac{\partial v_i}{\partial x_j} + \frac{\partial v_j}{\partial x_i} \right)$

$$\frac{\partial v_i}{\partial x_j} = \frac{1}{2} \left( \frac{\partial v_i}{\partial x_j} + \frac{\partial v_j}{\partial x_i} \right) + \frac{1}{2} \left( \frac{\partial v_i}{\partial x_j} - \frac{\partial v_j}{\partial x_i} \right)$$

Due to global rotational symmetry, only symmetric combinations are allowed.

Rotating bucket of water.



$\vec{v} = \vec{\omega}_0 \times \vec{r} \Rightarrow$  This corresponds to uniform rotation with  $\vec{\omega}_0$ .

$\left( \left( \left( \right. \right. \right. \text{move at same rotational speed, so no relative motion} \Rightarrow \text{so no friction.}$

so,  $\sigma'_{ij} = 0$

and it is only possible if

$$\sigma'_{ij} \propto \text{sym} \left( \frac{\partial v_i}{\partial x_j} \right)$$

$$\therefore \sigma'_{ij} \propto \frac{1}{2} \left( \frac{\partial v_i}{\partial x_j} + \frac{\partial v_j}{\partial x_i} \right)$$

$$\sigma'_{ij} = \eta \cdot \left( \frac{\partial v_i}{\partial x_j} + \frac{\partial v_j}{\partial x_i} \right)$$

↙ some constant.

Should take antisymmetric part, because that will not go to zero for cases like  $\vec{v} = \vec{\omega} \times \vec{x}$ .

Still can have,

$$\lambda \cdot \delta_{ij} \cdot \left( \frac{\partial x_k}{\partial x_k} \right)$$

$$= \lambda \cdot \delta_{ij} (\vec{\nabla} \cdot \vec{v})$$

$$\sigma'_{ij} = \eta \cdot \left( \frac{\partial v_i}{\partial x_j} + \frac{\partial v_j}{\partial x_i} \right) + \lambda \cdot \delta_{ij} \cdot \vec{\nabla} \cdot \vec{v}$$

Conventionally written in following form

$$\sigma'_{ij} = \eta \cdot \left( \frac{\partial v_i}{\partial x_j} + \frac{\partial v_j}{\partial x_i} - \frac{2}{3} \delta_{ij} \cdot \frac{\partial v_k}{\partial x_k} \right) + \lambda' \cdot \delta_{ij} \cdot \frac{\partial v_k}{\partial x_k}$$

↙  $\lambda$  piece is split in this form, so that the trace of

$$\left( \frac{\partial v_i}{\partial x_j} + \frac{\partial v_j}{\partial x_i} - \frac{2}{3} \delta_{ij} \cdot \frac{\partial v_k}{\partial x_k} \right) \text{ vanishes.}$$

often written as

$$\sigma'_{ij} = \eta \cdot \left( \frac{\partial v_i}{\partial x_j} + i \leftrightarrow j - \frac{2}{3} \delta_{ij} \cdot \frac{\partial v_k}{\partial x_k} \right) + \lambda \cdot \delta_{ij} \cdot \frac{\partial v_k}{\partial x_k}$$

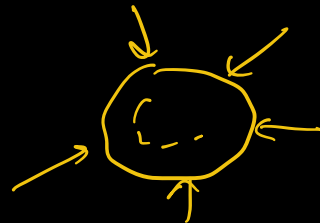
has no change in volume. ; but does volume deformation.

This associated to volume change



$\eta$   $\equiv$  Dynamic Viscosity.

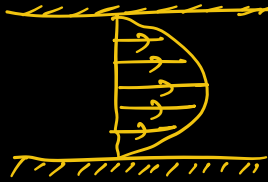
$\lambda$   $\equiv$  Viscosity due to compressibility



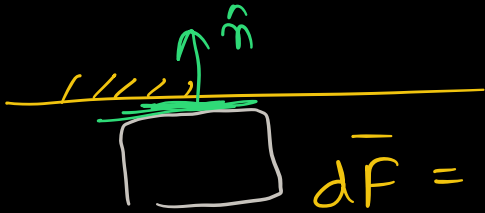
The term  $\frac{\partial v_k}{\partial x_k}$  is due to compression of the fluid. , and contribute to non-ideal part of fluid.

## Application

①



what are kind of forces acting on solid-liquid interface.

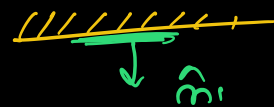


$$d\bar{F} = \bar{\sigma} \cdot \hat{n}$$

= force on the fluid element below due to the solid.

Force on the solid due to liquid  
(Reaction force)

→ called Drag.



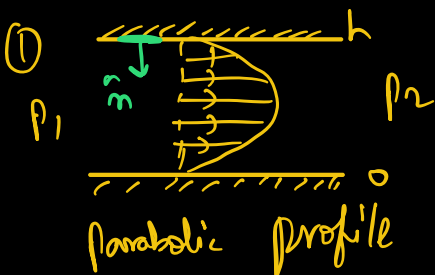
$$d\bar{F}_1 = \bar{\sigma} \cdot \hat{n}_1$$

$$= -\bar{\sigma} \cdot \hat{n}$$

note  $\hat{n}_1 = -\hat{n}$

## Application

①



$$v_x = - \left( \frac{p_1 - p_2}{2\eta L} \right) y \cdot (y - h)$$



Drag force on solid (at  $y=h$ )

$$= \bar{F}_h = \bar{\sigma} \cdot \hat{n}$$

for incompressible fluid

$$\bar{\sigma} = \eta \cdot \left( \frac{\partial v_i}{\partial x_j} + \frac{\partial v_j}{\partial x_i} \right) + \text{pressure term.}$$

$$F_i = \left[ \eta \cdot \left( \frac{\partial v_j}{\partial x_j} + \frac{\partial v_j}{\partial x_i} \right) - p \cdot \delta_{ij} \right] \cdot \hat{n}_j$$

here,

$$\hat{n} = \begin{pmatrix} 0 \\ -1 \end{pmatrix} = -\hat{y}$$

$$F_x = \eta \cdot \left[ \underbrace{\left( \frac{\partial v_x}{\partial x} + \frac{\partial v_x}{\partial x} \right) n_x}_{j=x} + \underbrace{\left( \frac{\partial v_x}{\partial y} + \frac{\partial v_y}{\partial x} \right) n_y}_{j=y} \right]$$

$$= -\eta \cdot \left( \frac{\partial v_x}{\partial y} + \frac{\partial v_y}{\partial x} \right)$$

$$= -\eta \cdot \frac{\partial v_x}{\partial y}$$

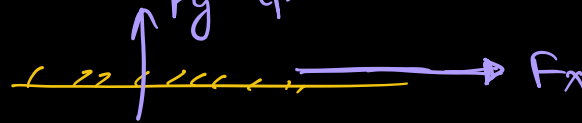
note

$$\frac{\partial v_x}{\partial y} \Big|_{y=h} < 0$$

$$\Rightarrow F_x = - \frac{(p_1 - p_2)}{2\eta L} \cdot (2y - h) (-1) \Big|_{y=h}$$

$$\Rightarrow F_x = \frac{(P_1 - P_2)h}{2\eta L}$$

$F_y$  (pressure) : coming due to pressure term.



A horizontal line represents a plate. A blue arrow labeled  $F_x$  points to the right from the center of the plate. A blue arrow labeled  $F_y$  points upwards from the center of the plate. To the left of the plate, there are several small blue arrows pointing to the right, representing a pressure distribution.



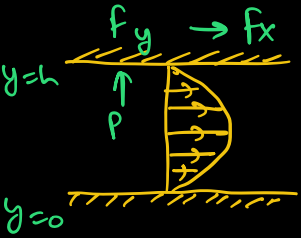
A horizontal line represents a plate. Above the line, there are several small blue arrows pointing to the right, representing a pressure distribution.

What is  $F_y$  on upper plate.

# Continuum Mechanics

- Shoail Akhtar

Lec 11 (19<sup>th</sup> Feb, 2021 ; 3:30pm to 5:00pm)



$$\bar{F} = -\tilde{\sigma} \cdot \hat{n}$$

volume is not changed under pure shear.

Has zero trace

$$\rho \left[ \frac{\partial \bar{v}}{\partial t} + (\bar{v} \cdot \nabla) \bar{v} \right] = -\nabla p + \eta \nabla^2 \bar{v} + \bar{f}_{ext}$$

$$\sigma_{ij} = -p \cdot \delta_{ij} + \eta \cdot \left( \frac{\partial v_i}{\partial x_j} + i \leftrightarrow j - \frac{2}{3} \delta_{ij} \cdot \frac{\partial v_k}{\partial x_k} \right)$$

shear viscosity

$$+ \xi \cdot \delta_{ij} \cdot \frac{\partial v_k}{\partial x_k}$$

Bulk viscosity

(not physically a friction...)

\* related to expansion & compression of volume.

for incompressible fluid.

$\rho \frac{D \bar{v}}{Dt}$  = force on fluid element which we are following in space.

$$\int \rho \cdot \frac{D \bar{v}}{Dt} \cdot dV = \oint \tilde{\sigma} \cdot d\bar{s} + \int f_{ext} \cdot dV$$

↑  
surface
↑  
volume.



$$= \int (\nabla \cdot \tilde{\sigma}) dV + \int f_{ext} dV$$

$$\rho \frac{D\bar{v}}{Dt} = -\bar{\nabla} \cdot \tilde{\sigma} + \bar{f}_{ext}$$

$$\rho \frac{Dv_i}{Dt} = \frac{\partial}{\partial x_j} \tilde{\sigma}_{ij} + \bar{f}_{ext,i}$$

lets work only with  $\rho \frac{Dv_i}{Dt} = \frac{\partial}{\partial x_j} \tilde{\sigma}_{ij}$

$$= \frac{\partial}{\partial x_j} \left[ -p \cdot \delta_{ij} + \eta \cdot \left( \frac{\partial v_i}{\partial x_j} + \frac{\partial v_j}{\partial x_i} - \frac{2}{3} \delta_{ij} \cdot \frac{\partial v_k}{\partial x_k} \right) + \xi \cdot \delta_{ij} \cdot \frac{\partial v_k}{\partial x_k} \right]$$

$$\Rightarrow \rho \cdot \frac{Dv_i}{Dt} = -\frac{\partial p}{\partial x_i} + \eta \cdot \left( \sum_j \frac{\partial^2 v_i}{\partial x_j^2} + \frac{\partial}{\partial x_i} \left( \frac{\partial v_j}{\partial x_j} \right) + \left( \xi - \frac{2\eta}{3} \right) \frac{\partial}{\partial x_i} \left( \frac{\partial v_k}{\partial x_k} \right) \right)$$

$$\rho \cdot \frac{D\bar{v}}{Dt} = -\bar{\nabla} p + \eta \cdot (\nabla^2 \bar{v} + \bar{\nabla} (\bar{\nabla} \cdot \bar{v})) + \left( \xi - \frac{2\eta}{3} \right) \cdot \bar{\nabla} (\bar{\nabla} \cdot \bar{v})$$

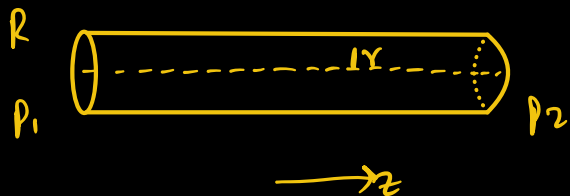
for incompressible fluid  $\bar{\nabla} \cdot \bar{v} = 0$

$$\rho \cdot \frac{D\bar{v}}{Dt} = -\bar{\nabla} p + \eta \nabla^2 \bar{v} + \left( \xi + \frac{\eta}{3} \right) \bar{\nabla} (\bar{\nabla} \cdot \bar{v}) + \bar{f}_{ext}$$

Full Navier Stokes Equation.

ex 11 for fluid in gravity  $\vec{f}_{\text{ext}} = -\rho g \hat{z}$

ex 11



$$p_1 > p_2$$

$$\text{Symmetry} \Rightarrow V_z = V_z(r)$$

$$V_\phi = 0$$

$$V_r = 0$$

$$V_z(r=R) = 0 \quad (\text{No slip boundary condition})$$

The continuity equation;

$$\frac{1}{r} \cdot \frac{\partial}{\partial r} (r \cdot V_r) + \cancel{\frac{1}{r} \cdot \frac{\partial V_\phi}{\partial \phi}} + \cancel{\frac{\partial V_z}{\partial z}} = 0$$

$V_z = V_z(r)$  due to  
translational symmetry  
along  $z$  axis.

$$\Rightarrow \frac{\partial}{\partial r} (r \cdot V_r) = 0$$

$$\Rightarrow r \cdot V_r = \text{constant}$$

constant  $\neq 0 \Rightarrow \boxed{V_r(R) \neq 0}$  contradicts no slip  
boundary condition.

$\Rightarrow$  There is no radial velocity

so; constant = 0, and we get  $V_r = 0$ .

$$V_z(r), \quad p(r)$$

$$r \text{ equation: } 0 = -\frac{1}{\rho} \cdot \frac{\partial p}{\partial r} + g \Rightarrow p \text{ not function of } r$$

$\phi$  equation:  $0 = -\frac{1}{\rho \cdot r} \cdot \frac{\partial \rho}{\partial \phi} = 0 \Rightarrow \rho$  is not function of  $\phi$ .

$z$  equation:  $\cancel{\nu r} \cdot \frac{\partial}{\partial r} \nu_z = \underbrace{-\frac{1}{\rho} \cdot \frac{\partial \rho(z)}{\partial z}}_{\text{function of } z} + \underbrace{\frac{\nu}{r} \cdot \frac{\partial}{\partial r} \left( r \cdot \frac{\partial \nu_z}{\partial r} \right)}_{\text{function of } r}$

$$\boxed{\nu = \frac{\eta}{\rho}}$$

$\Rightarrow$  They are constant.

$$\underbrace{= \frac{1}{\rho} \cdot \frac{\partial \rho(z)}{\partial z}}_z = \underbrace{\frac{\nu}{r} \cdot \frac{\partial}{\partial r} \left( r \cdot \frac{\partial \nu_z(r)}{\partial r} \right)}_r$$

So, They are some constant  $C_1$

$$\underbrace{\frac{1}{\rho} \cdot \frac{\partial \rho}{\partial z} = C_1}_{\rightarrow \rho(z) \text{ linear function of } z.}$$

with  $\rho_1$  &  $\rho_2$  at  $z=0$  &  $L$  respectively.

$$\rightarrow \frac{\eta}{\rho} \left( r \cdot \frac{\partial \nu_z}{\partial r} \right) = C_1 \cdot \frac{r^2}{2} + C_2$$

$$\frac{\partial \nu_z}{\partial r} = \frac{\rho}{\eta} \left[ C_1 \cdot \frac{r}{2} + \frac{C_2}{r} \right]$$

$$\Rightarrow \nu_z = \frac{\rho}{\eta} \left[ C_1 \cdot \frac{r^2}{4} + C_2 \cdot \ln r \right] + C_3$$

$C_2$  must be zero, because then  
 $V_z(r=0) \rightarrow -\infty$

To avoid  $\infty$  at  $r=0$ ,  $C_2=0$

$$\Rightarrow V_z = \frac{\rho}{\eta} \cdot C_1 \cdot \frac{r^2}{4} + C_3$$

$$\frac{1}{\rho} \cdot \frac{\partial p}{\partial z} = C_1$$

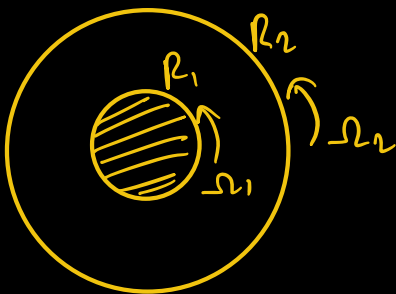
$C_1$  is known from here

$$V_z(R) = 0 \Rightarrow V_z(r) = \frac{C_1 \rho}{4\eta} (r^2 - R^2)$$

$$V_z(r) = \frac{C_1 \rho}{4\eta} (r^2 - R^2)$$

Parabolic flow

Example



concentric circular cylinders (rotating)

This is viscous driven force

Polar analogue of  
 Couet flow.

$\text{-----} \rightarrow v_1$

$\text{~~~~~} \rightarrow v_2$

$$v_1 \neq v_2$$

$\infty$  cylinders ;  $z$  ranges over  $(-\infty, \infty)$

$$V_\phi(r), p(r)$$

$$r \text{ equation: } -\frac{V_\phi^2}{r} = -\frac{1}{f} \cdot \frac{\partial p}{\partial r}$$

$$\phi \text{ equation: } V_\phi \cdot \frac{1}{r} \cdot \frac{\partial V_\phi}{\partial \phi} = \nu \left[ \frac{1}{r} \cdot \frac{\partial}{\partial r} \left( r \cdot \frac{\partial V_\phi}{\partial r} \right) - \frac{V_\phi}{r^2} \right]$$

→ solve this  
& get  $V_\phi$

$$\frac{\partial V_\phi}{\partial r} + r \cdot \frac{\partial^2 V_\phi}{\partial r^2} - \frac{V_\phi}{r} = 0$$

$$V_\phi = r^n$$

$$n r^{n-1} + r \cdot (n)(n-1) \cdot r \cdot r^{n-2} - r^{n-1} = 0$$

$$\Rightarrow n + n(n-1) - 1 = 0$$

$$\Rightarrow n + n^2 - n - 1 = 0$$

$$\Rightarrow \boxed{n = \pm 1}$$

$$\boxed{V_\phi = Ar + \frac{B}{r}}$$

$$\left. \begin{aligned} V_\phi(R_2) &= \Omega_2 R_2 \\ V_\phi(R_1) &= \Omega_1 R_1 \end{aligned} \right\} \rightarrow \text{get } A \text{ and } B.$$

$$V_\phi(r) = \frac{\Omega_2 R_2^2 - \Omega_1 R_1^2}{R_2^2 - R_1^2} \cdot r + \frac{\Omega_1 - \Omega_2}{R_2^2 - R_1^2} \cdot R_1^2 \cdot R_2^2 \cdot \frac{1}{r}$$



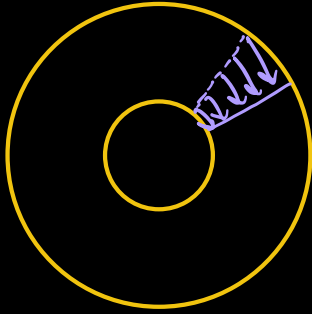
## limiting cases

①  $\Omega_1 = \Omega_2 \equiv \Omega$

Then

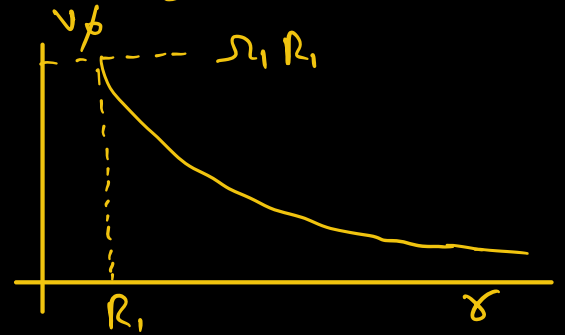
$$V_\phi(r) = \Omega r$$

linear profile in  $r$



②  $\Omega_2 = 0, R_2 \rightarrow \infty$  (no outer boundary)

$$V_\phi(r) = (\Omega_2 - \Omega_1) R_1^2 \cdot \frac{1}{r}$$



# Continuum Mechanics

— Shoail Akhtar

Lec 12 (5<sup>th</sup> March, 2021 ; 3:30 pm to 5:00 pm)



$$\rho \left[ \frac{\partial v_i}{\partial t} + (\bar{v} \cdot \bar{\nabla}) v_i \right] = -\bar{\nabla}_i p + \rho g_i \quad \text{under some force.}$$

$$+ \eta \nabla^2 v_i + \left( \zeta + \frac{1}{3} \right) \bar{\nabla} (\bar{\nabla} \cdot \bar{v}) \quad \bar{f} = -\bar{\nabla} \phi \quad (\text{potential force})$$

① when  $\bar{v}$  small ( $Re \ll 1$ ) + steady state  
drop the LHS  $\Rightarrow$  Stokes equation.

② Incompressible  $\Rightarrow \bar{\nabla} \cdot \bar{v} = 0$

③  $\eta = 0$   
 $\zeta = 0 \Rightarrow$  Ideal fluid.  $\Rightarrow$  Euler Equation.

Here (a) Incompressible (b) Ideal  $\Rightarrow \eta$  &  $\zeta$  are zero.

$$\bar{v} \times (\bar{\nabla} \times \bar{v}) = \bar{\nabla} \left( \frac{v^2}{2} \right) - (\bar{v} \cdot \bar{\nabla}) \bar{v}$$

$$\rho \left[ \frac{\partial \bar{v}}{\partial t} - \bar{v} \times (\bar{\nabla} \times \bar{v}) \right] = -\rho \bar{\nabla} \left( \frac{v^2}{2} \right) - \bar{\nabla} p + \rho \bar{g}$$

divide by  $\rho$

$$\frac{\partial \bar{v}}{\partial t} - \bar{v} \times (\bar{\nabla} \times \bar{v}) = -\bar{\nabla} \left( \frac{v^2}{2} \right) - \bar{\nabla} p + \bar{g}$$

$$\text{where } \bar{\nabla} p \equiv \frac{\bar{\nabla} p}{\rho}$$

for incompressible

$$\bar{\nabla} p = \bar{\nabla} \left( \frac{p}{\rho} \right)$$

$$\Rightarrow p = \frac{P}{\rho}$$

$$\left[ \frac{\partial \bar{v}}{\partial t} - \bar{v} \times (\bar{\nabla} \times \bar{v}) \right] = -\bar{\nabla} \left[ \frac{v^2}{2} + \frac{P}{\rho_0} + z \cdot g \right]$$

$\rightarrow \frac{\phi(\bar{r})}{\rho_0}$  in general.

Now, if steady state + Curl free flow (Irrotational flow)  
 $\frac{\partial \bar{v}}{\partial t} = 0$   $\bar{\nabla} \times \bar{v} = 0$

Then  $0 = -\bar{\nabla} \psi(\vec{x}, t)$  → since steady state, it can't be t dependent

$$\Rightarrow 0 = -\bar{\nabla} \psi(\vec{x})$$

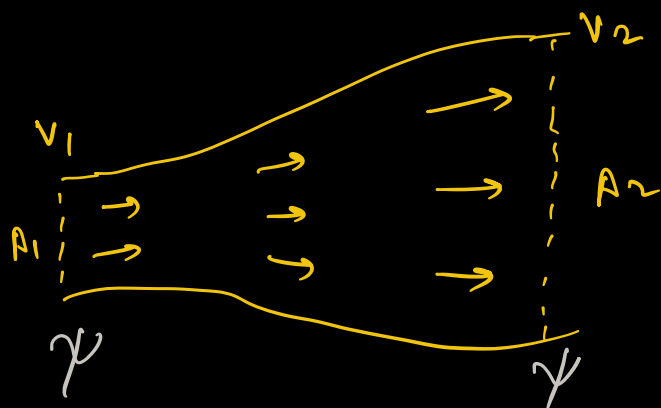
where  $\psi(\vec{x}) = \frac{v^2}{2} + \frac{P}{\rho_0} + z \cdot g$

The function  $\psi(\vec{x})$  is called Stream Function.

Stream Function:  $\psi = \frac{v^2}{2} + \frac{P}{\rho_0} + \frac{\Omega}{\rho_0}$

where  $\bar{f} = -\bar{\nabla} \Omega$

since  $\bar{\nabla} \psi = 0 \Rightarrow \psi$  does not vary in space  
 i.e.  $\psi$  is constant.



$A_1 v_1 = A_2 v_2$  from continuity.

$$\psi = \frac{v_1^2}{2} + \frac{P_1}{\rho_0} + \frac{\Omega_1}{\rho_0} = \frac{v_2^2}{2} + \frac{P_2}{\rho_0} + \frac{\Omega_2}{\rho_0}$$

Let  $\Omega_1 = \Omega_2 \Rightarrow \frac{v_1^2}{2} + \frac{P_1}{\rho_0} = \frac{v_2^2}{2} + \frac{P_2}{\rho_0}$

$$\psi_1 = \frac{v_1^2}{2} + \frac{p}{\rho_0} = \psi_2 = \frac{v_2^2}{2} + \frac{p_2}{\rho_0}$$

$$\psi_1 = \frac{v_1^2}{2} + \frac{p_1}{\rho_0} + \frac{\Omega_1}{\rho_0} = \psi_2 = \frac{v_2^2}{2} + \frac{p_2}{\rho_0} + \frac{\Omega_2}{\rho_0}$$



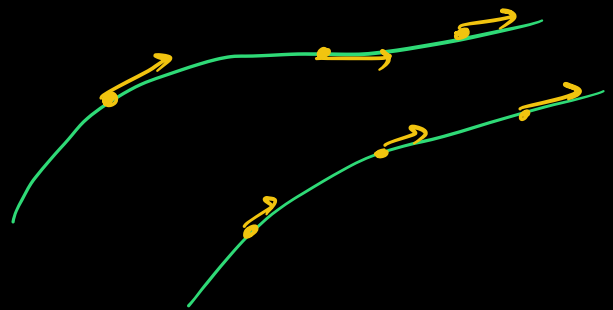
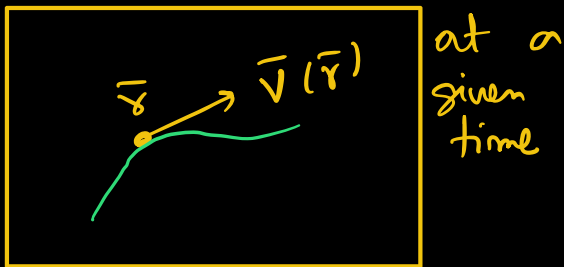
$$\frac{\nabla \times \bar{v}}{2} = \bar{\omega} \quad (\text{vorticity})$$

$$\frac{\partial \bar{v}}{\partial t} - 2\bar{v} \times \bar{\omega} = -\nabla \psi$$

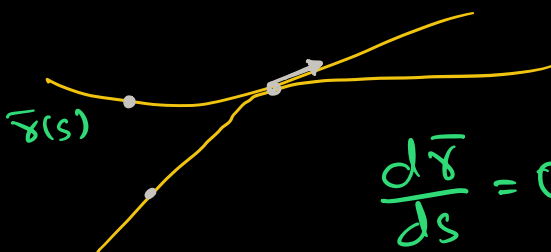
$$\psi = \text{stream function} = \frac{v^2}{2} + \frac{P}{\rho} + \Omega$$

$$\text{At steady state } 2\bar{v} \times \bar{\omega} = \nabla \psi$$

Stream Line: A curve at a given time such that its tangent at any point gives, the direction of  $\bar{v}$ .



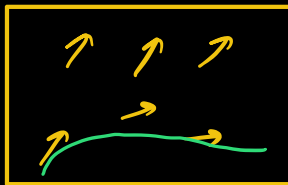
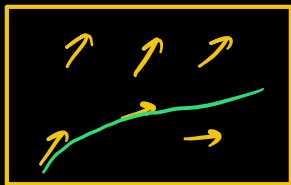
Streamlines can't cross at a point otherwise we will have two velocity for a particle ; which is unphysical.



$$\frac{d\bar{r}}{ds} = \alpha \cdot \bar{v}(\bar{r})$$

$$\left( \frac{d\bar{r}_s(s)}{ds} = \alpha \bar{v}(\bar{r}) \right)$$

$\bar{r}_s$  : s is for stream line.



Two possibility ??  
which one should we take.

check that the points on a streamline have same  $\alpha$ .  
 $\alpha$  characterizes the stream line.

Streamlines can change with time.

$$2\bar{v} \times \bar{\omega} = \bar{\nabla} \Psi, \quad \Psi = \frac{v^2}{2} + \frac{p}{\rho_0} + \Omega.$$

Along this stream line, how much  $\Psi$  change

$$d\Psi = \sum_i \frac{\partial \Psi}{\partial x_i} \cdot dx_i = \bar{\nabla} \Psi \cdot d\bar{r}$$

Now  $d\bar{r}$  is movement along streamline.

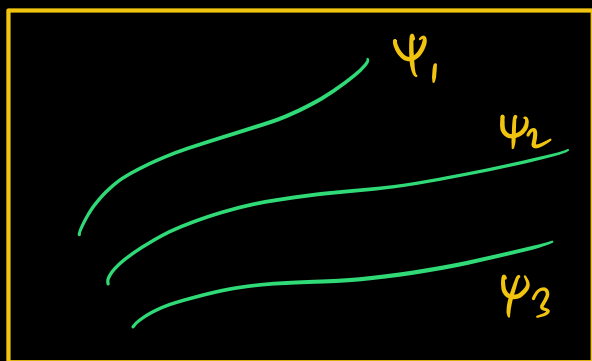
$$d\bar{r} = \alpha \bar{v}(\bar{r}) ds \quad \left( \frac{d\bar{r}}{ds} = \alpha \bar{v} \right)$$

$$\Rightarrow d\Psi = \bar{\nabla} \Psi \cdot (\alpha \cdot \bar{v}(\bar{r}) ds)$$

$$\Rightarrow d\Psi = (2\bar{v} \times \bar{\omega}) \cdot (\alpha \bar{v} ds) \\ = 0$$

$$\Rightarrow d\Psi = 0 \text{ along streamline.}$$

$\Psi$  does not change along streamline. Stream function  $\Psi$  remains constant on a Stream line.



## Bernoulli's Theorem

① Body Forces are conservative.

② Incompressible.

(because then  $-\frac{\nabla P}{\rho} = -\nabla\left(\frac{P}{\rho}\right)$ )

only for  $\rho = \rho_0$  constant

③ Ideal fluid:  $\begin{cases} \eta = 0 \\ \xi = 0 \end{cases}$

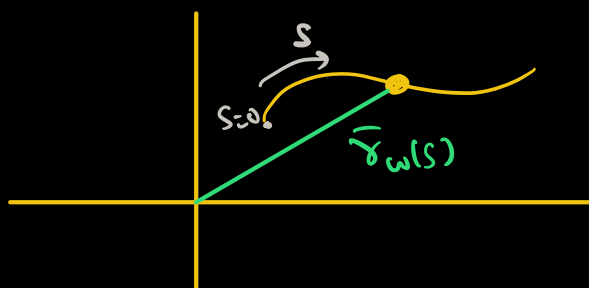
④ Steady flow  $\frac{\partial \mathbf{v}}{\partial t} = 0$

$\Rightarrow \psi$  is constant along stream lines.

Vortex Lines: tangent to this line gives direction of  $\bar{\omega} = \left(\frac{\nabla \times \bar{U}}{2}\right)$

$$\frac{d\bar{x}_\omega(s)}{ds} = \beta \bar{\omega}$$

$\bar{x}_\omega$  is position vector along the vortex line.



$$\psi = \frac{v^2}{2} + \frac{P}{\rho_0} + \Omega(\vec{r})$$

$$\frac{f}{\rho_0} = -\nabla \Omega$$

Force per unit mass.

$$\text{ex } f = -\rho_0 g$$

$$\text{so: } \Omega = gz$$

$$\Omega = gz$$

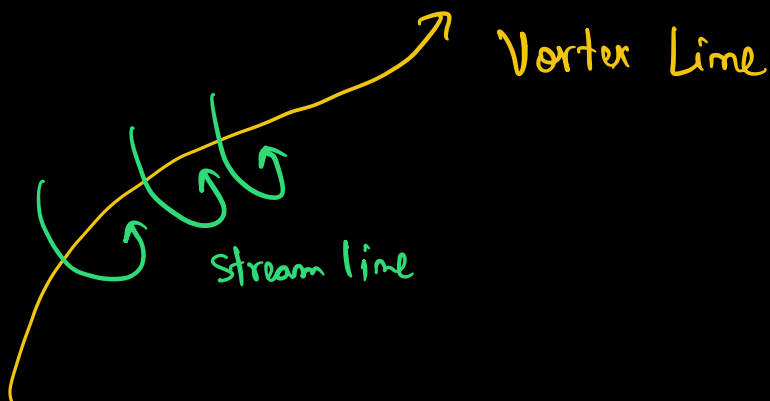
$$\begin{aligned}
 d\psi &= \bar{\nabla} \psi \cdot (\beta \bar{\omega} ds) \\
 &= (2\bar{v} \times \bar{\omega}) \cdot (\beta \bar{\omega} ds) \\
 &= 0
 \end{aligned}$$

: Change of  $\psi$  along the Vortex.

$d\psi = 0$  along vortex line.

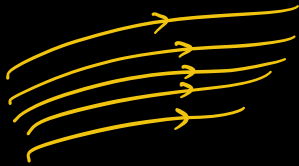
$\psi$  does not change along vortex line.

$\psi$  is constant along streamline & Vortex line.

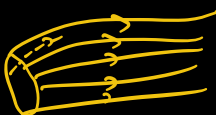


Stream lines go around the vortex line.

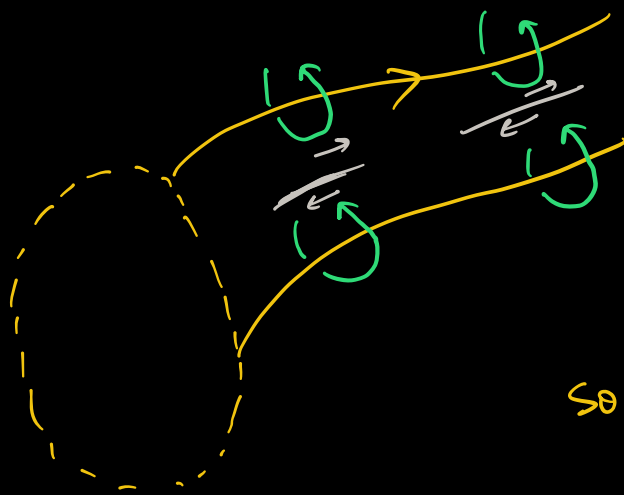
Vortex Sheet (when many vortex lines are parallel to each other, and form a sheet)



Vortex Tube (Vortex sheet curl up to form tube)

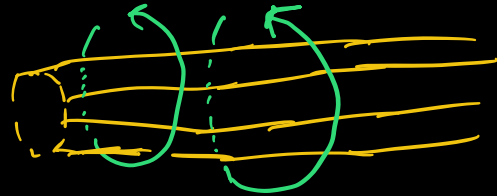






$\Rightarrow$  There transverse motion along the surface is zero.

So.



ie: The fluid is going round the vortex tube.

Stream Line :  $\frac{d\bar{x}_s(s)}{ds}$  —  $s$  stands for stream.

$$\frac{d\bar{x}_s(s)}{ds} = \alpha \bar{V}$$

$$\text{so: } \frac{dx_s}{ds} = \alpha v_x, \quad \frac{dy_s}{ds} = \alpha v_y, \quad \frac{dz_s}{ds} = \alpha v_z$$

so:

$$\boxed{\frac{dx_s}{v_x} = \frac{dy_s}{v_y} = \frac{dz_s}{v_z} = \alpha \cdot ds}$$

lets say  $(v_x, v_y, v_z) = \bar{V}(\bar{x})$  given to us

Now draw the stream lines. (at a given time)

$$v_x = \frac{x}{1+t}, \quad v_y = \frac{y}{1+2t}, \quad v_z = 0$$

$$\frac{dx_s}{ds} = \alpha v_x = \frac{\alpha x_s}{1+t}$$

$$\frac{dy_s}{ds} = \alpha v_y = \frac{\alpha y_s}{1+2t}$$

let  $\alpha = 1$        $\frac{dx_s}{ds} = \frac{x_s}{1+t}$  ,  $\frac{dy_s}{ds} = \frac{y_s}{1+2t}$

Note that time is fixed here

$$\int_{x_0}^{x_s} \frac{dx_s}{x_s} = \int_0^s \frac{ds}{1+t}$$

$$\Rightarrow \ln\left(\frac{x_s}{x_0}\right) = \frac{s}{1+t}$$

$$\text{Similarly } \ln\left(\frac{y_s}{y_0}\right) = \frac{s}{1+2t}$$

}  $\Rightarrow$  now eliminate  $s$   
& get equation  
between  $y_s$  &  $x_s$   
including a parameter  $t$ .

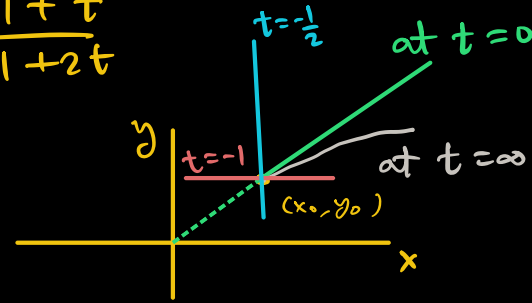
$$\Rightarrow \frac{y_s}{y_0} = \left(\frac{x_s}{x_0}\right)^n \quad \text{where } n = \frac{1+t}{1+2t}$$

alternate:  $\frac{dx_s}{dy_s} = \frac{x_s}{y_s} \cdot \frac{(1+2t)}{(1+t)} \Rightarrow \int_{x_0}^{x_s} \frac{dx_s}{x_s} = \frac{1}{n} \int_{y_0}^{y_s} \frac{dy_s}{y_s}$

$$\Rightarrow \ln\left(\frac{y_s}{y_0}\right) = n \cdot \ln\left(\frac{x_s}{x_0}\right) \Rightarrow \left(\frac{y_s}{y_0}\right) = \left(\frac{x_s}{x_0}\right)^n$$

So;  $\frac{y}{y_0} = \left(\frac{x}{x_0}\right)^n$  where  $n = \frac{1+t}{1+2t}$

at  $t=0 \Rightarrow n=1 \Rightarrow \frac{y}{y_0} = \frac{x}{x_0}$



at  $t=\infty, n=\frac{1}{2} \quad \frac{y}{y_0} = \left(\frac{x}{x_0}\right)^{1/2}$

at  $t=-1, n=0 \quad y=y_0$

at  $t=-\frac{1}{2}, n=\infty \quad x=x_0$

Path Line : actual path travelled by an individual fluid particle.



This is not  
streamline.

Tracking a particular  
particle.

The particle  
moves on path  
line as time passes.

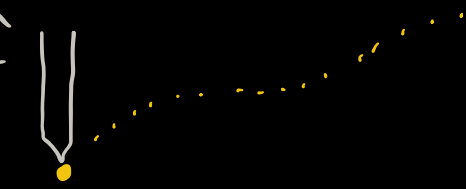
Streak line at a particle point  $(x_0, y_0)$

all the particles which pass through  $(x_0, y_0)$  before  $t$ ,  
where they are now.

Streak line is the present locus of particles which have passed through  $(x_0, y_0)$  at earlier times.

a color dye at a point.

The fluid parcel takes away the color and you can track them.



for steady flow

path line  $\equiv$  stream line

Chandrasekhar book (Astrophysical calculation)

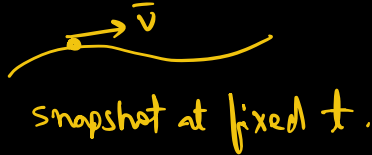
$\rightarrow$  some stability calculation.



Lee 14 (12<sup>th</sup> March, 2021 ; 3:30 pm to 5:00 pm)

Second Quiz : fluid mech only ; polar & spherical polar  
(specifically)

① Streamline

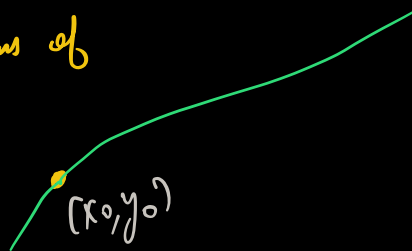


② Pathline (tag a particle & follow it in time)



③ Streak line

gives present locus of  
all the particles  
which pass  
through  
(x<sub>0</sub>, y<sub>0</sub>) at  
t < t.



all particles passed through (x<sub>0</sub>, y<sub>0</sub>)  
before t

Example Velocity given  $u = \dot{x}$  ,  $v = \dot{y}$   $\vec{v} = (u, v)$

Streamline  $\vec{r}(s)$   
(a space curve)

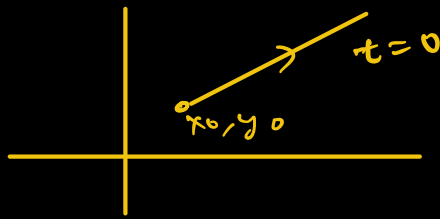
$$\frac{dx}{ds} = u \quad , \quad \frac{dy}{ds} = v$$

$$u = \frac{x}{1+t} \quad v = \frac{y}{1+2t} \quad \text{given.}$$

$$x = c_1 \cdot e^{s/(1+t)} \quad , \quad y = c_2 \cdot e^{s/(1+2t)}$$

$$\Rightarrow y = y_0 \cdot \left(\frac{x}{x_0}\right)^m \quad \text{where } m = \frac{1+t}{1+2t}$$

at  $t=0$ ,



path line  $\frac{dx}{dt} = u, \quad \dot{y} = v \quad \Rightarrow \quad \frac{dx}{dt} = \frac{x}{1+t}$

$$\Rightarrow \ln\left(\frac{x}{c_1}\right) = \ln(1+t)$$

at  $t=0 \Rightarrow x = c_1$

$$\Rightarrow \boxed{x = c_1 \cdot (1+t)}$$

but  $x(t=0) = x_0 \Rightarrow c_1 = x_0$

$$\Rightarrow \boxed{x = x_0 \cdot (1+t)}$$

$$\frac{dy}{dt} = \frac{y}{1+2t} \Rightarrow \frac{dy}{y} = \frac{dt}{1+2t} = \frac{dz}{z} \left( \frac{1}{1+z} \right) \quad z = 2t$$

$$\Rightarrow \ln\left(\frac{y}{c_2}\right) = \frac{1}{2} \ln(1+z)$$

$$\Rightarrow \frac{y}{c_2} = (1+2t)^{1/2}$$

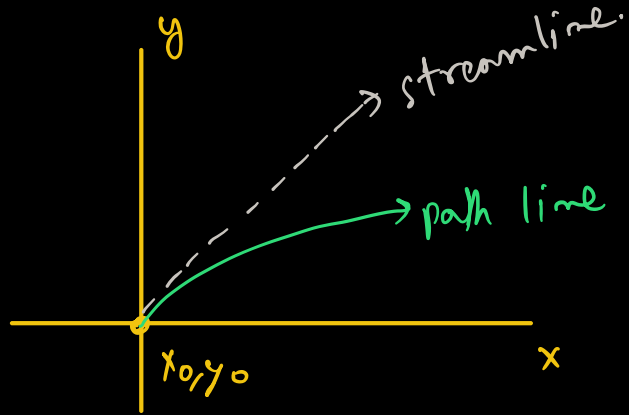
$t=0, y=y_0 \Rightarrow c_2 = y_0$

$$\Rightarrow \boxed{y = y_0 \cdot (1+2t)^{1/2}}$$

eliminate  $t$  & get trajectory  $f(x, y) = 0$

$$\Rightarrow y = y_0 \cdot \left[ 1 + 2 \left( \frac{x}{x_0} - 1 \right) \right]^{1/2}$$

at  $(t=0, (x, y) = (x_0, y_0))$



Streak line

$$x = c_1(1+t), \quad y = c_2 \cdot (1+2t)^{1/2}$$

All those particles which passed through  $(x_0, y_0)$  at  $\tau < t$

$$x_0 = c_1(1+\tau), \quad y_0 = c_2(1+2\tau)^{1/2}$$

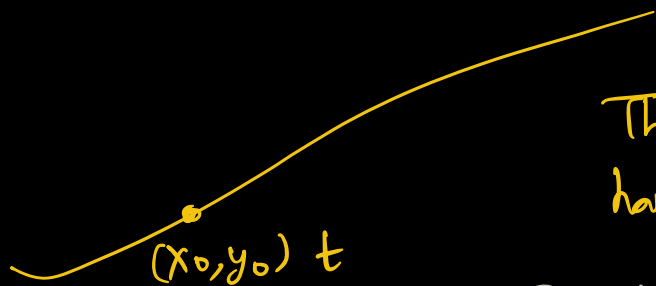
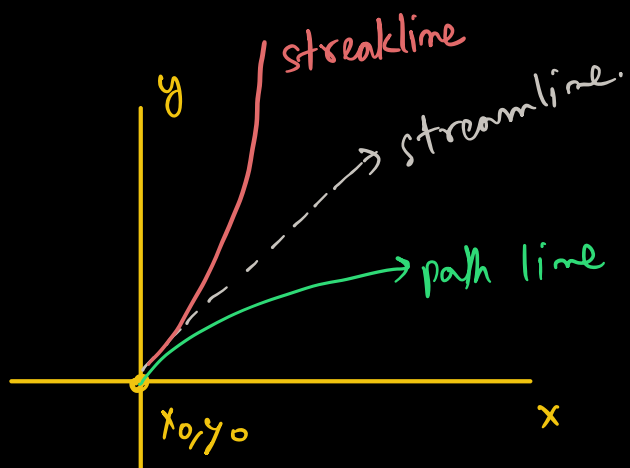
$$c_1 = \frac{x_0}{1+\tau}, \quad c_2 = \frac{y_0}{(1+2\tau)^{1/2}}$$

$$\boxed{x = \frac{x_0}{1+\tau} \cdot (1+t)} \quad \Rightarrow \quad \tau = \frac{x_0}{x} \cdot (1+t) - 1$$

$$y = \frac{y_0}{(1+2\tau)^{1/2}} (1+2t)^{1/2} \quad \Rightarrow \quad \boxed{y = \frac{y_0}{(1+2\tau)^{1/2}} \cdot (1+2t)^{1/2}}$$

$$\Rightarrow \boxed{\left( \frac{y_0}{y} \right)^2 = \frac{1+2t}{1+2(t+1)\left(\frac{x_0}{x}\right) - 1}}$$

at  $t=0$   $\frac{y}{y_0} = \frac{1}{[1 + 2(\frac{x_0}{x} - 1)]^{1/2}}$  } streamline



This means, that all these particles have passed  $(x_0, y_0)$  before  $t$ .

Streak line contains information about the particle which crossed  $(x_0, y_0)$  before  $t$ .

Whereas stream line tells the future of particle which crossed  $(x_0, y_0)$  at  $t$ .

Time Reversal Symmetry is broken in Navier's Stokes Equation.

$$\rho \left[ \frac{\partial \bar{v}}{\partial t} + (\bar{v} \cdot \nabla) \bar{v} \right] = -\nabla \bar{p} + \eta \nabla^2 \bar{v}$$

Breaks  $t \rightarrow -t$  symmetry

because  $\bar{v} \rightarrow -\bar{v}$  and so  $\frac{\partial \bar{v}}{\partial t} \rightarrow \frac{\partial \bar{v}}{\partial t}$



$$(\bar{\mathbf{v}} \cdot \bar{\nabla}) \bar{\mathbf{v}} \rightarrow (\bar{\mathbf{v}} \cdot \bar{\nabla}) \bar{\mathbf{v}}$$

$$\text{but } \nabla^2 \bar{\mathbf{v}} \rightarrow -\nabla^2 \mathbf{v}$$

so, LHS =  $\rho \left[ \frac{\partial \bar{\mathbf{v}}}{\partial t} + (\bar{\mathbf{v}} \cdot \bar{\nabla}) \bar{\mathbf{v}} \right]$  is time reversal invariant

$$\eta \nabla^2 \mathbf{v} \rightarrow -\eta \nabla^2 \mathbf{v}$$

so, on RHS  $\eta \nabla^2 \mathbf{v}$  term breaks the Time Reversal Symmetry.

$$\bar{\nabla} p \rightarrow \bar{\nabla} p$$

The viscosity is not symmetric under  $t \rightarrow -t$ .

and this causes the Navier's Stokes equation to break time reversal symmetry.

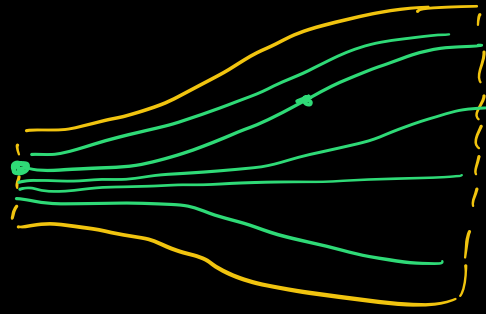
solve it.

$$\begin{aligned} p_x &= -\frac{\rho_0 x^2}{2(1+t)^2} \\ p_y &= -\frac{\rho_0 y^2}{2(1+2t)^2} \end{aligned}$$

$$\frac{\partial v_x}{\partial t} = \frac{-x}{(1+t)^2}$$

$$\frac{\partial v_y}{\partial t} = \frac{-y}{(1+2t)^2} x^2$$

For time independent  
flows,  
streamline &  
pathline are same.  
& streakline

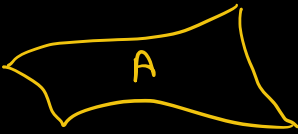


For time independent flows.

and streakline & pathline grows with time and  
follow the streamline.

Streakline & Pathline grows with time and will follow  
streamline.





$$E = \sigma A$$

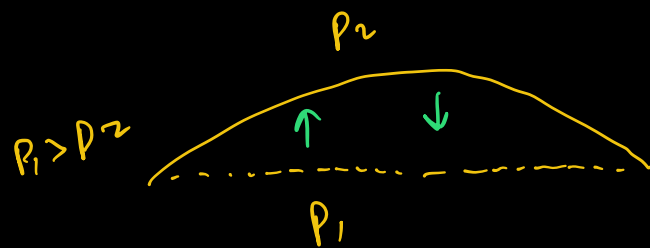
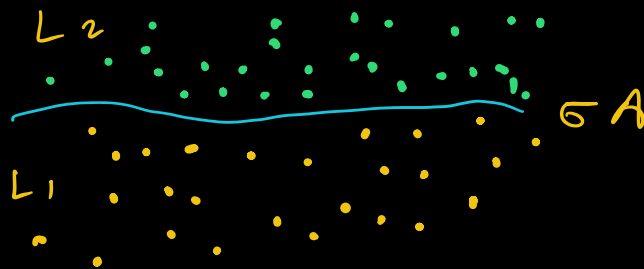
area

energy

Surface tension

surface stretching energy  $\frac{1}{2} K (A - A_0)^2$

Interface

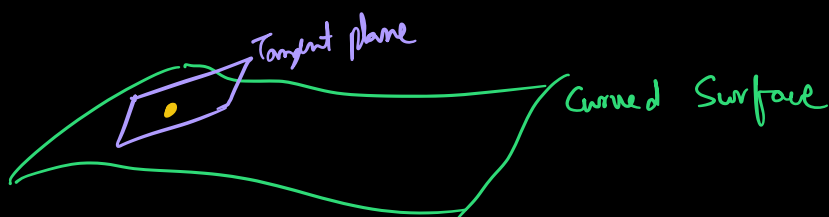


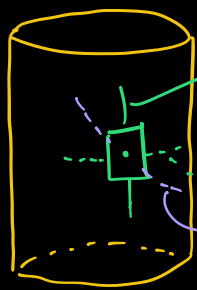
curved interface

(has more energy, due to more area locally)

$$\Delta P = \sigma \cdot \left( \frac{1}{R_1} + \frac{1}{R_2} \right)$$

Principal Curvatures





radius of curvature along the flat direction  $R_1 = \infty$

but  $R_2 = R$  finite.

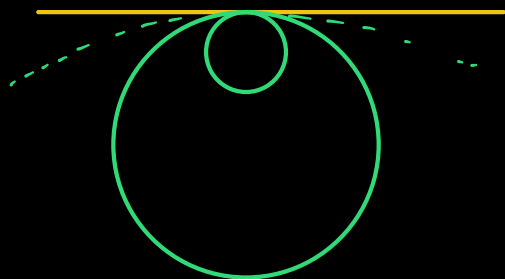
curvature between  $R_1$  &  $R_2$

at a point we can draw many circles.

Principle Curvature: The maximum and minimum radii of the circles that we can draw through a point tangent to the local tangent plane.

$$R_1 = \max, \quad R_2 = \min$$

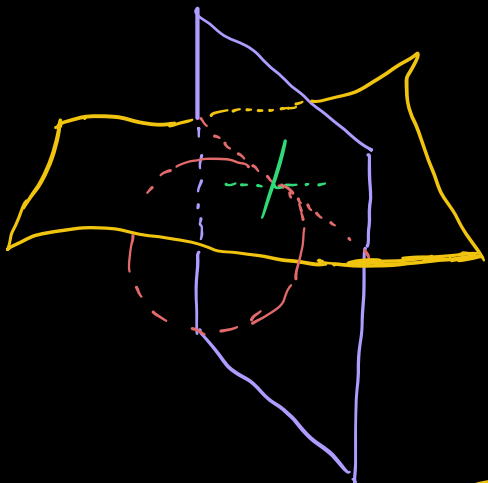
ex



straight line.

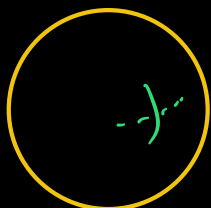
$$R_1 = \infty, \quad R_2 = 0$$

ex



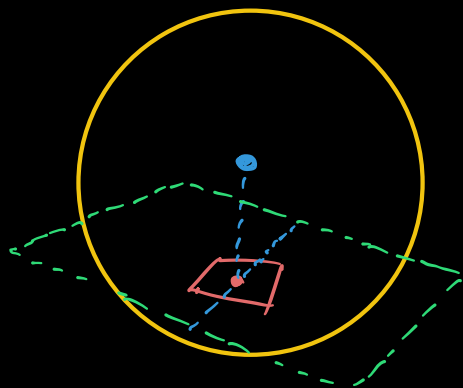
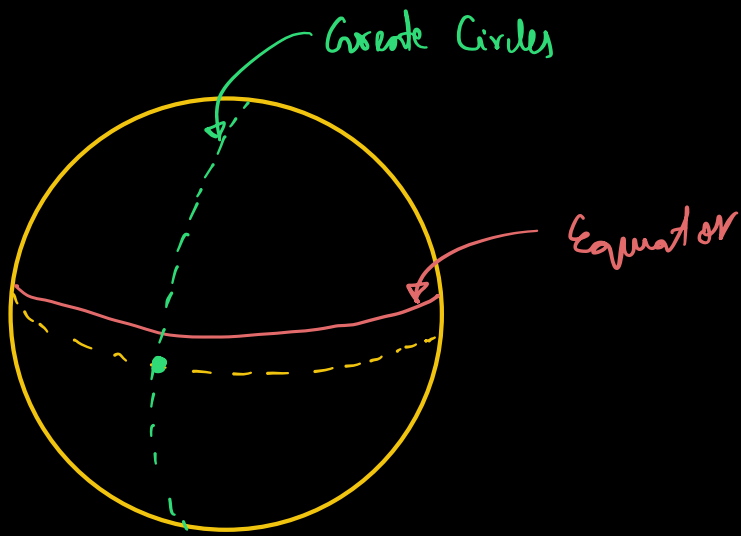
example 1 sphere

$R$

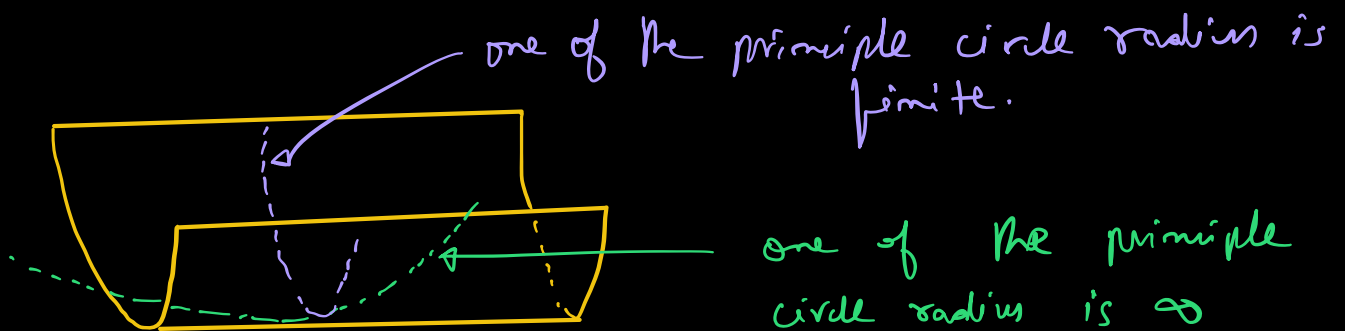


both circles have curvature  $1/R$

both principle curvatures are same



ex



$$\Delta P = \sigma \left( \frac{1}{R_1} + \frac{1}{R_2} \right) = \frac{2\sigma}{R}$$

↑  
Laplace  
Pressure  
(comes at interface)

for spherical interface  
(locally the area will be  
a path of a sphere)

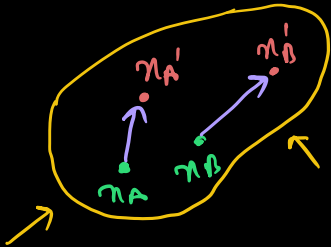
The pressure needed to prevent curvature of the  
surface.



## ELASTICITY



when released, it comes back.

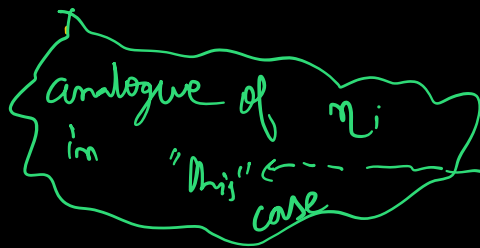


$$x_A \rightarrow x_{A'}$$

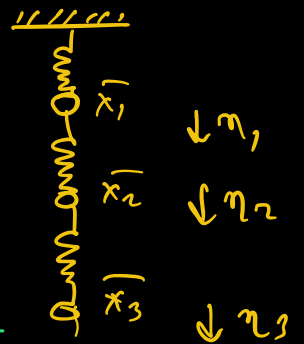
$$x_B \rightarrow x_{B'}$$

displacement field

$$\bar{u} = x' - x$$



flex



$$\begin{aligned} (x'_A - x'_B) &= (x_A + u_A) - (x_B + u_B) \\ &= (x_A - x_B) + (u_A - u_B) \end{aligned}$$

Continuum :  $x_A$  &  $x_B$  are close by.

then,  $x_{A'} - x_{B'} = dx'$

$$x_A - x_B = dx$$

and  $u_A - u_B = du$

$$\Rightarrow dx' = dx + du$$

$dx'_i = dx_i + du_i$

 putting indices for vector components.

$$du_i = \frac{\partial u_i}{\partial x_j} \cdot dx_j$$

$$\Rightarrow dx_i' = dx_i + \frac{\partial u_i}{\partial x_j} dx_j$$

$$dx_i' = \left( \delta_{ij} + \frac{\partial u_i}{\partial x_j} \right) dx_j$$

Assume:  $u_i(\bar{x})$  is a smooth function.

Change in distance:

How does the distance change;  $dx_i'^2 = ?$

$$dx_i' = \left( \delta_{ij} + \frac{\partial u_i}{\partial x_j} \right) dx_j$$

$$\frac{\partial u_i}{\partial x_j} \equiv U_{i,j} \text{ or } U_{ij}$$

Now we write the line element.

$$dx_i'^2 = \left( \delta_{ij} + \frac{\partial u_i}{\partial x_j} \right) dx_j \cdot \left( \delta_{in} + \frac{\partial u_i}{\partial x_n} \right) dx_n$$

$$= \underbrace{(\delta_{ij} dx_j \delta_{in} dx_n)}_{dx_i^2} + \delta_{in} \frac{\partial u_i}{\partial x_j} dx_j dx_n$$

$$+ \delta_{ij} \cdot \frac{\partial u_i}{\partial x_n} \cdot dx_j \cdot dx_n$$

$$+ O\left(\left(\frac{\partial u_i}{\partial x_j}\right)^2\right)$$

if  $u$  is smooth, then its derivative is small number.

Assumption that;  $u$  is small field.

$\Rightarrow$  so its gradient is small.

$$dx_i'^2 = dx_i^2 + \frac{\partial v_u}{\partial x_j} dx_u dx_j + \frac{\partial v_j}{\partial x_u} dx_j dx_u$$

$$\frac{\partial v_u}{\partial x_j} dx_u dx_j$$

(renaming dummy index)

$$dx_i'^2 = dx_i^2 + 2 dx_i dx_j \cdot \frac{\partial v_i}{\partial x_j}$$

$$2 \cdot \frac{\partial v_i}{\partial x_j} = \left( \frac{\partial v_i}{\partial x_j} + \frac{\partial v_j}{\partial x_i} \right) + \left( \frac{\partial v_i}{\partial x_j} - \frac{\partial v_j}{\partial x_i} \right)$$

$$= \underbrace{S_{ij}}_{\text{symmetric}} + \underbrace{A_{ij}}_{\text{antisymmetric.}}$$

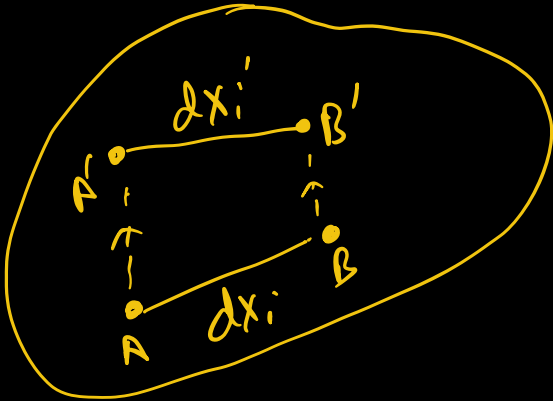
$$\begin{aligned} dx_i'^2 &= dx_i^2 + dx_i dx_j (S_{ij} + A_{ij}) \\ &= dx_i^2 + dx_i dx_j S_{ij} + \cancel{dx_i dx_j A_{ij}} \end{aligned} \rightarrow 0$$

$$dx_i'^2 = dx_i^2 + \left( \frac{\partial v_i}{\partial x_j} + \frac{\partial v_j}{\partial x_i} \right) dx_i dx_j$$



NOTATION :  $\frac{1}{2} \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) \equiv e_{ij}$  STRAIN TENSOR

$$dx_i'^2 = dx_i^2 + 2 dx_i dx_j \cdot e_{ij}$$



$$dx_i'^2 = dx_i^2 + 2 e_{ij} dx_i dx_j$$

$$e_{ij} = \frac{1}{2} \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right)$$

$$= \frac{1}{2} \left( \frac{\partial u_i}{\partial x_j} + i \leftrightarrow j \right)$$

Strain:  $\frac{dl}{l}$

Go to a frame, in which  $e_{ij}$  is diagonal.  
 (so, will need to choose different frames at different point to make it diagonal)

here sum over  $i$ .

$$d\tilde{x}_i'^2 = d\tilde{x}_i^2 + 2 \cdot \tilde{e}_{ii} d\tilde{x}_i^2$$

$\rightarrow$   $i^{\text{th}}$  element; no sum over  $i$

$$e_{ij} \rightarrow \delta_{ij} \tilde{e}_{ij} \quad (\text{in diagonal frame})$$

(no sum over  $i$ )

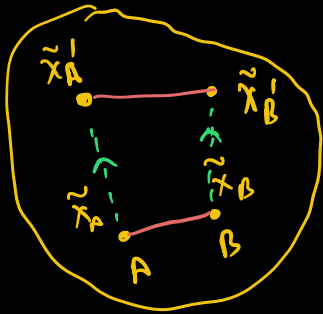
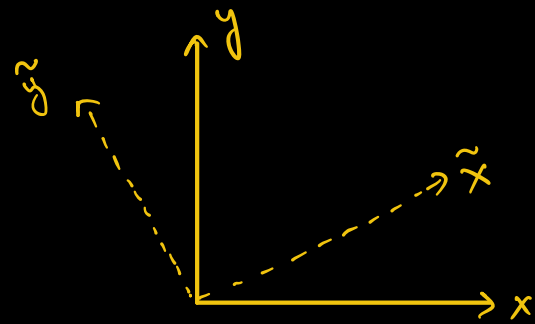
$$\tilde{e} = \begin{bmatrix} \tilde{e}_{11} & 0 & 0 \\ 0 & \tilde{e}_{22} & 0 \\ 0 & 0 & \tilde{e}_{33} \end{bmatrix}$$

$$d\tilde{x}'^2 = d\tilde{x}^2 (1 + 2\underbrace{\tilde{e}_{ii}}_{\rightarrow \text{no sum over } i})$$

Because  $e_{ij}$  was real symmetric,

it can always be diagonalized by a rotation;  
and also all the eigenvalues are real & eigenvectors are  
orthogonal.

$d\tilde{x}_i$  is along that orthogonal eigenvectors.



$\Downarrow$

$$\begin{aligned} d\tilde{x}'_i &= d\tilde{x}_i \cdot \sqrt{1 + 2\tilde{e}_{ii}} \\ &\approx d\tilde{x}_i \cdot (1 + \tilde{e}_{ii}) \end{aligned}$$



$$d\tilde{x}_i' = d\tilde{x}_i \cdot (1 + \tilde{e}_{ii})$$

change in volume  $dx dy dz \rightarrow dx' dy' dz'$   
omitting the tilde.

$$dx' dy' dz' = dx(1 + \tilde{e}_{11}) \cdot dy(1 + \tilde{e}_{22}) dz(1 + \tilde{e}_{33})$$

$$1, 2, 3 \equiv x, y, z$$

$$e_{ij} \rightarrow \tilde{e}_{ij}$$

↓

$$\frac{1}{2} \left( \frac{\partial u_i}{\partial x_j} + i \leftrightarrow j \right)$$

$u_i$  are small,  $\frac{\partial u_i}{\partial x_j}$  even smaller

$$\Rightarrow dx' dy' dz' \approx dx dy dz (1 + e_{11} + e_{22} + e_{33}) + O(e^2)$$

$$dV' = dV \cdot (1 + \text{Tr}(\tilde{e}))$$

$\text{Tr}(e)$  does not change under rotation of frame.

$$\text{Tr}[e_{ij}] = \text{Tr}[\tilde{e}_{ij}]$$

so,

$$dV' = dV \cdot (1 + \text{Tr}(e))$$

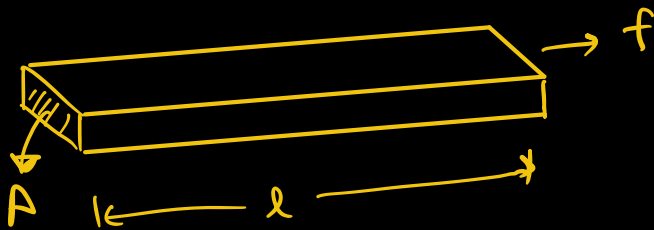
$\text{Tr}(\epsilon)$  is fractional volume change

$$\text{Tr}(\epsilon) = \frac{dV' - dV}{dV}$$

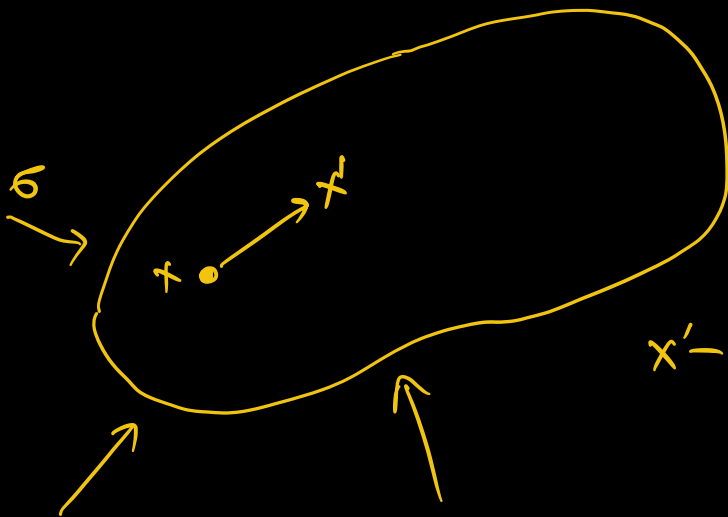
---

Connecting Stress to Strain  
(Constitutive Relation)

$$\sigma_{ij} = f(\epsilon_{ij})$$



$$\sigma = \frac{f}{A} = \gamma \cdot \frac{A l}{l}$$



$$x' - x = u(x) \text{ displacement}$$

$$\epsilon_{ij} = \frac{1}{2} \left( \frac{\partial u_i}{\partial x_j} + i \leftrightarrow j \right)$$

Strain is effect.

Stress is cause.

(assume linear response)

Stress  $\longleftrightarrow$  Strain ?  
 $\sigma_{ij}(\bar{x})$   $\epsilon_{ij}(\bar{x})$

$\epsilon_{ij}$  is symmetrized  
form of  $\frac{\partial u_i}{\partial x_j}$

$$\sigma_{ij}(\bar{x}) = 2\mu \epsilon_{ij}(\bar{x}) + \lambda \cdot \delta_{ij} \epsilon_{kk}$$

Isotropic

Assuming linearity in  
simple way.

the  $2\mu \cdot \epsilon_{ij}(\bar{x})$  term  
(we could have  $A_{ijkl} \epsilon_{kl}$ )

$\mu$  and  $\lambda$  are Lamé Constants.

- ①  $\sigma_{ij}$  is a linear function of  $\epsilon_{ij}$  (because  $\epsilon_{ij}$  is small)
- ②  $\sigma_{ij}$  should not have constant on RHS.

$$\text{if } \sigma_{ij} = f_{ij}(\epsilon) + C$$

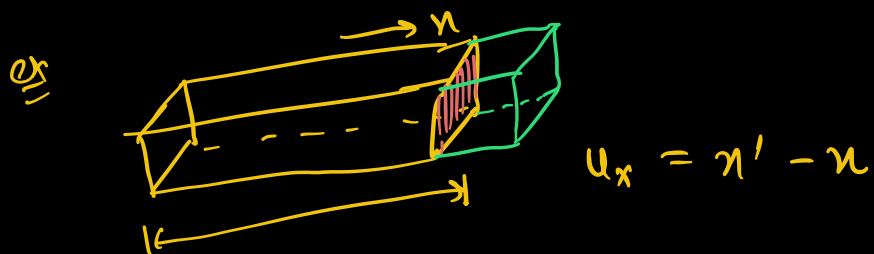
$$\text{because if } \sigma_{ij} = 0 \Rightarrow f_{ij}(\epsilon) = -C$$

$$\text{i.e., } \epsilon_{ij} \neq 0$$

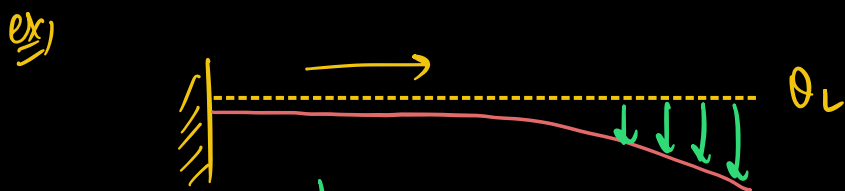
- ③ can't have  $u_i v_j$  on the RHS, since  $u_i = \text{constant}$   
i.e., uniform translation  
we will have stress.

so we have to always use their  
gradient.

although  $\partial u_i$  is small,  
but  $u_i u_j$  can be large



$u_L$  is large, but  $\frac{\partial u}{\partial x}$  is still small.



bending due to gravity.

$$\frac{d\theta}{dx} \ll 1$$

, but  $\theta_L$  may not be small.

$\sigma_{ij}$  must be symmetric for angular momentum conservation  
(when there is no external torque)

→ This is why we don't just add  $\frac{\partial u_i}{\partial x_j}$  in RHS of

$\sigma_{ij}$ ; but  $\left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right)$ ; so we indeed

use  $e_{ij}$

ex 11 For a Bar

A B



A' B'

$$x_{A'} - x_{B'} = (x_A - x_B) + (u_A - u_B)$$

$$\text{ie, } dx' = dx + du$$

$$\text{ie, } l' = l + du$$

$$du = \Delta l$$

$$\Rightarrow \frac{\Delta u}{l} = \frac{du}{l} = \frac{du}{dx} \quad \text{Fractional change of length.}$$

$$\boxed{\frac{F}{A} = \sigma = \gamma \cdot \frac{\Delta l}{l}} \quad \text{found before.}$$

$$\hookrightarrow \sigma_{ij} = \gamma \cdot \left( \frac{\partial u_i}{\partial x_j} + i \leftrightarrow j \right)$$

$$\sigma_{ij} = 2\mu \cdot e_{ij} + \lambda \cdot \delta_{ij} \cdot e_{kk}$$

$$\text{Tr}(e) = e_{kk} = \frac{\partial u_k}{\partial x_k}$$

$$\sigma_{ij} = 2\mu e_{ij} + \lambda \delta_{ij} \cdot \left( \frac{\partial u_x}{\partial x} + \frac{\partial u_y}{\partial y} + \frac{\partial u_z}{\partial z} \right)$$

$$dV' = dV \cdot (1 + e_{kk})$$

$$\lambda \equiv e_{kk} = \frac{dV' - dV}{dV}$$

$$e_{ij} = \left( e_{ij} - \frac{1}{3} \cdot \Delta \cdot \delta_{ij} \right) + \frac{1}{3} \cdot \Delta \cdot \delta_{ij}$$

has  $\text{Tr}() = \Delta$

has  $\text{Tr}() = 0$

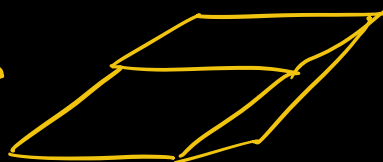
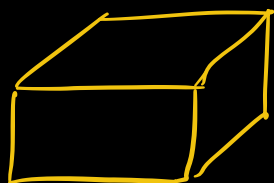
has  $\text{Tr}() = \Delta$

$$\boxed{\text{Tr}(e_{ij}) = \Delta}$$

Traceless part of Strain Tensor.

This does not lead to volume change;  
but it changes the shape of the object.

ex 3D



Shape change; but not volume change due to traceless part of strain.

so,  $\frac{1}{3} \Delta \cdot \delta_{ij}$  leads to volume change.

$$\sigma_{ij} = 2\tilde{\mu} \cdot \left( e_{ij} - \frac{\Delta}{3} \cdot \delta_{ij} \right) + K \cdot \delta_{ij} \cdot \Delta$$

Shear Modulus

Bulk Modulus.

2D





Show:  $\tilde{\mu} = \mu$ , but  $\lambda = -\frac{2}{3}\mu + K$

$\Delta \cdot K \cdot \delta_{ij}$  leads to no shape change, but only volume change.

Imagine pure strain of type  $e_{ij} = \frac{1}{3} \cdot \delta_{ij} \cdot \Delta$

$$\Rightarrow e_{ii} = \frac{\Delta}{3} \quad (\text{no sum over } i)$$

$$\& e_{ij} = 0 \quad \text{for } i \neq j$$

$$\text{so, } \frac{\partial u_x}{\partial x} = \frac{\partial u_y}{\partial y} = \frac{\partial u_z}{\partial z} = \frac{\Delta}{3} = c$$

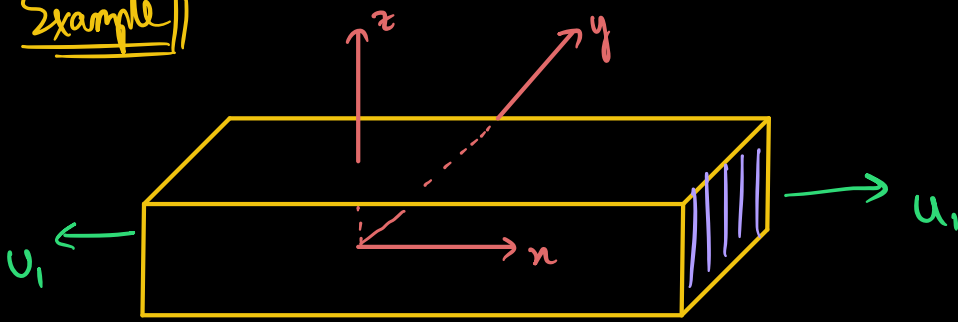
$$\Rightarrow \left. \begin{array}{l} u_x = c \cdot x \\ u_y = c \cdot y \\ u_z = c \cdot z \end{array} \right\} \Rightarrow \text{so, we have just scale transformation in all direction, by the scale factor } c.$$

any possible constants like

$$u_x = c \cdot x + c_1$$

$c_1$  in this can be gotten rid of  
(this just means all points shift in  $x$  direction (so no shape change for sure); just pure translation) by shifting  $x$ -origin.

### Example 11



put force in x direction.

displacement field.

$$u_1 = a x_1$$

$$u_2 = b x_2$$

$$u_3 = b x_3$$

} y & z directions are similar.  
(Assume symmetry in y-z)

Strain Field: 
$$e_{ij} = \frac{1}{2} \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right)$$

$$[e_{ij}] = \begin{bmatrix} a & 0 & 0 \\ 0 & b & 0 \\ 0 & 0 & b \end{bmatrix}$$

$$\sigma_{ij} = 2\mu \cdot \left( e_{ij} - \delta_{ij} \cdot \frac{\Delta}{3} \right) + K \cdot \Delta \cdot \delta_{ij}$$

$$\Delta = a + b + b = a + 2b$$

↓  
Diagonal

$$[\sigma_{ij}] = 2\mu \cdot \begin{pmatrix} a - \frac{(a+2b)}{3} & 0 & 0 \\ 0 & b - \frac{(a+2b)}{3} & 0 \\ 0 & 0 & b - \frac{(a+2b)}{3} \end{pmatrix} + K \cdot \begin{pmatrix} a+2b & 0 & 0 \\ 0 & a+2b & 0 \\ 0 & 0 & a+2b \end{pmatrix}$$

$$\sigma_{11} = K \cdot (a+2b) + \frac{4\mu}{3} \cdot (a-b)$$

$$\sigma_{22} = K \cdot (a+2b) - \frac{2\mu}{3} \cdot (a-b) = \sigma_{33}$$

No stress on  $y$  and  $z$  surfaces.  $\Rightarrow \sigma_{22} = 0, \sigma_{33} = 0$   
 ( $\hat{y}$ ) ( $\hat{z}$ )

Since  $\sigma_{22}$  or  $\sigma_{33}$  does not depend on  $x$

and we satisfy boundary condition

$$\Rightarrow \sigma_{22} = 0 \text{ \& } \sigma_{33} = 0$$

$$\Rightarrow K = \frac{2\mu}{3} \cdot \frac{(a-b)}{(a+2b)}$$

$$\Rightarrow \frac{b}{a} = \frac{1}{2} \cdot \left( \frac{2\mu - 3K}{\mu + 3K} \right) = -\sigma$$

minus; because we want to keep  $\sigma$  positive for normal material.

$\sigma$  is called Poisson Ratio.

if  $3K > 2\mu$  then  $b < 0 \Rightarrow$  shrinks along  $y$  and  $z$  direction.  
 (if expand along  $x$ )

If volume has to be conserved, i.e. incompressible

we need  $K = \infty$  ( $K$  is bulk Modulus)

$$\frac{\mu}{K} \rightarrow 0$$

$$\Rightarrow \frac{b}{a} = -\frac{1}{2}$$

for large  $K$ , the cost of expansion or contraction is large.

( $K$  is small for compressible)

$$\Rightarrow \sigma = 1/2$$

Usually; when material expand in one direction, it gets shrinked in other.



If shear modulus  $= \mu \rightarrow \infty \Rightarrow \frac{b}{a} = 1 \Rightarrow \sigma = -1$

so range for  $\sigma \in [-1, 1/2]$

Regular materials have  $\sigma > 0 \Rightarrow$  shrinks along  $y, z$  if pulled along  $x$ .

Can calculate  $\sigma_{||}$

$$\sigma_{||} = K(a + 2b) + \frac{4\mu}{3} \cdot (a - b)$$

using relation  $\frac{b}{a} = \frac{1}{2} \cdot \frac{(2\mu - 3K)}{(\mu + 3K)}$

$$\Rightarrow \sigma_{||} = a \cdot \left( \frac{9\mu \cdot K}{3K + \mu} \right)$$

stress along  
x direction

strain

$$\Rightarrow \frac{9\mu K}{3K + \mu} \equiv Y$$

Young's Modulus

for simple linear elasticity.

## Volume Change

$$\Delta = \text{Tr}(e_{ij}) = a + 2b = \begin{cases} 0 & \text{only for } K \rightarrow \infty \\ a + 2b \neq 0 & \text{in general.} \end{cases}$$

$$\begin{aligned} a + 2b &= a \cdot \left( 1 + \frac{2\mu - 3K}{\mu + 3K} \right) \\ &= a \cdot \left( \frac{\mu + 3K + 2\mu - 3K}{\mu + 3K} \right) \\ &= a \cdot \left( \frac{3\mu}{\mu + 3K} \right) \end{aligned}$$

$$a + 2b = a \cdot \frac{3\mu}{\mu + 3K}$$

$$= \frac{\bar{\sigma}_{11}}{3K} \quad (\text{volume change})$$

When incompressible, we have  $K \rightarrow \infty$ .

In fluids

$$\rho \left[ \frac{\partial \bar{v}}{\partial t} + (\bar{v} \cdot \nabla) \bar{v} \right] = -\nabla p + \eta \nabla^2 \bar{v} + \xi \nabla(\nabla \cdot \bar{v}) + \bar{f}_{\text{ext}}$$

Shear viscosity  $\eta$  (for incompressible fluid,  $\xi \rightarrow \infty$ )

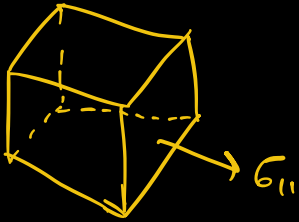
(for incompressible fluid,  $\xi \rightarrow \infty$ )  
 $\hookrightarrow$  so system chooses to have  $\nabla \cdot \bar{v} = 0$ , otherwise equation becomes singular

for incompressible, this resistance is  $\infty$ .  
 so:  $\xi = \infty$ .

(resist compression or Bulk viscosity expansion)



## Shear Stress



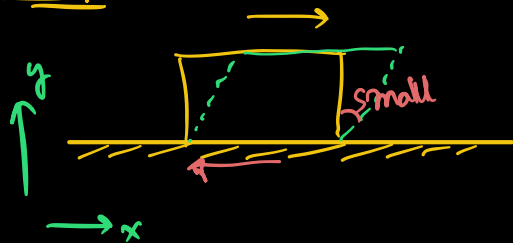
$$\sigma_{11}, \sigma_{22}, \sigma_{33} \neq 0$$

$$\sigma_{i \neq j} = 0 \quad (\text{no shear})$$

$$a + 2b = \frac{\sigma_{11}}{3K} = \text{fractional volume change} = \Delta = \frac{\delta V}{V}$$

$$a + 2b = \frac{\sigma_{11}}{3K} \xrightarrow{k \rightarrow \infty} 0$$

Example



$$u_x = \gamma \cdot y$$

$$u_y = 0$$

assume no volume (area) change,

$$\text{ie; } \epsilon = \frac{1}{2}$$

$$e_{ij} = \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) \cdot \frac{1}{2}$$

$$[e_{ij}] = \begin{bmatrix} 0 & \gamma/2 \\ \gamma/2 & 0 \end{bmatrix} = \begin{bmatrix} \epsilon_{xx} & \epsilon_{xy} \\ \epsilon_{yx} & \epsilon_{yy} \end{bmatrix}$$

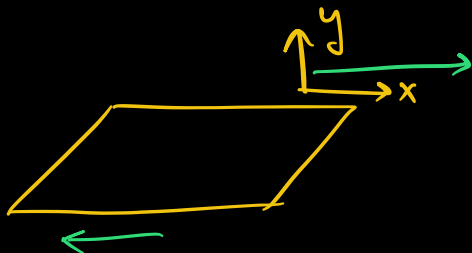
Strain Matrix

$$\sigma_{ij} = 2\mu \cdot (e_{ij} - \frac{\delta_{ij}}{3} \cdot \Delta) + K \cdot \Delta \cdot \delta_{ij}$$

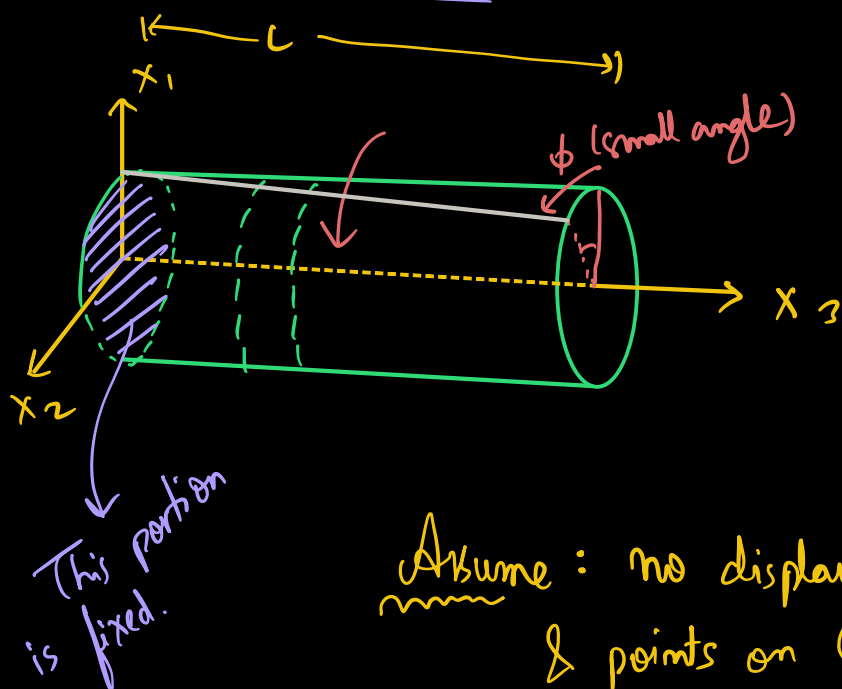
here  $\Delta = 0$

$$\Rightarrow \sigma_{ij} = 2\mu \cdot \epsilon_{ij}$$

$$\Rightarrow [\sigma] = \begin{bmatrix} 0 & \mu \cdot \gamma \\ \mu \cdot \gamma & 0 \end{bmatrix}$$



### Torsion in a Beam

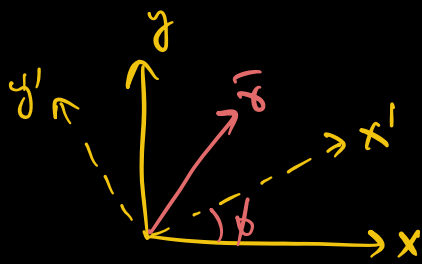


There is shear force between every layer.

Assume: no displacement along  $\hat{x}_3$ .  
 & points on Circular Cross section rotates by small angle  $\phi$ .

$$\begin{pmatrix} x_1 \\ x_2 \\ x_3 \end{pmatrix} \xrightarrow{\text{after rotating}} \begin{pmatrix} x'_1 \\ x'_2 \\ x'_3 \end{pmatrix} = \begin{pmatrix} \cos\phi & -\sin\phi & 0 \\ \sin\phi & \cos\phi & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} x_1 \\ x_2 \\ x_3 \end{pmatrix}$$

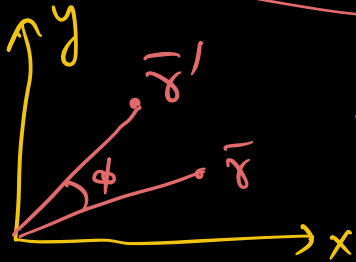
we are moving vector ;  
 so active rotation.  
 (not rotating the frame)



Passive Rotation (because the frame rotated; not the vector)

$$\vec{r} \rightarrow \vec{r}' = \begin{pmatrix} \cos \phi & \sin \phi \\ -\sin \phi & \cos \phi \end{pmatrix} \vec{r}$$

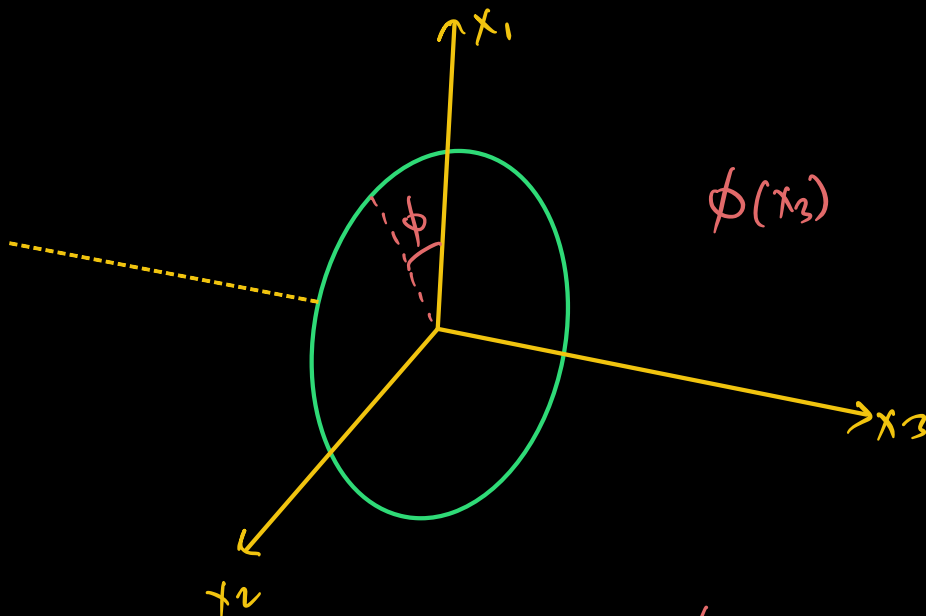
$\vec{r}' = \vec{r}$  equal as vector. equation among components.



Active Rotation (the vector rotates)

$\vec{r}' \neq \vec{r}$  not equal as vector.

$$\vec{r}' = \begin{pmatrix} \cos \phi & -\sin \phi \\ \sin \phi & \cos \phi \end{pmatrix} \vec{r}$$



Assumption  $\phi \propto x_3$

ie,  $\phi = \alpha \cdot x_3$

some elastic modulus.

(not Bulk Modulus K)

$$\sigma_{ij} = K \cdot \Delta \cdot \delta_{ij} + \left( \sigma_{ij} - \frac{\Delta \cdot \delta_{ij}}{3} \right) 2\mu$$

$$\begin{pmatrix} x'_1 \\ x'_2 \\ x'_3 \end{pmatrix} = \begin{pmatrix} \cos \alpha x_3 & -\sin \alpha x_3 & 0 \\ \sin \alpha x_3 & \cos \alpha x_3 & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} x_1 \\ x_2 \\ x_3 \end{pmatrix}$$



$$\begin{pmatrix} x_1' \\ x_2' \\ x_3' \end{pmatrix} = \begin{pmatrix} x_1 \cdot \cos \alpha x_3 & -x_2 \cdot \sin \alpha x_3 \\ x_1 \cdot \sin \alpha x_3 & +x_2 \cdot \cos \alpha x_3 \\ x_3 \end{pmatrix}$$

$$u_1 = x_1' - x_1 = x_1 \cdot (\cos \alpha x_3 - 1) - x_2 \cdot \sin \alpha x_3$$

$$u_2 = x_2' - x_2 = x_2 \cdot (\cos \alpha x_3 - 1) + x_1 \cdot \sin \alpha x_3$$

$$u_3 = x_3' - x_3 = 0$$

Assume  $\alpha$  is small.

$$\cos \alpha x_3 - 1 = -\frac{(\alpha x_3)^2}{2}$$

So,

$u_1 \approx -x_2 \cdot \alpha x_3$	$\Rightarrow$	$u_1 \approx -x_2 x_3 \cdot \alpha$
$u_2 \approx x_1 \cdot \alpha x_3$		$u_2 \approx x_1 x_3 \cdot \alpha$
$u_3 = 0$		$u_3 \approx 0$

$$[\alpha] = L^{-1}$$

$\alpha$  small means  $\alpha \ll L^{-1}$

$L$  is length of beam.

$$e_{ij} = \frac{1}{2} \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right)$$

$$\frac{\partial u_i}{\partial x_j} = \begin{pmatrix} 0 & -\alpha x_3 & -\alpha x_2 \\ \alpha x_3 & 0 & \alpha x_1 \\ 0 & 0 & 0 \end{pmatrix}$$

$$[e_{ij}] = \begin{pmatrix} 0 & 0 & -\frac{\alpha x_2}{2} \\ 0 & 0 & \alpha x_1/2 \\ -\frac{\alpha x_2}{2} & \frac{\alpha x_1}{2} & 0 \end{pmatrix}$$

$$\Delta = \text{tr}(e_{ij}) = 0 \Rightarrow \text{no volume change.}$$

$$\sigma_{ij} = 2\mu \left( e_{ij} - \cancel{\Delta \frac{\delta_{ij}}{3}} \right) + K \cdot \cancel{\Delta \frac{\delta_{ij}}{3}}$$

$$\sigma_{ij} = 2\mu e_{ij}$$

$$[\sigma] = \begin{pmatrix} 0 & 0 & -\mu \alpha x_2 \\ 0 & 0 & \mu \alpha x_1 \\ -\mu \alpha x_2 & \mu \alpha x_1 & 0 \end{pmatrix}$$

$$\text{force density: } f_i = \frac{\text{force}}{\text{volume}} = \frac{\partial}{\partial x_j} \sigma_{ij}$$

$$f_1 = \frac{\partial}{\partial x_1} \sigma_{11} + \frac{\partial}{\partial x_2} \sigma_{12} + \frac{\partial}{\partial x_3} \sigma_{13}$$

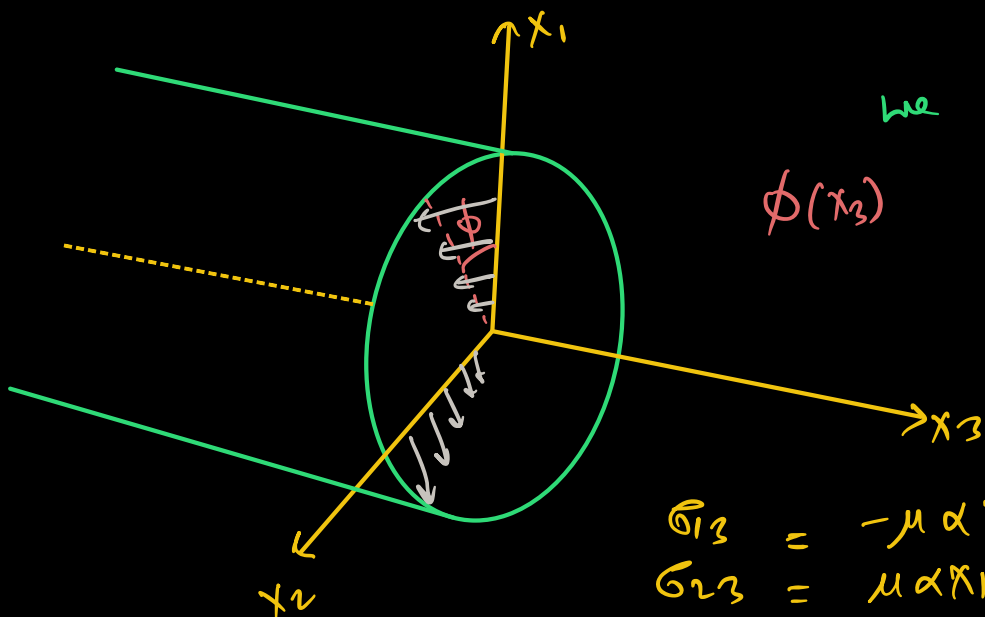
$$= 0 + 0 + 0$$

$$f_2 = 0$$

$$f_3 = 0$$

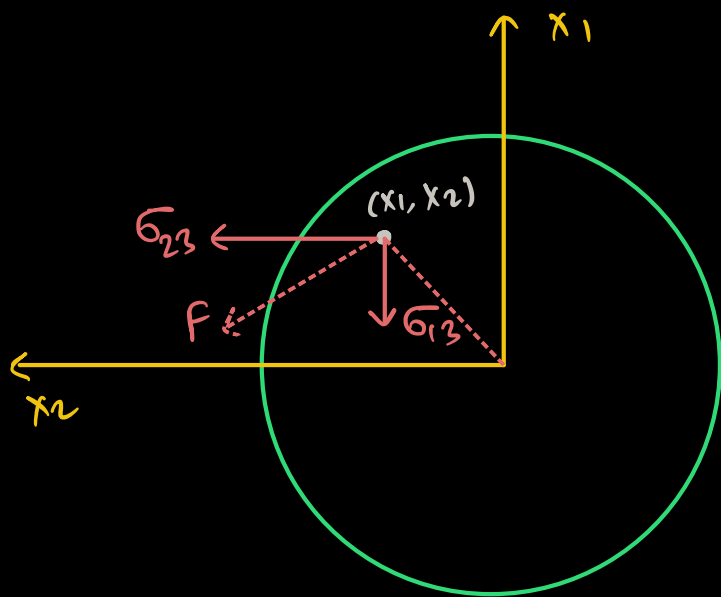
$\bar{f} = 0$  force per unit volume is zero everywhere  
(and it is consistent; because the body  
is at equilibrium)

→ we applied torque & had hold it.  
given point are at rest & internal  
stresses have developed.



we want to find  
 $\phi(x_3)$  forces on surface.

$$\begin{aligned}\sigma_{13} &= -\mu \alpha x_2 \\ \sigma_{23} &= \mu \alpha x_1 \\ \sigma_{33} &= 0\end{aligned}$$



Force / Area

$$\vec{F} = \mu \alpha (-x_2 \hat{i} + x_1 \hat{j})$$

$$\vec{r} = (x_1 \hat{i} + x_2 \hat{j})$$

$$\Rightarrow \vec{F} \perp \vec{r}$$

Torque per unit area  $\vec{r} \times \vec{F} dA$

$$= \sqrt{x_1^2 + x_2^2} \times \mu \cdot \alpha \sqrt{x_2^2 + x_1^2} \cdot dA \hat{z}$$

$$= \mu \alpha (x_1^2 + x_2^2) dA \hat{z}$$

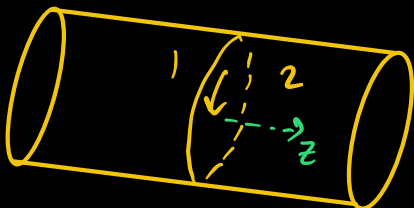
$$= \mu \alpha r^2 \cdot dA \cdot \hat{z}$$

$$\mu \alpha r^2 \hat{z} = \text{Torque per unit area.}$$

$$\text{Torque on the cross section} = \int_0^{2\pi} \int_0^R \mu \cdot \alpha \cdot r^2 \cdot r \cdot dr d\theta$$

$$= \mu \cdot \alpha \cdot 2\pi \int_0^R r^3 dr$$

$$= \frac{\mu \alpha \pi R^4}{2}$$



Torque applied by section 2 on section 1

